

ELECTRODYNAMICS

C1 MAXWELL EQUATIONS - THEY ARE COMPLETE ^{SYSTEM} - NO OTHER EQ.

- LINKS FIELDS \vec{E}, \vec{B} TO THE CHARGES ρ, \vec{j}

ρ - CHARGE DENSITY ; \vec{j} - FLUX

- IN AN INERTIAL FRAME, MAXWELL'S EQUATIONS ASSUMES THE SHAPE

① $\text{div } \vec{D} = 4\pi\rho$ GAUSS-LAW

$$\int_V dV \text{div } \vec{D} = \int_V dV \vec{\nabla} \cdot \vec{D} = \psi = 4\pi \int_V dV \rho = 4\pi Q$$

EL. FLUX ↑ TOTAL CHARGE

$|\psi|$

EL. CHARGES ARE SOURCES OF THE ELECTRIC FIELD, EL. FLUX ~~IS~~ IS EQUAL TO THE ENCLOSED CHARGE

② $\text{div } \vec{B} = 0$

$$\int_V dV \text{div } \vec{B} = \int_V dV \vec{\nabla} \cdot \vec{B} = \phi = 0$$

MAG. FLUX (IN CLOSE SURFACE)

THERE ARE NO MAGNETIC CHARGES, MAGNETIC FLUX THROUGH A CLOSED SURFACE IS ALWAYS ZERO.

③ $\text{rot } \vec{E} = -\frac{1}{c} \frac{\partial}{\partial t} \vec{B}$

FARADAY'S LAW OF INDUCTION

$$\int_A d\vec{A} \cdot \text{rot } \vec{E} = \int_{\partial A} d\vec{s} \cdot \vec{E} = -\frac{1}{c} \frac{\partial}{\partial t} \int_A d\vec{A} \cdot \vec{B} = -\frac{1}{c} \frac{\partial}{\partial t} \phi$$

INDUCED VOLTAGE ALONG LOOP ↑ MAG. FLUX DEFINED OVER SOME SURFACE

IN THE MAGNETIC FLUX THROUGH A LOOP CHANGES, THERE IS AN INDUCED ELECTRIC FIELD.

MINUS SIGN FROM LENZ RULE.

ELECTRODYNAMICS

④ $\text{rot } \vec{H} = \frac{4\pi}{c} \vec{j} + \frac{1}{c} \frac{\partial}{\partial t} \vec{D}$ AMPÈRE'S LAW

ELECTRODYNAMICS

$$\int_A d\vec{A} \text{ rot } \vec{H} = \int_{\partial A} d\vec{s} \vec{H} = \frac{4\pi}{c} \int_A d\vec{A} \cdot \vec{j} + \frac{1}{c} \frac{d}{dt} \int_A d\vec{A} \cdot \vec{D}$$

$$= \frac{4\pi}{c} \cdot I + \frac{1}{c} \frac{d}{dt} \psi$$

THERE IS A MAGNETIC FIELD AROUND A CURRENT OR AROUND CHANGING ELECTRIC FLUX.

C2 ELECTRODYNAMICS IN VACUUM

- THE FIELDS \vec{D} AND \vec{H} ARE RELATED TO \vec{E} AND \vec{B}

BY THE DIELECTRIC CONSTANT ϵ AND THE PERMEABILITY μ

$$\vec{D} = \epsilon \cdot \vec{E} \quad \vec{B} = \mu \cdot \vec{H}$$

\uparrow DIELECTRIC CONSTANT \uparrow PERMEABILITY $\epsilon = \mu = 1$ IN VACUUM

C3 CHARGE CONSERVATION

- EL. CHARGE IS CONSERVED, SO THE DYNAMICS OF FIELD MUST RESPECT THAT!

$$\text{rot } \vec{H} = \frac{4\pi}{c} \left(\vec{j} + \frac{1}{4\pi} \frac{d}{dt} \vec{D} \right) = \frac{4\pi}{c} \cdot \vec{z}$$

$\vec{z} \sim$ DISPLACEMENT CURRENT

$$\text{div } \vec{z} = \text{div } \vec{j} + \frac{1}{4\pi} \frac{d}{dt} \text{div } \vec{D} = \text{div } \vec{j} + \frac{\partial}{\partial t} \rho = 0$$

CONSERV. LAW IS ALREADY IN MAXWELL'S EQUATIONS

THE 4π 'S

BECAUSE $\text{DIV ROT } \vec{H} = 0$


C4 OHM'S LAW - PHENOMENOLOGICAL RELATION

$$\vec{j} = \sigma \cdot \vec{E}, \text{ WITH CONDUCTIVITY (TENSOR) } \sigma$$

[USUAL FOR]

$$\underbrace{j \cdot A}_I = \sigma \cdot A \cdot \frac{z}{z} E = \frac{\sigma \cdot A}{z} \cdot \frac{E \cdot z}{z}$$

\downarrow $\frac{1}{R}$ \uparrow $j_i = \sigma_{ij} \cdot E_j$



C5 LORENZ FORCE

$$\vec{f} = \rho \cdot \vec{E} + \frac{1}{c} \vec{j} \times \vec{B} \quad (\text{CAREFUL, RELATIVISTIC FORCE})$$

- IT'S ONLY VALID IN A SPECIFIC INERTIAL FRAME AND ONE NEEDS A RELATIVISTIC EQUATION OF MOTION, OTHERWISE $\beta \ll 1$, $|v| \ll c$

C6 ENERGY CONSERVATION IN ELECTRODYNAMICS → POYNTING LAW

(ASSUME VACUUM $\epsilon = \mu = 1 \rightarrow \vec{E} = \vec{D}$ AND $\vec{B} = \vec{H}$)

TAKE THE TWO ROT-EQUATIONS AND ASSEMBLE

$$\vec{E} \cdot \text{rot} \vec{B} - \vec{B} \cdot \text{rot} \vec{E} = \frac{4\pi}{c} \vec{j} \cdot \vec{E} + \frac{1}{c} \left(\vec{E} \cdot \frac{\partial}{\partial t} \vec{E} + \vec{B} \cdot \frac{\partial}{\partial t} \vec{B} \right) = -\text{div}(\vec{E} \times \vec{B})$$

↑
SIGN SWITCH, LENZ-RULE

ENERGY DENSITIES OF ELECTRIC PART OF THE FIELD:

$$w_{\text{el}} = \frac{\vec{E}^2}{8\pi} \rightarrow \frac{\partial}{\partial t} w_{\text{el}} = \frac{1}{4\pi} \vec{E} \cdot \frac{\partial}{\partial t} \vec{E}$$

MAGNETIC PART OF THE FIELD

$$w_{\text{mag}} = \frac{\vec{B}^2}{8\pi} \rightarrow \frac{\partial}{\partial t} w_{\text{mag}} = \frac{1}{4\pi} \vec{B} \cdot \frac{\partial}{\partial t} \vec{B}$$

$$\text{BECAUSE } w_{\text{el}} = -\frac{1}{2} \int_V \text{div} \phi = -\frac{1}{8\pi} \int_V \text{div} \Delta \phi \cdot \phi = \frac{1}{8\pi} \int_V \text{div} (\nabla \phi)^2$$

$$\rightarrow \frac{c}{4\pi} \text{div} (\vec{E} \times \vec{B}) + \vec{j} \cdot \vec{E} = -\frac{\partial}{\partial t} (w_{\text{el}} + w_{\text{mag}}) =$$

$$= \text{div} \vec{S} + \vec{j} \cdot \vec{E}$$

ENERGY CONSERVATION IN THE FORM OF CONTINUITY RELATION.

