

THEORETICAL ASTROPHYSICS

A1 FLUID MECHANICS

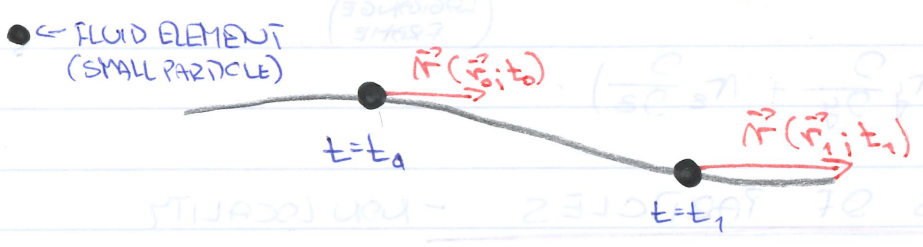
- FIRST WE WILL TALK ABOUT FLUID MOTION, CONTINUUM (FIELD) THEORY

- CONTINUUM WILL BE DESCRIBED BY DENSITY $\rho(\vec{r}; t)$ AND VELOCITY $(\vec{v}; t)$. MASS WILL BE SMOOTHLY DISTRIBUTED.

- FLUID ELEMENT (CENTRAL CONCEPT) - LET'S IMAGINE WE HAVE A CLOUD OF PARTICLES (CLOUD SUFFICIENTLY LARGE)

- ELEMENT IS LARGE ENOUGH TO CONTAIN MANY PARTICLES
- WE WILL DESCRIBE VARIATION IN DENSITY AND VELOCITY, SO IT HAS TO BE SMALL ENOUGH COMPARED TO LENGTH
- SMALL VELOCITIES $|\vec{v}| \ll c$ (NO RELATIVITY)
- GALILEAN RELATIVITY
 - ADDITIVITY OF VELOCITIES
 - UNIVERSAL TIME
- SMALL POTENTIALS - GRAVITATIONAL POTENTIAL $|\phi| \ll c^2$

A2 EULER & LAGRANGE FRAME



TWO WAYS HOW TO DESCRIBE THIS:

1) EULER FRAME - FIXED LABORATORY COORDINATES, WE ARE STANDING ON A SHORE (BRIDGE). WE HAVE VELOCITIES

$$v(\vec{r}; t) \text{ AT ALL POSITION } \vec{r} \text{ AT A FIXED TIME}$$

2) LAGRANGE FRAME - WE OBSERVE WHILE MOVING WITH THE FLOW,

$$v^vec(r(t); t)$$

A3 ACCELERATION OF A FLUID

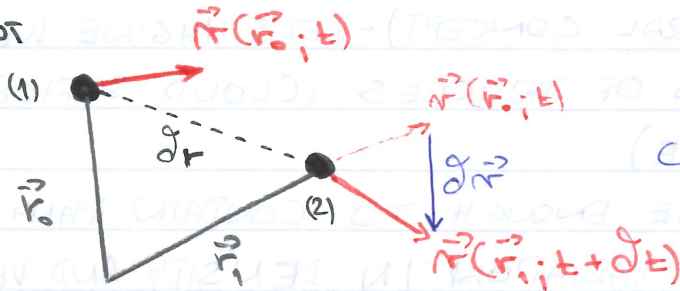
- FIRST LAGRANGE FRAME - WE FEEL ACCELERATION WE MOVE TO A PLACE WHERE THE VELOCITY IS LARGER (EXPLICIT VARIATIONS OF \vec{v})

$$\delta \vec{v} = \vec{v}(\vec{r}_0; t + \delta t) - \vec{v}(\vec{r}_0; t)$$

- SECOND EULER FRAME - ONE "STATIC" PLACE, AND FLUID MOVES ALONG, BUT LATER THE HIGHER VELOCITY APPEARS. FLUID ELEM. MOVES TO A NEW POSITION

$$\delta \vec{v} = \vec{v}(\vec{r}_1; t + \delta t) - \vec{v}_0(\vec{r}_0; t + \delta t)$$

FLUID ELEMENT



CHANGE IN A VELOCITY

HOW THE ACCELERATION LOOKS LIKE:

$$\delta \vec{v} = \vec{v}(\vec{r}_1; t + \delta t) - \vec{v}(\vec{r}_0; t) \stackrel{\text{TAYLOR EXPAN}}{=} \frac{\partial \vec{v}}{\partial t} \delta t + \frac{\partial \vec{v}}{\partial \vec{r}} \delta \vec{r}$$

$$\frac{d\vec{v}}{dt} = \lim_{\delta t \rightarrow 0} \frac{\delta \vec{v}}{\delta t} = \frac{\partial \vec{v}}{\partial t} + \underbrace{\frac{\partial \vec{v}}{\partial \vec{r}} \cdot \frac{d\vec{r}}{dt}}_{\text{ADVECTIVE DERIVATIVE}} = \left(\frac{\partial}{\partial t} + \vec{v} \cdot \nabla \right) \vec{v}$$

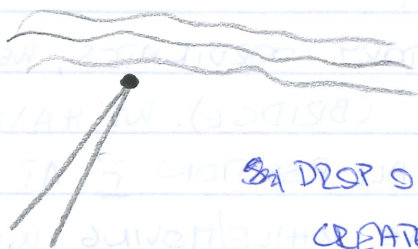
ACCELEROMETER (LAGRANGE FRAME)

FLOW BRINGS "ACCELERATION" (EULER FRAME)

(NON-LINEARITY)

$$\vec{v} \cdot \nabla = \left(v_x \frac{\partial}{\partial x} + v_y \frac{\partial}{\partial y} + v_z \frac{\partial}{\partial z} \right)$$

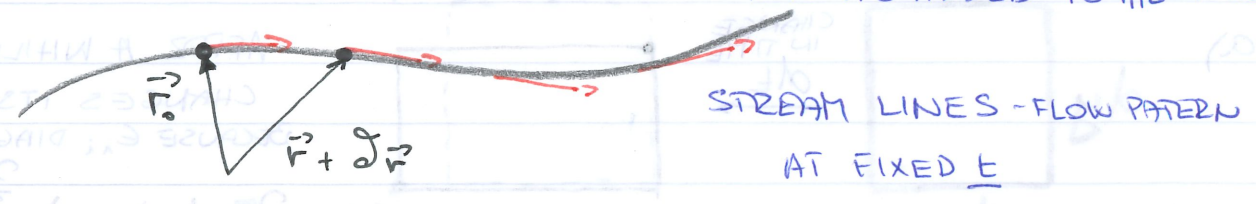
ALL TRAJECTORIES OF PARTICLES - NON LOCALITY



NEEDLE WITH INK

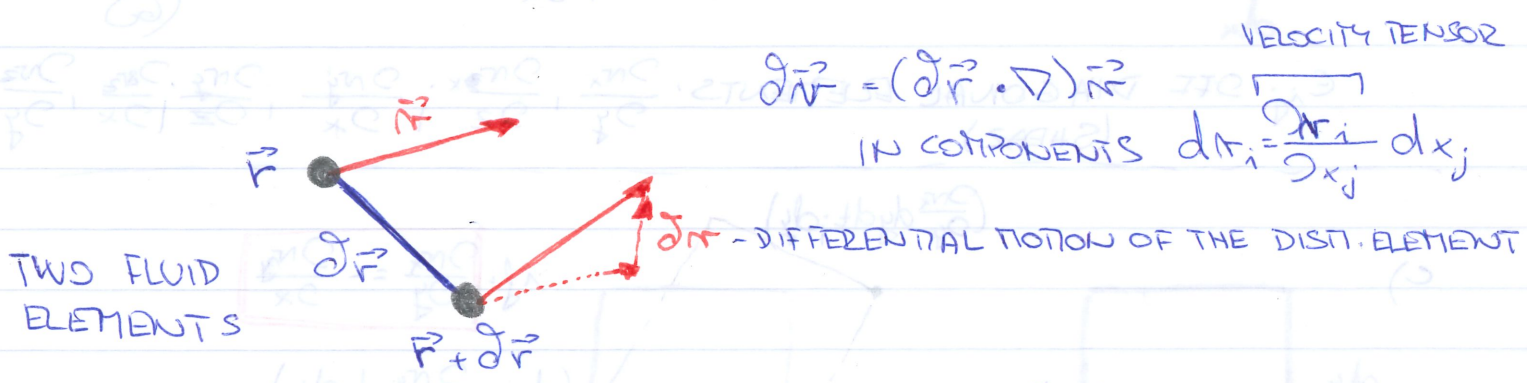
A DROP OF INK FOLLOWS THE STREAM AND CREATES STREAK LINE (STREAM LINE)