\vec{S} - POYNTING VECTOR (DESCRIBES ENERGY TRANSPORT OF ENERGY SPECTRUM)

ENERGY CONSERVATION: CHANGE IN ENERGY IS EQUAL TO THE ENERGY FLUX ACROSS THE BOUNDARY AND TO THE DISSIPATION.

C2. MOMENTUM CONSERVATION

LORENTZ FORCE. $\vec{F} = \int dV \vec{f} = \int dV (\rho \vec{E} + \frac{1}{c} \vec{j} \times \vec{B}) = \frac{d}{dt} \vec{P}$
FORCES ACTING ON A CHARGE AND ACCELERATE IT.

LORENTZ FORCE DENSITY

EQUATION OF MOTION

WILL E AND B EXPRESS P AND P?

$\text{div } \vec{E} = 4\pi\rho \rightarrow \rho = \frac{1}{4\pi} \text{div } \vec{E}$

$\text{rot } \vec{B} = \frac{4\pi}{c} \vec{j} + \frac{1}{c} \frac{\partial \vec{E}}{\partial t} \rightarrow \vec{j} = \frac{c}{4\pi} (\text{rot } \vec{B} - \frac{1}{c} \frac{\partial \vec{E}}{\partial t})$

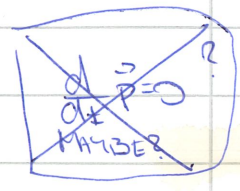
$\rightarrow \frac{d}{dt} \vec{P} = \frac{1}{4\pi} \int dV (\vec{E} \cdot \text{div } \vec{E} + \text{rot } \vec{B} \times \vec{B} - \frac{1}{c} \frac{\partial \vec{E}}{\partial t} \times \vec{B}) =$

$\frac{\partial}{\partial t} (\vec{E} \times \vec{B}) - \vec{E} \times \frac{\partial \vec{B}}{\partial t}$
 $\uparrow - \frac{1}{c} \frac{\partial \vec{B}}{\partial t} = \text{rot } \vec{E}$

$\rightarrow \frac{d}{dt} \vec{P} = \frac{1}{4\pi} \int dV (\vec{E} \cdot \text{div } \vec{E} + \vec{B} \cdot \text{div } \vec{B} + \text{rot } \vec{B} \times \vec{B} - \frac{1}{c} \frac{\partial \vec{E}}{\partial t} \cdot$

$\cdot (\vec{E} \times \vec{B}) - \vec{E} \times \text{rot } \vec{E})$

CAN BE ADDED BECAUSE ITS ZERO



$\rightarrow \frac{d}{dt} (\vec{P} + \frac{1}{c^2} \int dV \vec{S}) = \frac{1}{4\pi} \int dV (\vec{E} \cdot \text{div } \vec{E} + \vec{B} \cdot \text{div } \vec{B} + \text{rot } \vec{B} \times \vec{B} + \text{rot } \vec{E} \times \vec{E})$

POINTING VECTOR

RATE OF CHANGE OF MOMENTUM CARRIED BY THE FIELD
 → SOURCE IF (BRACKETS) CAN BE WRITTEN AS A DIVERGENCE.

$(\vec{E} \cdot \text{div } \vec{E})_i = E_i \partial_j E_j$ (VECTORIAL PROPERTY CARRIED BY E_i)

$(\vec{E} \times \text{rot } \vec{E})_i = \epsilon_{ijk} E_j \cdot \epsilon_{kmn} \partial_l E_m = (-\partial_{im} \partial_{jl} + \partial_{il} \partial_{jm}) \cdot$

$\cdot E_j \partial_l E_m = E_j \partial_j E_i - E_j \partial_i E_j$

(3)

$$\rightarrow (\vec{E} \cdot \text{div} \vec{E} + \text{rot} \vec{E} \times \vec{E})_i = E_i \nabla_j E_j + E_j \nabla_j E_i - E_j \nabla_i E_j =$$

$$= \nabla_j (E_i E_j) - \frac{1}{2} \nabla_i (E_j E_i) = \nabla_j (E_i E_j - \frac{1}{2} \delta_{ij} (E_k E_k))$$

SAME ARGUMENT FOR $\vec{B} \cdot \text{div} \vec{B} + \text{rot} \vec{B} \times \vec{B}$

$$\rightarrow \frac{d}{dt} \left(\vec{P} + \frac{1}{c^2} \int dV \vec{S} \right) = \frac{1}{4\pi} \int dV \nabla_j (E_i E_j + B_i B_j - \frac{1}{2} \delta_{ij} \cdot$$

$$\cdot (E_k E_k + B_k B_k))$$

ENERGY DENSITIES: $w_{\text{el}} = \frac{1}{8\pi} \vec{E}^2$ AND $w_{\text{mag}} = \frac{1}{8\pi} \vec{B}^2$

$$\rightarrow W = \vec{E}^2 + \vec{B}^2 = 8\pi \quad (\text{TOTAL ENERGY DENSITY})$$

MAXWELL STRESS - TENSOR

$$T_{ij} = \frac{1}{4\pi} (E_i E_j + B_i B_j - \frac{1}{2} \delta_{ij} W)$$

TRACE $\text{tr} T = \sum_i T_{ii} = \frac{1}{4\pi} (\vec{E}^2 + \vec{B}^2 - \frac{1}{2} (\vec{E}^2 + \vec{B}^2)) = w_{\text{el}} + w_{\text{mag}} =$

$$= \frac{W}{4\pi}$$

MOMENTUM CONSERVATION: COLLECT ALL RESULTS

$$\rightarrow \frac{d}{dt} \left(\vec{P} + \frac{1}{c^2} \int dV \vec{S} \right)_i = \int dV \nabla_j T_{ij} = \int dA_j T_{ij}$$

(NOT CONSERVED!)

$T_{ij} \sim$ STRESS SURFACE - NORMALISED FORCE INTO i -DIRECTION
 PULLING ON THE SURFACE WITH NORMAL INTO j -DIRECTION
 CORRESPONDS TO STRESS TENSOR $\underline{\sigma}$ (REYNOLDS IN FLUID MECHANICS).

C8 ELECTROMAGNETIC WAVES IN VACUUM

- LETS INTRODUCE VECTOR POTENTIAL \vec{A} AS $\vec{B} = \text{rot} \vec{A}$

$$\text{rot} \vec{E} = -\partial_{ct} \vec{B} \rightarrow \text{rot} (\vec{E} + \partial_{ct} \vec{A}) = 0$$

\downarrow
 $\rightarrow \partial_{ct} = \frac{1}{c} \frac{\partial}{\partial t}$

- IMPLYING THAT THERE MUST BE A SCALAR

POTENTIAL $\vec{E} + \partial_{ct} \vec{A} = -\nabla \phi$ BECAUSE $\text{rot} \nabla \phi = 0$
 \uparrow
CONVENTION

$$\vec{E} = -\nabla \phi - \partial_{ct} \vec{A} \rightarrow \text{div} \vec{E} = -\Delta \phi - \partial_{ct} \text{div} \vec{A} = 4\pi \rho$$

$$\vec{B} = \text{rot} \vec{A} \rightarrow \text{rot} \text{rot} \vec{A} = \frac{4\pi}{c} \vec{j} + \partial_{ct} (-\nabla \phi - \partial_{ct} \vec{A}) =$$
$$= \nabla \text{div} \vec{A} - \Delta \vec{A}$$

$$\text{rot} \text{rot} \vec{A} = \nabla \text{div} \vec{A} - \Delta \vec{A} : \text{TWO COUPLED EQUATIONS}$$

$$\Delta \phi + \partial_{ct} \text{div} \vec{A} = -4\pi \rho$$

$$\Delta \vec{A} - \partial_{ct}^2 \vec{A} = \frac{4\pi}{c} \vec{j} + \nabla (\text{div} \vec{A} + \partial_{ct} \phi)$$

C9 GAUGE TRANSFORMATION

\vec{B} (THE PHYSICAL FIELD [MAGNETIC]) IS NOT CHANGED

IF $\nabla \chi$ IS ADDED TO THE POTENTIAL \vec{A}

$$\text{rot} \nabla \chi = 0$$

\vec{E} CAN'T CHANGE EITHER! ONE MUST CHANGE ϕ

BY $-\partial_{ct} \chi$ (OR IN OTHER WORDS, SUBTRACT IT FROM ϕ)

GAUGE TRANSFORMATION

$$\vec{A} \rightarrow \vec{A}' = \vec{A} + \nabla \chi$$

$$\phi \rightarrow \phi' = \phi - \partial_{ct} \chi$$

THE GAUGE TRANSFORMATION LEAVE THE FIELDS INVARIANT, BUT CAN BE USED TO DECOUPLE THE RELATIONS FOR THE POTENTIALS (FIELD EQUATIONS).

A SPECIFIC GAUGE TRANSFORMATION CAN BE CHOSEN TO FULFILL A GAUGE FIXING CONDITION.

LORENZ - GAUGE $\partial_{ct} \phi + \text{div} \vec{A} = 0$

$$\Delta \phi - \partial_{ct}^2 \phi = -4\pi \rho = \square \phi$$

$$\Delta \vec{A} - \partial_{ct}^2 \vec{A} = \frac{4\pi}{c} \vec{j} = \square \vec{A}$$

~ WAVE EQUATIONS WITH SOURCES ρ AND \vec{j} .

D'ALEMBERT OPERATOR:

$$\square = \partial_{ct}^2 - \nabla^2 = \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \frac{\partial^2}{\partial x^2} - \frac{\partial^2}{\partial y^2} - \frac{\partial^2}{\partial z^2}$$

LAPLACE OPERATOR IN MINLOWSKI SPACE = D'ALEMBERT OPERATOR

WHAT HAPPENS TO THE GAUGE FIELD χ ?

$$\partial_{ct} (\phi - \partial_{ct} \chi) + \text{div} (\vec{A} + \nabla \chi) = 0 \text{ IN LORENZ GAUGE}$$

$$\text{IT IS CONSISTENT IF } \Delta \chi - \partial_{ct}^2 \chi = 0$$

→ POTENTIALS ϕ AND \vec{A} ARE LINKED TO THE SOURCES \vec{j} AND ρ WITH A WAVE EQUATION. CHANGES IN \vec{j} OR ρ PROPAGATION AT THE SPEED c . (LIGHT SPEED 299792458 m/s).

CLO COUPLING THE POTENTIALS TO THE CHARGES

(\square REPLACES Δ , AND WE NEED A NEW TIME-DEPENDENT GREEN FUNCTION)

$$\text{TYPE OF EQUATION } \Delta \psi - \partial_{ct}^2 \psi = -4\pi f$$

CONSTRUCT A GREEN-FUNCTION FOR DEALING WITH THE CHARGE/CURRENT DISTRIBUTION ~~AND WITH~~

GREEN-FUNCTION G , DEFINED THROUGH

$$(\Delta - \partial_{ct}^2) G(\vec{r}; t; \vec{r}'; t') = -4\pi \delta_D(\vec{r} - \vec{r}') \delta_D(t - t')$$

THE FIELD OF ANY CHARGE DISTRIBUTION IS CONSTRUCTED BY SUPERPOSITION, BECAUSE MAXWELL-ELECTRODYNAMICS IS LINEAR.

$$\psi(\vec{r}; t) = \int d^3r' \int dt' G(\vec{r}; t; \vec{r}'; t') \cdot f(\vec{r}'; t')$$

~ CONVOLUTION RELATION

C11 LIENARD-WIECHERT POTENTIALS

→ CONSTRUCT THE GREEN FUNCTION FOR LINEAR POTENTIAL PROBLEMS.

FOURIER SPACE - TRANSFORMATION THE DIFFERENTIAL EQVAT. INTO ALGEBRAIC (INVERSION IS EASY)

$$G(\vec{r}; t; \vec{r}'; t') = \int d^3k \int d\omega G(\vec{k}; \omega) \cdot \exp[i\vec{k}(\vec{r} - \vec{r}')] \cdot \exp[-i\omega(t - t')]$$

MINUS CONVENTION

$$\square G(\vec{r}; t; \vec{r}'; t') = \int d^3k \int d\omega G(\vec{k}; \omega) \square \exp[i\vec{k}(\vec{r} - \vec{r}')] \cdot$$

WITH $\square = \Delta - \partial_{ct}^2$

D'A LEMBERT OPERATOR

$$\cdot \exp[-i\omega(t - t')] = \int d^3k \int d\omega G(\vec{k}; \omega) \cdot$$

$$\cdot \left[k^2 - \left(\frac{\omega}{c}\right)^2 \right] \exp[i\vec{k}(\vec{r} - \vec{r}')] \cdot \exp[-i\omega(t - t')]$$

MEZI KROK

$$= \int d^3k \int d\omega G(\vec{k}; \omega) (\Delta \exp[i\vec{k}(\vec{r} - \vec{r}')] - \partial_{ct}^2 \exp[-i\omega(t - t')])$$

5

LET'S COMPARE IT WITH THE RIGHT SIDE OF THE WAVE EQ.:

$$= \int \frac{d^3k}{(2\pi)^3} \cdot \int \frac{d\omega}{2\pi} \cdot (4\pi) \cdot \exp(i\vec{k}(\vec{r}-\vec{r}')) \cdot \exp(-i\omega(t-t'))$$

$$\rightarrow G(\vec{k}; \omega) = \frac{1}{4\pi^3} \cdot \frac{1}{k^2 - \frac{\omega^2}{c^2}} \quad \text{GREEN FUNCTION IN FOURIER SPACE}$$

TRANSFORM BACK TO REAL SPACE:

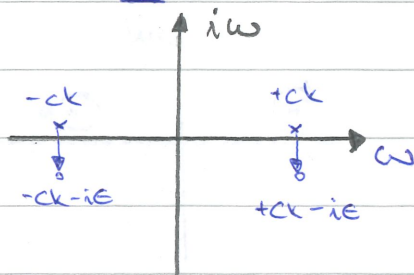
$$G(\vec{r}-\vec{r}'; t-t') = \frac{1}{4\pi^3} \cdot \int d^3k \int d\omega \underbrace{\frac{c^2}{(ck)^2 - \omega^2}}_{= G(\vec{k}; \omega)} \exp[i\vec{k}(\vec{r}-\vec{r}')] \cdot \exp[i\omega(t-t')]$$

BUT THE INTEGRAND IS NOT DEFINED AT $\omega = \pm ck$

→ USE TOOLS FROM COMPLEX-CALCULUS FOR PERFORMING THE INTEGRATION AT $\omega = \pm ck$

$d\omega$ - INTEGRATION MOVE TO THE COMPLEX PLANE

LET'S OFFSET THE POLES OF THE INTEGRAND BY A SMALL AMOUNT ϵ .