$$\vec{r}(t) - \vec{r}(t_0) = \int_{t_0}^t dt' \cdot \underbrace{\vec{v}(\vec{r}_0; t')}_{\text{LAGRANGE VELOCITY - THIS WHAT DROP OF INK FEELS WHEN IT'S ADDED TO THE WATER}}$$



A5 DEFORMATION OF THE FLUID ELEMENTS DUE TO VELOCITY VARIATIONS

- AT FIXED t :
- ① A FLUID ELEMENT AT \vec{r} WITH VELOCITY $\vec{v}(\vec{r}; t)$
 - ② A FLUID ELEMENT AT $\vec{r} + \delta\vec{r}$ WITH VELOCITY $\vec{v} + \delta\vec{v}$



VELOCITY TENSOR DESCRIBES IN WHAT WAY ^{FLUID} ELEMENT SQUIZES...
 RELATIVE MOTION IS DESCRIBED BY THE VELOCITY TENSOR $\frac{\partial v_i}{\partial x_j}$
 IT CAN BE DECOMPOSED:

1) SYMMETRIC PART $\epsilon_{ij} = \frac{1}{2} \left(\frac{\partial v_i}{\partial x_j} + \frac{\partial v_j}{\partial x_i} \right)$, $\epsilon_{ij} = \epsilon_{ji}$
 (SHEAR) (SMYKOVÉ)

2) ANTISYMMETRIC PART $\omega_{ij} = \frac{1}{2} \left(\frac{\partial v_i}{\partial x_j} - \frac{\partial v_j}{\partial x_i} \right)$, $\omega_{ij} = -\omega_{ji}$
 (VORTICITY) (VROBOSTI?)

IN SYMMETRIC PART WE CAN TAKE A LOOK AT THE TRACE:
 $\epsilon_{ii} = \frac{\partial v_i}{\partial x_i} = \text{div } \vec{v} = \nabla \cdot \vec{v} \rightarrow$ RATE OF CHANGE OF A FLUID ELEMENT.

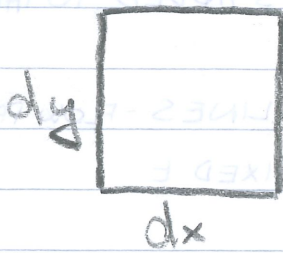
IF $\text{div } \vec{v} = 0$ THEN THE FLOW IS INCOMPRESSIBLE
 \hookrightarrow VOLUME OF THE FLUID DOES NOT CHANGE

NIC S TROKAT

A6 VISUALISATION OF THE VELOCITY TENSOR

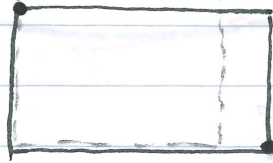
FLUID ELEMENT

a)



CHANGE IN TIME dt

$$(0; dy - \frac{\partial v_y}{\partial y} dy dt)$$

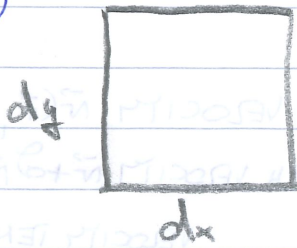


$$(dx + \frac{\partial v_x}{\partial x} dx dt; 0)$$

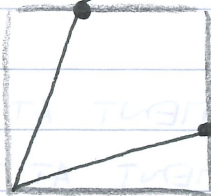
AFTER A WHILE (dt) ELEMENT CHANGES ITS SHAPE (VOLUME) BECAUSE E_{ii} DIAG. ELEMENTS:

$$\frac{\partial v_x}{\partial x}, \frac{\partial v_y}{\partial y}, \frac{\partial v_z}{\partial z} \quad (3)$$

b)