FOR $d\omega \frac{\exp(-i\omega t)}{(ck)^2 - (\omega - i\epsilon)^2}$

AS PART OF THE GREEN'S FUNCTION

$$G(\vec{r}; t) = \frac{1}{4\pi^3} \int d^3k \int d\omega \frac{\exp[i(\vec{k}\vec{r} - \omega t)]}{k^2 - (\frac{\omega - i\epsilon}{c})^2}$$

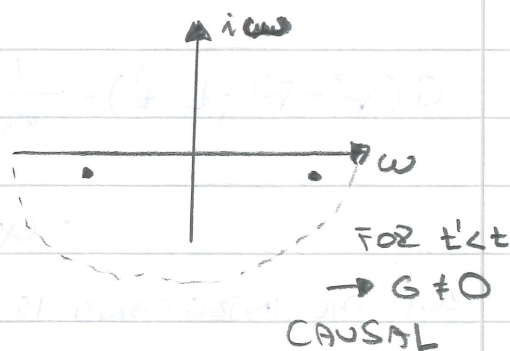
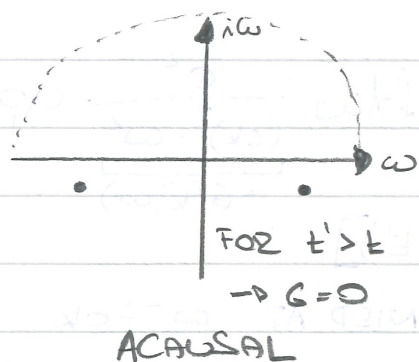
FOR APPLYING THE RESIDUE THEOREM ONE NEEDS A CLOSED

LOOP → INTEGRAND SHOULD DROP AT LEAST LIKE $1/\omega^3$,

OTHERWISE THE CLOSING ...

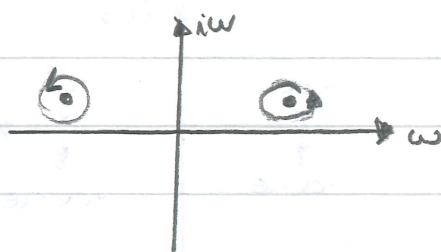
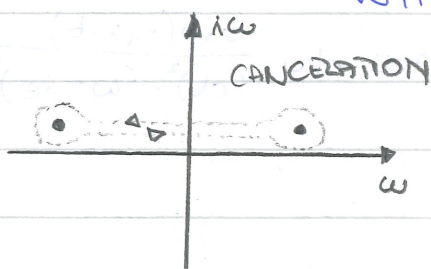
IF THE LOOPS DOES NOT CONTAIN ANY OF THE POLES, THE RESULT = 0 DUE TO HOLOMORPHY (COMPLEX DIFFERENTIABILITY).

WE CAN CHOOSE THE CLOSING ARE TO CONTAIN THE POLES FOR $t > t'$; $\mathcal{C} > 0$ (CAUSAL SOLUTION) OR NOT TO CONTAIN THE POLES FOR $t < t'$; $\mathcal{C} < 0$ (ACAUSAL SOLUTION)



FOR A HOLOMORPHIC FUNCTION THE INTEGRATION PATH DOES NOT MATTER:

~~MORPHOLOGICAL~~ INTEGRATION PATH TO \rightarrow WE ARE LEFT WITH TWO INTEGRALS AROUND TWO POLES



EFFECTIVELY, THERE ARE TWO INDIVIDUAL LOOPS LEFT WITH INDIVIDUAL SENSES OF ROTATION \rightarrow IDENTICAL RESIDUES, DUE TO SYMMETRY.

C13 COVARIANT FORMULATION OF ELECTRODYNAMICS

LORENTZ - FORCE + FIELD TENSOR (FARADAY TENSOR)

$$\frac{d\vec{p}}{dt} = q \cdot (\vec{E} + \frac{\vec{v}}{c} \times \vec{B})$$

$\frac{d\vec{p}}{dt} \rightarrow$ MOTION OF A CHARGE

6.

ONE OBSERVER SEES ELECTRIC & MAGNETIC FIELD
 A OBSERVES A CHARGE, IN THAT FRAME THE
 MOTION OF A CHARGE IS GIVEN BY $\frac{d\vec{p}}{dt}$,
 ANOTHER OBSERVER MIGHT SEES DIFFERENT \vec{E} & \vec{B} ?
 BUT FOR THAT OBSERVER THE TRAJECTORY IS
 OF COURSE DIFFERENT. THE \vec{E} & \vec{B} CHANGE
 CONSISTENTLY AS WELL AS THE TRAJECTORY OF
 THAT PARTICULAR PARTICLE (LORENTZ TRANSFORM).

LET'S TRANSFORM ~~IT~~ INTO A COVARIANT EQUATION,
 WE CAN DO TWO THINGS - FIRST MULTIPLY WITH

$\frac{1}{\gamma} \frac{d\vec{r}}{dt}$; AND MULTIPLYING WITH $\frac{1}{\gamma}$

COMBINE

$$\left\{ \begin{array}{l} m \frac{1}{\gamma} \frac{d}{dt} (\gamma c) = \frac{q}{c} \vec{E} \cdot \vec{r} \quad \frac{1}{\gamma} \frac{d\vec{r}}{dt} \\ m \frac{1}{\gamma} \frac{d}{dt} (\gamma \vec{r}) = q \cdot (\vec{E} + \frac{1}{c} \vec{r} \times \vec{B}) \quad \frac{1}{\gamma} \end{array} \right.$$

$$m \frac{d}{dt} u^\mu = q \cdot \begin{matrix} \overbrace{\phantom{F_{\mu\nu}}}^{F_{\mu\nu}} \\ \begin{pmatrix} 0 & -E_x & -E_y & -E_z \\ E_x & 0 & -B_x & B_y \\ E_y & B_x & 0 & -B_z \\ E_z & -B_y & B_z & 0 \end{pmatrix} \cdot u_\nu \end{matrix}$$

WITH FOUR VELOCITY $u^\mu = \frac{1}{\gamma} (c, \vec{r}) = (\gamma c, \gamma \vec{r})$ AND
 PROPER TIME $\frac{dt}{d\tau} = \gamma$; WHERE $F_{\mu\nu}$ IS ANTI SYMMETRIC
 TENSOR CONSERVE NORMALISATION OF u^μ (FOUR VECTOR OF A VELOCITY)
 IT TRACE IS ZERO.

FOUR VELOCITY $u_\mu u^\mu = c^2$, VELOCITIES IN MINKOWSKI
 SPACE DON'T FORM VECTOR SPACE, WE CANNOT ^{SIMPLY} ADD
 FOUR VELOCITIES. IF WE WANT TO ADD VELOCITIES WE
 HAVE TO CONSERVE THIS NORMALISATION $u_\mu u^\mu = c^2$,
 IF FORCE LIKE "LORENTZ FORCE" ACTS ON PARTICLE.

$$\frac{1}{2} \frac{d}{dt} (m_\mu u^\mu) = 0 = m_\mu \frac{du^\mu}{dt} = m_\mu \cdot \frac{q}{m_0} F^{\mu\nu} u_\nu =$$

$$= \frac{q}{m_0} F^{\mu\nu} m_\mu u_\nu = 0$$

$\underbrace{\hspace{10em}}_{\text{SYMMETRIC}}$
 $\underbrace{\hspace{10em}}_{\text{ANTISYMMETRIC}}$ - THIS TENSOR COMES IN IN FIELD EQUATION

• FIELD EQUATION AND CHARGE CONSERVATION.

$$\partial_\mu F^{\mu\nu} = \frac{4\pi}{c} \partial_\nu j^\nu = 0$$

MAXWELL FIELD EQUATION

~~CONSERVATION OF CHARGE~~

STRONG CONSTRAINT HOW PARTICLES SHOULD MOVE, SIMILAR CONSTRAINT HOW THE FIELDS GENERATED BY THE PARTICLES.

$$\partial_\nu \partial_\mu F^{\mu\nu} = \frac{4\pi}{c} \partial_\nu j^\nu = 0$$

CONSERVATION OF CHARGE

$\partial_\nu \partial_\mu$ - SYMMETRIC TENSOR

$F^{\mu\nu}$ - ANTISYMMETRIC TENSOR - HIDDEN CONSERV. OF CHARGE.

CONSTRUCTION OF THE FIELD EQUATION:

- LINEAR

- FIRST ORDER IN F , SECOND ORDER IN A

VECTOR FIELD THAT IS LORENTZ VECTOR

POSTULATE A 4-POTENTIAL A_μ AS A LORENTZ-VECTOR TO CONTAIN ϕ AND \vec{A}

$$A_\mu \rightarrow F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu \quad \text{IS AUTOMATICALLY ANTISYMMETRIC}$$

~~$$\partial^\mu F_{\mu\nu} = \partial^\mu (\partial_\mu A_\nu - \partial_\nu A_\mu) =$$~~

$$= \underbrace{\partial^\mu \partial_\mu A_\nu}_{= \square \text{ D'ALAMB. OP.}} - \underbrace{\partial_\nu \partial^\mu A_\mu}_{= 0 \text{ IN LORENTZ GUAGE TRANSF}} = \frac{4\pi}{c} j^\nu$$

GAUGE TRANSFORMATION

$$A_\mu \rightarrow A'_\mu = A_\mu + \partial_\mu \chi$$

↑ GAUGE FUNCTION

$$F_{\mu\nu} \rightarrow F'_{\mu\nu} = F_{\mu\nu} = \partial_\mu (A_\nu + \partial_\nu \chi) - \partial_\nu (A_\mu + \partial_\mu \chi) =$$

$$= \partial_\mu A_\nu - \partial_\nu A_\mu + \underbrace{\partial_\mu \partial_\nu \chi - \partial_\nu \partial_\mu \chi}_0 = F_{\mu\nu}$$

ANTISYMMETRY OF F AND THE DUAL TENSOR \tilde{F}
 $F_{\mu\nu}$ HAS 6 ENTRIES IN 4 DIMENSIONS DUE TO ANTISYMMETRY $\rightarrow 3 \times \vec{E}$ & $3 \times \vec{B}$.

$$F_{\mu\nu} = -F_{\nu\mu} \rightarrow \underbrace{\partial_\lambda F_{\mu\nu} + \partial_\mu F_{\nu\lambda} + \partial_\nu F_{\lambda\mu}}_{\text{CYCLIC PERMUTATION}} = 0 = * \text{BIANCHI-IDENTITY}$$

$$* = \partial_\lambda (\partial_\mu A_\nu - \partial_\nu A_\mu) + \partial_\mu (\partial_\nu A_\lambda - \partial_\lambda A_\nu) + \partial_\nu (\partial_\lambda A_\mu - \partial_\mu A_\lambda) = 0$$

(TERMS CANCEl PAIRWISE)

THE BIANCHII IDENTITY IS THE FIELD EQUATION FOR VACUUM SOLUTIONS AND PROVIDES MEANS OF PROPAGATION OF THE FIELD:

$$\partial_\lambda F_{\mu\nu} + \partial_\mu F_{\nu\lambda} + \partial_\nu F_{\lambda\mu} = 0 \quad / \cdot \partial^\lambda$$

$$\underbrace{\partial^\lambda \partial_\lambda F_{\mu\nu}}_{\square} + \underbrace{\partial^\lambda \partial_\mu F_{\nu\lambda}}_0 + \underbrace{\partial^\lambda \partial_\nu F_{\lambda\mu}}_0 = 0 \rightarrow \square F_{\mu\nu} = 0$$

WITH $\square = \partial_\mu \partial^\mu$ WAVE EQUATION
 IN VACUUM

IF ONE HAS A STATIC SITUATION AND CHOOSES THE CONECT FRAME IN WHICH THE CHARGE IS AT REST.

$$\Delta F_{\mu\nu} = 0 \rightarrow \Delta \vec{E} = 0 = \Delta(\nabla\phi) = \nabla(\Delta\phi) \rightarrow \Delta\phi = 0$$

WITH THE CORRECT BOUNDARY CONDITION AT INFINITY.
FOR STATIC CASE $\Delta F_{\mu\nu} = 0$.

FIELD CAN ACTUALLY BE IN A REGION OF SPACE
WHERE THERE ARE NO CHARGES AROUND \rightarrow BIANCHI IDENT

FIELD DUAL: WRITE THE BIANCHI IDENTITY LIKE A
FIELD EQUATION

$$\partial_\lambda F_{\mu\nu} + \partial_\mu F_{\nu\lambda} + \partial_\nu F_{\lambda\mu} = 0$$

BECAUSE OF THE CYCLIC PROPERTY

$$\epsilon^{i\lambda\mu\nu} \partial_\lambda F_{\mu\nu} = \partial_\lambda \cdot \underbrace{(\epsilon^{i\lambda\mu\nu} F_{\mu\nu})}_{2 \cdot \tilde{F}^{i\lambda}} = 2 \cdot \partial_\lambda \tilde{F}^{i\lambda} = 0$$

\hookrightarrow TAKES THE SHAPE
OF FIELD EQUATION

ELECTRODYNAMICS: ① $\partial_\mu F^{\mu\nu} = \frac{4\pi}{c} j^\nu$ (FIELD EQUATION)

NAZARE DECIDES \longrightarrow
TO HAVE ONLY ELECTRON
CHARGE \rightarrow NO MONOPOLES.

② $\partial_\mu \tilde{F}^{\mu\nu} = 0$ (BIANCHI IDENTITY)

③ $\partial_\mu j^\mu = 0$

④ A^μ AND j^μ ARE LORENTZ VECTORS

IN VACUUM: $\partial_\mu F^{\mu\nu} = \partial_\mu \tilde{F}^{\mu\nu} = 0$

$\vec{E} \rightarrow \vec{B}$; $\vec{B} \rightarrow -\vec{E}$ ELECTROMAGNETIC DUALITY

- IN VACUUM BOTH WOULD BE IDENTICAL, ELECTROMAGNETICS

IN VACUUM IS IDENTICAL IF $\vec{E} \rightarrow \vec{B}$; $\vec{B} \rightarrow -\vec{E}$

"ELECTROMAGNETIC DUALITY"

(OF FIELDS)

INVARIANTS UNDER LORENTZ TRANSFORMATION

$$F_{\mu\nu} F^{\mu\nu} = 2 \cdot (\vec{E}^2 - \vec{B}^2) = \tilde{F}_{\mu\nu} \tilde{F}^{\mu\nu} \quad (\text{SCALAR})$$

$$F_{\mu\nu} \tilde{F}^{\mu\nu} = 4 \cdot \vec{E} \cdot \vec{B} \quad (\text{PSEUDO SCALAR})$$

ELECTRODYNAMICS FROM A VARIATIONAL PRINCIPLE

- LAGRANGE DENSITY - LORENTZ SCALARS \rightarrow MAKES SURE

THAT FIELD EQUATION IS COVARIANT.

~~is~~

LAGRANGE FUNCTION NOT A STATEMENT

ABOUT ENERGY BUT RATHER ABOUT CAUSALITY

- SQUARES OF FIRST DERIVATIVES ^(OF A_μ) \rightarrow SECOND ORDER FIELD EQ.