AFTER A TIME dt



$$(dx; \frac{\partial v_x}{\partial x} dx dt)$$

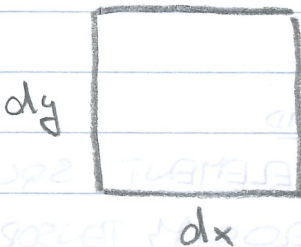
$$(\frac{\partial v_x}{\partial y} dy dt; dy)$$

$$\text{if } \frac{\partial v_x}{\partial y} = \frac{\partial v_y}{\partial x}$$

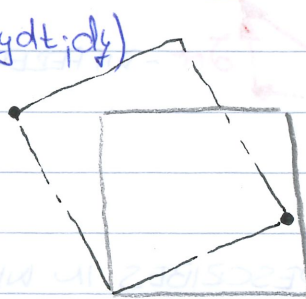
(2)

E_{ij} OFF DIAGONAL ELEMENTS (SHEAR) $\frac{\partial v_x}{\partial y}, \frac{\partial v_x}{\partial z}, \frac{\partial v_y}{\partial x}, \frac{\partial v_y}{\partial z}, \frac{\partial v_z}{\partial x}, \frac{\partial v_z}{\partial y}$

c)



IN A TIME dt



$$(dx; \frac{\partial v_y}{\partial x} dx dt)$$

$$\text{if } \frac{\partial v_x}{\partial y} = -\frac{\partial v_y}{\partial x}$$

CHANGE

ω_{ij} OFF DIAGONAL ELEMENTS (VORTICITY)

b) & c) BOTH CHANGE SHAPE BUT CONSERVE VOLUME, BOTH E_{ij}, ω_{ij} AND $\text{div}(\vec{v})$ HAVE UNITS OF ANGULAR VELOCITIES.

A7 DECOMPOSITION OF THE VELOCITY TENSOR

MOTION $\vec{r}(t+\Delta t) = \vec{r}_0(t) + \vec{v} \Delta t$

RELATIVE MOTION $\frac{\partial v_i}{\partial x_j} = \underbrace{J_{ij}} + \frac{\partial v_i}{\partial x_j} \Delta t + O(\Delta t^2) + \dots = \exp(\frac{\partial v_i}{\partial x_j} \Delta t)$

HOW MY NEIGHBORING POINT CHANGE'S ITS POSITION IN DIFFERENT RATE THAN I DO.

TOHLE JE BACHA'ER EXHONENCIJELY

CHANGE IN VOLUME $d^3x(t+\Delta t) = \det\left(\frac{\partial r_i}{\partial x_j}\right) \cdot d^3x(t)$