- COUPLING INVOLVING THE POTENTIAL \rightarrow LINEAR FIELD EQ.

LAGRANGE DENSITY \mathcal{L}

$$\mathcal{L} = \frac{1}{16} F_{\mu\nu} F^{\mu\nu} - \frac{1}{c} j_{\mu} A^{\mu} \rightarrow S = \int d^4x \mathcal{L}$$

\hookrightarrow FIRST DERIVATIVE OF A

HAMILTON'S PRINCIPLE $\delta S = 0 \rightarrow \partial_{\mu} F^{\mu\nu} = \frac{4\pi}{c} j^{\nu}$

$$\delta S = \int d^4x \delta F_{\mu\nu} F^{\mu\nu} = - \int d^4x A_{\mu} \square A^{\mu}; \quad \delta S = 0 \rightarrow \square A_{\mu} = \frac{4\pi}{c} j_{\mu}$$

HAMILTONIAN DENSITY $\mathcal{H} = \frac{1}{2} F_{\mu\nu} F^{\mu\nu}$

$$\mathcal{H} = \frac{\partial \mathcal{L}}{\partial(\partial_{\mu} A_{\nu})} \partial_{\mu} A_{\nu} - \mathcal{L} = \dots = \frac{1}{2} \cdot (\vec{E}^2 + \vec{B}^2) = F_{\mu\nu} \partial_{\mu} A_{\nu} + \frac{1}{4} F_{\mu\nu} F^{\mu\nu}$$

$F_{\mu\nu}$

ENERGY DENSITY OF THE FIELD, NOT A LORENTZ-INVARIANT

* WHEN WE DO LORENTZ TRANSFORMATION THIS IS NOT CONSERVED THEREFORE IT IS LORENTZ-INVARIANT.

PROPAGATION OF PHOTONS

- PHOTONS ARE A WAY OF PROPAGATION OF A FIELD ^{ITSELF} AWAY FROM A CHARGE,

WHAT WE KNOW ABOUT PHOTONS SHOULD FOLLOW FROM BIANCHI

IDENTITY.

$$(\text{curl} - \vec{\nabla} \times)$$

$$\square F_{\mu\nu} = 0, \text{ SOLVE WITH } F_{\mu\nu} = Q_{\mu\nu} \cdot \exp(i k_{\alpha} x^{\alpha})$$

$$\square F_{\mu\nu} = Q_{\mu\nu} \square \exp(i k_{\alpha} x^{\alpha}) = Q_{\mu\nu} \cdot \exp(i k_{\alpha} x^{\alpha}) \cdot \underbrace{k_{\alpha} k^{\alpha}}_0 = 0$$

WAVE VECTOR IS A MULTI VECTOR

$$\vec{k} \cdot \vec{k} = \left(\frac{\omega}{c}\right)^2 - \vec{k}^2 = 0 \rightarrow \omega = \pm ck$$

NO DISPERSION IN VACUUM

$$\text{GROUP VELOCITY } v_g = \frac{\partial \omega}{\partial k} = c = \frac{\omega}{k} = v_p = \text{PHASE VELOCITY}$$

PHOTON TRAJECTORY $x^\mu(\lambda)$

$$\vec{k}^\mu = \frac{dx^\mu}{d\lambda} \quad \text{- TANGENT TO THE TRAJECTORY}$$

$$x^\mu x_\mu = \frac{dx^\mu}{d\lambda} \cdot \frac{dx_\mu}{d\lambda} (d\lambda)^2 = \underbrace{k^\mu k_\mu}_{=0} (d\lambda)^2 = 0$$

$$\rightarrow x = \pm ct \quad \text{FROM } x_c - x^\mu = 0$$

C14 THE PLANCK SPECTRUM

- PHOTON ENERGY DISTRIBUTION FROM AN IDEALLY ABSORBING/EMITTING BODY IN THERMAL EQUILIBRIUM

HISTORICALLY: SECOND QUANTUM MECHANICAL SYSTEM

AFTER THE QUANTUM MECHANICAL

KEPLER-PROBLEM - HYDROGEN

- STATISTICAL MECHANICS

$$P(\epsilon; T) \sim \exp\left(-\frac{\Delta \epsilon}{kT}\right) \quad \text{BOLTZMANN-FACTOR}$$

PROBABILITY OF A THERMAL FLUCTUATION $\Delta \epsilon$ AT TEMPERATURE T

$$kT = \frac{1}{10} \text{ eV} \quad \text{AT } T = 300\text{K} \rightarrow \text{THERMAL FLUCTUATIONS ARE}$$

RELEVANT FOR ATOMS

$$k = 10^{-23} \text{ J/K}$$

WHY THIS SHAPE OF THE BOLTZMANN FACTOR?

→ TRANSITIVITY

RATIO $\frac{P(E_2; T)}{P(E_1; T)}$ CAN ONLY BE A FUNCTION OF $E_2 - E_1$

$$g(E_3 - E_1) = \frac{P(E_3; T)}{P(E_1; T)} = \frac{P(E_3; T)}{P(E_2; T)} \cdot \frac{P(E_2; T)}{P(E_1; T)} = g(E_3 - E_2) g(E_2 - E_1)$$

UNIQUELY SOLVED BY $g(E) = \exp(-\beta_i E)$

(LOOK AT THE LOGARITHM OF g)

① MINUS SIGN IS BY OBSERVATION: KINETIC SYSTEMS ARE ENERGETICALLY BOUNDED BELOW →

THIS IS DIFFERENT IN QUANTUM MECHANICAL SYSTEMS

② $\beta \sim \frac{1}{T}$, P SHOULD BE SMALL IF T IS SMALL

③ k FIXES UNITS ~ CONCEPT OF THERMAL ENERGY

(THINK OF BAROMETRIC FORMULA)

• THE PLANCK-SYSTEM IS ONE OF THE SIMPLEST QUANTUM SYST.

QUANTUM MECHANICS → SUPERPOSITION + INTERFERENCES

THERMAL PHYSICS → THERMAL ENERGY

THERE IS EQUIPARTITION IN SYSTEMS IN THERMAL EQUILIBRIUM

$$\frac{\vec{p}^2}{2m} = \frac{3}{2} kT \quad (\text{FOR A CLASSICAL SYSTEM})$$

$$d\vec{p} = \frac{3}{2} kT \quad (\text{FOR A RELATIVISTIC SYSTEM})$$

$$\hbar + p = \frac{h}{\lambda} = \hbar k \quad \text{de BROGLIE WAVE LENGTH}$$

$$\rightarrow \text{THERMAL WAVELENGTH } \lambda_{TH} = \frac{h}{p} = \frac{2}{3} \frac{hc}{k_B T}$$

PARTICLE SEPARATION $\gg \lambda_{TH}$ INTERFERENCE NOT IMPORTANT

SYSTEM BEHAVES CLASSICALLY (TRUE AT HIGH ENERGIES)

PARTICLE SEPARATION $\ll \lambda_{th}$ INTERFERENCE IS IMPORTANT

SYSTEM BEHAVES QUANTUM MECHANICALLY (TRUE AT LOW ENERGIES)

• PHOTON NUMBER STATISTICS

STATES $i \sim$ SORTED IN ENERGY

WEIGHTS $g_i \sim$ NUMBER OF PHOTONS EACH ~~STATE~~ STATE CAN CARRY

OCCUPATION $n_i \sim$ NUMBER OF PHOTONS IN A STATE

$w(n_i; g_i) \sim$ NUMBER OF REALISATIONS FOR n_i PHOTONS IN STATE i WITH STATISTICAL WEIGHT g_i

$\prod_i w(n_i; g_i) \sim$ NUMBER OF WAYS FOR PUTTING $N = \sum n_i$ PHOTONS IN THE STATES i , n_i EACH.

CONSTRUCT $w(n_i; g_i)$, THERE CAN BE ARBITRARILY MANY PHOTONS IN A SINGLE STATE.

$w(n, 1) = 1$ 1 WAY OF PUTTING n -PHOTONS IN 1 STATE

$w(n, 2) = n + 1 = \frac{(n+1)!}{n!}$ $n+1$ WAYS OF PUTTING n PHOTONS INTO 2 STATES

$$w(n, 3) = \sum_k w(n-k, 2) = \sum_k \frac{(n-k+1)!}{(n-k)!} = \frac{(n+2)!}{n! \cdot 2!}$$

$$w(n, g) = \frac{(n+g-1)!}{n! \cdot (g-1)!} \approx \frac{(n+g)!}{n! \cdot g!}$$

$$\prod_i w(n_i; g_i) = \prod_i \frac{(n_i + g_i)!}{n_i! \cdot g_i!} = \Omega$$

$$\ln \Omega = \sum_i (n_i + g_i) \cdot \ln(n_i + g_i) - n_i \ln n_i - g_i \ln g_i$$

WITH THE STIRLING-APPROXIMATION $\ln(n!) = n \ln n$

- FIND THE MAXIMUM ~ MOST LIKELY CONFIGURATION UNDER THE CONDITIONS $\sum_i n_i = N$ (TOTAL NUMBER) AND $\sum_i n_i \epsilon_i = E$ (TOTAL ENERGY)

$$X = \ln \Omega + \alpha (N - \sum_i n_i) + \beta (E - \sum_i n_i \epsilon_i)$$

WITH TWO LAGRANGE-MULTIPLIERS α AND β

$$\rightarrow \frac{dX}{dn_i} = 0 = \ln(n_i + g_i) - \ln n_i - \alpha - \beta \epsilon_i$$

$$\rightarrow \frac{n_i + g_i}{n_i} = \exp(-\alpha - \beta \epsilon_i)$$

$$n_i = g_i \frac{1}{\exp(\alpha + \beta \epsilon_i) - 1} \quad \text{BOSE FACTOR}$$

COMPARISON TO E.G. IDEAL GAS SHOWS

$$\alpha = \frac{\mu}{kT} \quad \text{NFUGACITY AD} \quad \beta = \frac{1}{kT} \quad \text{BOLTZMANN FACTOR}$$

- PHOTONS ARE ULTRARELATIVISTIC PARTICLES $\epsilon = cp$ 2 SPIN STATES

$$\text{DISPERSION: } p = \frac{h}{\lambda} = \frac{\hbar \omega}{c}$$

FUGACITY $\frac{\mu}{kT}$ (OR CHEMICAL POTENTIAL) = 0 FOR PHOTONS

$$N = 2 \cdot \int \frac{d^3p}{h^3} \frac{1}{\exp(\beta \epsilon p) - 1} = \frac{1}{\pi^2 c^3} \int_0^\infty d\omega \frac{\omega^2}{\exp(\beta \hbar \omega) - 1} = 4 \left\{ (3) \left(\frac{kT}{\hbar c} \right)^3 \right.$$

$$\left. \right\} = \frac{15}{15} \frac{\pi}{15 (\hbar c)^3} (kT)^4$$

PLANCK SPECTRUM $S(\omega)$ $N = \int_0^\infty d\omega S(\omega) \rightarrow$

$$\rightarrow S(\omega) = \frac{\hbar}{\pi^2 \cdot c^3} \cdot \frac{\omega^3}{\exp(\beta \hbar \omega) - 1}$$

TYPICAL INTEGRALS IN THIS CONTEXT

$$\int_0^{\infty} d\omega \frac{\omega^{n-1}}{\exp(\omega)-1} = \int_0^{\infty} d\omega \cdot \omega \cdot \frac{1}{\exp(\omega)-1} = \int_0^{\infty} d\omega \omega^{n-1} \sum_{m=0}^{\infty} \exp(-m\omega) =$$
$$= \sum_{m=0}^{\infty} m^{-n} \cdot \int_0^{\infty} dy y^{n-1} \cdot \exp(-y) = \zeta(n) \cdot \Gamma(n-1), \quad y = m\omega$$

RAYLEIGH-JEANS-LIMIT $\hbar\omega \ll kT$

$$S(\omega) = \frac{\omega^3}{\exp(\beta\hbar\omega)-1} \rightarrow \frac{\omega^3}{1+\beta\hbar\omega-1} \sim \omega^2$$

WIEN-LIMIT $\hbar\omega \gg kT$

$$S(\omega) = \frac{\omega^3}{\exp(\beta\hbar\omega)-1} \rightarrow \omega^3 \cdot \exp(-\beta\hbar\omega)$$

DISPLACEMENT LAW

$$\frac{ds}{d\omega} = 0 \rightarrow \frac{3-x}{e^x-1} = 0 \quad \text{WITH } x = \frac{\hbar\omega}{kT} = \beta\hbar\omega$$

$$\sim \text{SOLVED BY } x \approx 2.81, \quad \frac{\omega}{T} = \text{CONST.}$$

WIEN-RADIATION LAW. THE (-1) IN THE BOSE FACTOR

COMES FROM THE INDISTINGUISHABILITY OF PHOTONS IN THEIR QUANT. STATISTICS.

CLASSICAL PARTICLES JUST WOULD HAVE THE BOLTZMANN-FACTOR, E.G.

$$N = 2 \cdot \int \frac{d^3p}{h^3} \exp(-\beta cp) = \frac{1}{\pi^2 \cdot c^3} \cdot \int_0^{\infty} d\omega \cdot \omega^2 \cdot \exp(-\beta\hbar\omega) \sim T^3$$

THE BASIC INTEGRAL PROPERTIES ARE IDENTICAL "ONLY" THE PREFACTORS DON'T MATCH.

C15 THOMSON - SCATTERING AND THE THOMSON CROSS SECTION

- PLACE A TEST CHARGE IN THE WAY OF AN INCOMING ELECTROMAGNETIC WAVE \sim RADIATION PRESSURE EXPERIENCED BY THE CHARGE + SCATTERING OF THE INCIDENT RADIATION.

MOTION OF THE ELECTRON: IN THE NON-RELATIVISTIC LIMIT

$$m_e \cdot \frac{d^2 \vec{x}}{dt^2} = q \cdot (\vec{E} + \vec{v} \times \vec{B}) ; \quad q = -e \text{ FOR ELECTRON}$$

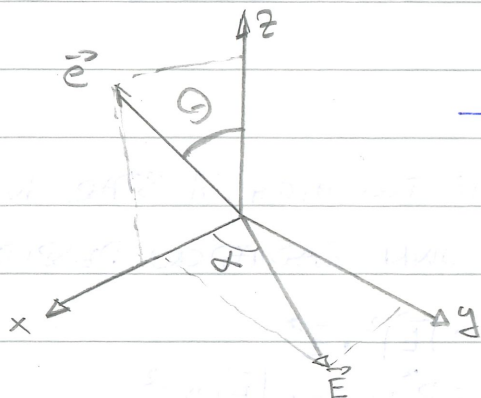
FOR VERY SMALL VELOCITIES $|\vec{E}| \gg |\vec{v} \times \vec{B}| \rightarrow$

$$c \cdot \vec{\beta} = \frac{q}{m_e} \vec{E} \quad \text{WITH} \quad \vec{\beta} = \frac{\vec{v}}{c}$$