FUNCTIONAL DETERMINANT

HANDY NOTE

$\ln \det A = \ln \prod_i \lambda_i = \sum_i \ln \lambda_i = \text{tr} \ln A$

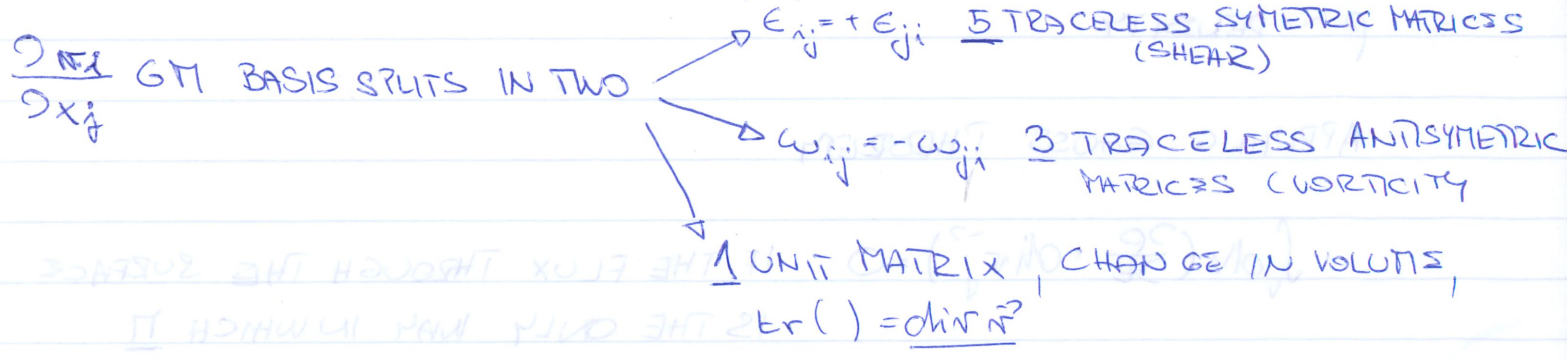
STORA - TRACE

$\ln \det\left(\frac{\partial r_i}{\partial x_j}\right) = \text{tr} \ln\left(\frac{\partial r_i}{\partial x_j}\right) = \text{tr}\left(\frac{\partial r_i}{\partial x_j}\right) \cdot \Delta t = \frac{\partial r_i}{\partial x_i} \Delta t = \text{div} \vec{v}$

→ INCOMPRESSIBLE FLOWS CONSERVE VOLUMES

VELOCITY TENSOR HAS A BASIS IN TERMS OF THE GEL-MANU MATRICES WHICH ARE ALL TRACELESS; VOLUMES ARE CONSERVED BY SHEAR AND VORTICITY

ONLY $\text{div} \vec{v}$ CHANGES VOLUMES



$\lambda_z = \begin{pmatrix} 0 & 1 \\ -1 & 0 \\ & & 0 \end{pmatrix}$

$\lambda_y = \begin{pmatrix} 0 & & 1 \\ & 1 & 0 \\ -1 & 0 & 0 \end{pmatrix}$

$\lambda_x = \begin{pmatrix} 0 & & 1 \\ & 0 & 1 \\ -1 & 0 & 0 \end{pmatrix}$

$\text{rot}(\vec{v})_z = \frac{\partial v_x}{\partial y} - \frac{\partial v_y}{\partial x}$

$\text{rot}(\vec{v})_y = \frac{\partial v_z}{\partial x} - \frac{\partial v_x}{\partial z}$; $\text{rot}(\vec{v})_x = \frac{\partial v_z}{\partial z} - \frac{\partial v_z}{\partial y}$

$\omega_{ij} = -\omega_{ji}$ HAS 3 INDEPENDENT AND IS ANTI SYMMETRIC

- CONTRACT WITH ϵ_{ijk} AND OBTAIN AN AXIAL VECTOR

$\omega_i = \epsilon_{ijk} \omega_{jk}$, $\vec{\omega} = \text{rot}(\vec{v}) = \nabla \times \vec{v}$

A8 EQUATION OF MOTION - DYNAMICS

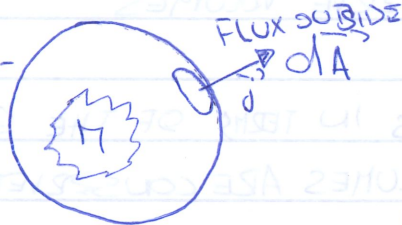
ACCELERATION OF FLUID ELEMENTS

- PRESSURE GRADIENTS ∇p
- GRAVITY $\nabla \Phi$
- ELECTRIC & MAGNETIC FORCES
- VISCOUS FORCES

} EULER } NAVIER
 } STOKES

A9 CONTINUITY EQUATION AND THE CONSERVATION OF MASS, VOLUME V , SURFACE ∂V , CONTAINS MASS m INSIDE

DECAY, CHEMICAL REACTION
 THAT'S HOW IT GETS OUT



CHANGE IN MASS

$$\frac{dm}{dt} = - \frac{d}{dt} \int d^3x \rho = - \int d\vec{A} \cdot \vec{j} =$$

$$= - \int d^3x \operatorname{div} \vec{j} \rightarrow \frac{\partial \rho}{\partial t} + \operatorname{div} \vec{j} = 0$$