RADIATION EMITTED BY THE TEST CHARGE

$$\frac{dP}{d\Omega} = \frac{q^2}{4\pi c} \cdot |\vec{e} \times \vec{\beta}|^2 \quad \text{INTO THE DIRECTION } \vec{e}$$

$$\rightarrow \frac{dP}{d\Omega} = \frac{q^4}{4\pi m_e^2 c^3} |\vec{e} \times \vec{E}|^2 \quad \text{WITH THE LORENTZ-FORCE SUBSTITUTED}$$



$$\rightarrow \frac{dP}{d\Omega} = \frac{q^4 E^2}{4\pi m_e^2 c^3} \cdot (1 - \cos^2 \alpha \sin^2 \theta)$$

- THOMSON CROSS SECTION: RATIO BETWEEN SCATTERED AND INCOMING RADIATION

$$\vec{S} = \frac{c}{4\pi} |\vec{E}|^2 \cdot \vec{e}_z$$

$$\frac{d\sigma}{d\Omega} = \frac{dP}{S \cdot d\Omega} = \frac{q^4}{(mc^2)^2} \cdot (1 - \cos^2 \alpha \sin^2 \theta)$$

DIFFERENTIAL CROSS-SECTION

ELECTROSTATIC ENERGY OF A HOMOGENEOUSLY CHARGED SPHERE

=

REST MASS ENERGY (RELATIVISTIC)

$$\frac{e^2}{r_e} = mc^2 \quad \text{DETERMINES THE CLASSICAL ELECTRON RADIUS}$$
$$r_e = \frac{e^2}{mc^2}$$

$$\rightarrow \frac{d\sigma}{d\Omega} = r_e^2 \cdot (1 - \cos^2 \alpha \sin^2 \theta)$$

\hookrightarrow 'AREA' ASSOCIATED WITH THE ELECTRON, $\sim 10^{-28} \text{ m}^2$

AVERAGE OVER ALL POLARISATION STATES AND DIRECTIONS

$$\frac{d\sigma}{d\Omega} = \frac{r_e^2}{2\pi} \cdot \int_0^{2\pi} d\alpha \cdot \frac{d\sigma}{d\Omega}(\alpha) = \frac{r_e^2}{2} (1 + \cos^2 \theta)$$

$$\sigma = \int d\Omega \frac{d\sigma}{d\Omega} = \pi \cdot r_e^2 \cdot \int_{-1}^{+1} du (1 + u^2) = \frac{8\pi}{3} \cdot r_e^2 = \sigma_T$$

$\sigma_T \approx 10^{-28} \text{ m}^2 \sim$ TOTAL THOMSON CROSS-SECTION

EVENT RATE = PARTICLE FLUX \cdot CROSS-SECTION

EDDINGTON-LIMIT

IF THE RADIATION PRESSURE IS TOO HIGH, A STAR WOULD BE BLOWN APART BY ITS OWN RADIATION DESPITE GRAVITY

$$\text{MOMENT DENSITY} \quad \vec{S} = P = \frac{1}{4\pi} \cdot |\vec{E}|^2 \cdot \vec{e}$$

$$\text{TOTAL LUMINOSITY} \quad L = \int d\vec{A} \cdot \vec{S} = 4\pi R^2 \cdot \frac{c}{4\pi} \cdot |\vec{E}(R)|^2$$

$$\rightarrow \frac{\vec{S}}{c} = \frac{L}{4\pi R^2 \cdot c} \vec{e}$$

$$\text{FORCE ON AN ELECTRON} \quad \vec{F}_s = \frac{\vec{S}}{c} \cdot \sigma_T = \frac{L \cdot \sigma_T}{4\pi \cdot R^2 \cdot c} \cdot \vec{e}$$

$$\text{VS. GRAVITY} \quad \vec{F}_g = -\frac{GMm}{R^2} \cdot \vec{e}$$

FORCE BALANCE DEFINES EDDINGTON LIMIT

$$L_{\text{EDD}} = \frac{4\pi GMm_e c}{\sqrt{\tau}} \approx 33 \cdot 10^3 \cdot L_{\odot} \left(\frac{M}{M_{\odot}} \right)$$

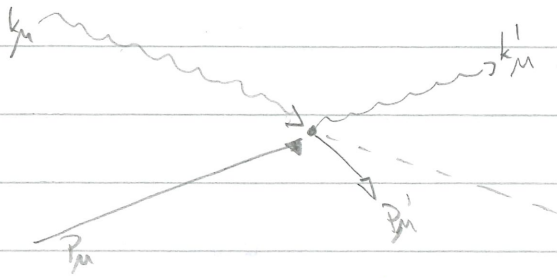
m SHOULD BE THE MASS OF THE IONS WHICH ARE TIGHTLY COUPLED TO THE ELECTRONS.

• COMPTON SCATTERING

-MOMENTUM TRANSFER FROM A PHOTON ONTO AN ELECTRON

PHOTON $ck^{\mu} = \omega \cdot \left(\frac{1}{c} \right)$

ELECTRON $p^{\mu} = \left(\frac{E/c}{p} \right) = \hbar \left(\frac{c}{v} \right)$



CONSERVATION OF 4-MOMENTUM

$$p^{\mu} + \hbar k^{\mu} = p'^{\mu} + \hbar k'^{\mu}$$

$$\begin{cases} E + \hbar\omega = E' + \hbar\omega' & \mu=0 \\ c\vec{p} + \hbar\omega\vec{e} = c\vec{p}' + \hbar\omega'\vec{e}' & \mu=1,2,3 \text{ SPHERE} \end{cases}$$

$$E'^2 = E^2 + 2\hbar c \vec{p} \cdot (\omega\vec{e} - \omega'\vec{e}') + \hbar^2 \cdot (\omega\vec{e} - \omega'\vec{e}')^2$$

~~(E^2 = c^2 p^2 + m^2 c^4, AND SUBSTITUTE p')~~

USE ENERGY CONSERVATION TO ELIMINATE E'

$$E(\omega - \omega') = \hbar\omega\omega' \cdot (1 - \cos\theta) + c\vec{p} \cdot (\omega\vec{e} - \omega'\vec{e}') \quad \theta = \angle(\vec{e}; \vec{e}')$$

ELECTRON AT REST, $\vec{p} = 0$; $E = mc^2 \rightarrow$

$$\frac{\omega'}{\omega} = \frac{1}{1 + E(1 - \cos\theta)} \quad \text{WITH } E = \frac{\hbar\omega}{mc^2}$$

- COMPUTE AVERAGE CHANGE IN ENERGY (OVER ALL ANGLES)

$$\frac{\langle \Delta E \rangle}{E} = \frac{\langle \omega' \rangle - 1}{\omega} = \frac{1}{v_T} \cdot \frac{r_e^2}{2} \cdot \int \sin \theta d\theta \int_0^{2\pi} d\phi \left(\frac{1 + \cos \theta}{1 + \epsilon(1 - \cos \theta)} - 1 \right) =$$

$$= \frac{\pi \cdot r_e^2}{v_T} \cdot \left(\frac{1}{E^3} \cdot (\ln(1 + 2\epsilon)(2E^2 + 2\epsilon + 1) - 2\epsilon(1 + \epsilon)) - 1 \right)$$

IF $\epsilon \ll 1$ (LOW ENERGY LIMIT)

$$\frac{\langle \Delta E \rangle}{E} \approx \frac{\int \pi}{3} \frac{r_e^2}{v_T} \cdot (1 - \epsilon - 1) \approx -\frac{h\omega}{mc^2} \approx -\epsilon$$