$\vec{j} \sim$ MATTER FLUX (OUTWARD)
 $\rho \sim$ DENSITY

APPLYING GAUSS THEOREM

$$\int dV \left(\frac{\partial \rho}{\partial t} + \operatorname{div} \vec{j} \right) = 0$$

IF THE FLUX THROUGH THE SURFACE IS THE ONLY WAY IN WHICH Π CAN CHANGE

$$\vec{j} = \rho \cdot \vec{v} \Rightarrow \operatorname{div}(\vec{j}) = \operatorname{div}(\rho \cdot \vec{v}) = \operatorname{div} \rho \cdot \vec{v} + \rho \cdot \operatorname{div}(\vec{v}) \quad \text{IN GENERAL}$$

AGAIN EQ. OF CONTINUITY:

$$\frac{\partial \rho}{\partial t} + \operatorname{div}(\rho \vec{v}) = 0$$

INCOMPRESSIBLE

COMPRESSIBLE

$$\frac{\partial \rho}{\partial t} + \vec{v} \cdot \nabla \rho = 0$$

$$\frac{\partial \rho}{\partial t} + \vec{v} \cdot \nabla \rho = -\rho \operatorname{div}(\vec{v})$$

ADVECTIVE DERIVATIVE

CONTINUITY IS NONLINEAR

(BUT IN ED IS LINEAR)

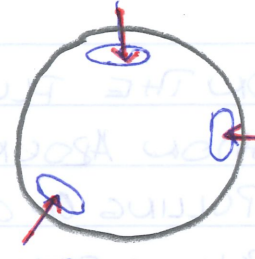
A10 KINEMATICS

$$\vec{a} = \frac{d\vec{v}}{dt} = \underbrace{\frac{\partial}{\partial t} \vec{v} + (\vec{v} \cdot \nabla) \vec{v}}_{\text{2 CONTRIBUTION LAGR. \& EULER FRAME}} = \sum \text{SPECIFIC FORCES ACTING OF FLUID ELEMENT}$$

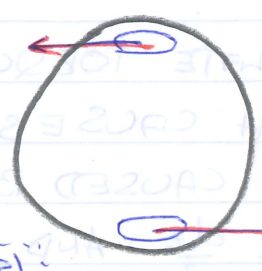
(FORCE DENSITIES)

A11 PRESSURE FORCES AND VISCOUS FORCES (STRESS & PRESSURE)

- THESE FORCES ~~ACT~~ ON A SURFACE OF THE ELEMENT



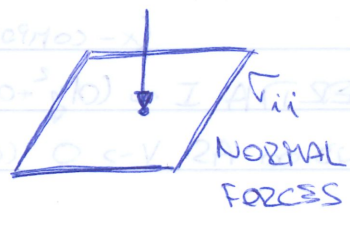
ISOTROPIC PRESSURE (FEELS PRESSURE FROM ALL SIDES)



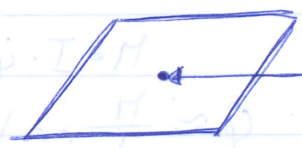
MANY FLUID ELEMENTS EXPERIENCE SHEARING FORCES WITH SURROUNDING FLUID

FLUID ELEMENT TRAPPED IN FLOW (VISCOUS FORCES)

STRESS TENSOR → SURFACE ELEMENT $d\vec{A}$
 → 3 COMPONENTS OF FORCE FOR 3 COMPONENTS OF $d\vec{A}$



τ_{ii}
NORMAL FORCES



$\tau_{ij} = \tau_{ji}$

TANGENTIAL FORCES

τ_{ij} - FORCE DENSITY IN i -DIRECTION ACTING ON A SURFACE ELEMENT IN j -DIRECTION.

$$d\vec{F} = \sigma \cdot d\vec{A} \rightarrow \frac{d\vec{F}}{dA} = \sigma \cdot \vec{n} \text{ WITH } d\vec{A} = \vec{n} dA$$

σ - LINEAR TRANSFORMATION BETWEEN ORIENTATION \vec{A} AND FORCE \vec{F} .

τ_{ij} IS SYMMETRIC $\tau_{ij} = \tau_{ji}$ (ORIGIN OF THIS IS IN REVERSIBILITY)
 THIS SYMMERY LINKS TO A VELOCITY TENSOR e_{ij}

ISOTROPIC PART OF $\tau_{ij} \sim$ PRESSURE

$$\tau_{ij} = \underbrace{\tau_{ij}'}_{\text{VISCOSITY}} - \underbrace{p \delta_{ij}}_{\text{PRESSURE}}$$

p - IS PUSHING AGAINST $d\vec{A}$ \rightarrow MINUS SIGN

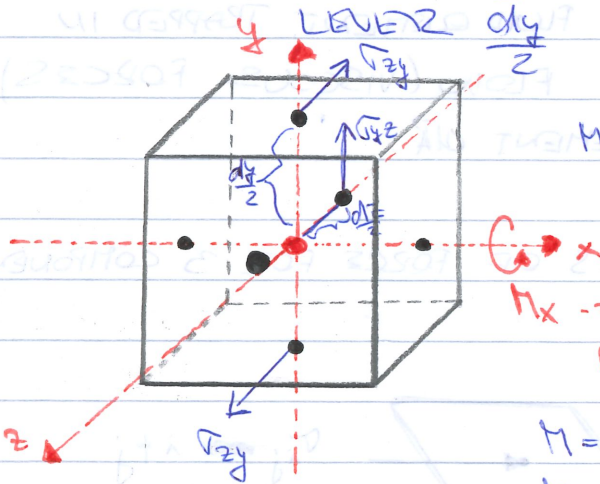
- AND IS ISOTROPIC

τ_{ij} CAN STILL HAVE ELEMENTS ON THE DIAGONAL

$$\text{tr}(\tau) = \tau_{ii} = -3p \quad \text{tr}(\tau') = \tau'_{ii} = 0$$

$\tau_{ij} = \tau_{ji} \rightarrow$ CALCULATE TORQUE M_x ON THE FLUID ELEMENT WHICH CAUSES A ROTATION AROUND THE X-AXIS

M_x IS CAUSED BY τ_{yz} PULLING ON $dx dy$ WITH LEVER $\frac{dz}{2}$ AND BY τ_{zy} PULLING ON $dx dz$ WITH LEVER $\frac{dy}{2}$



τ_{yz} & τ_{zy} GENERATE TORQUE

$$M_x = \tau_{zy} (dx dz) \frac{dy}{2} - \tau_{yz} (dx dy) \frac{dz}{2} = (\tau_{zy} - \tau_{yz}) \frac{dV}{2}$$

SYSTEM STARTS ROTATING.

M_x - TORQUE IN X-AXIS

(x - COMPONENT)

$$M = I \cdot \dot{\varphi} \quad \text{WITH INERTIA } I \sim (dy^2 + dz^2) \frac{dV}{2}$$

$$\dot{\varphi} \sim \frac{M}{I} \sim V^{-2/3} \rightarrow \text{AND } \infty \text{ AS } V \rightarrow 0 \text{ (GOES TO ZERO)}$$

THEREFORE $\tau_{ij} = \tau_{ji}$, STRESS TENSOR IS SYMMETRIC OTHERWISE SMALL FLUID ELEMENTS WOULD START SPINNING.

AIR VISCOSITY

FOCUS ON NEWTONIAN FLUIDS: τ_{ij} DEPENDS LINEARLY AND INSTANTANEOUSLY ON THE (LIKEWISE SYMMETRIC) VELOCITY TENSOR ϵ_{ij}

NEWTONIAN FLUID - DRY PAINT, KETCHUP

$$\tau_{ij}' = 2\eta \left(\epsilon_{ij} - \frac{1}{3} \text{tr}(\epsilon) \cdot \delta_{ij} \right) + \zeta \text{tr}(\epsilon) \delta_{ij} \quad \text{LAME' ANZAC}$$

η - SHEAR VISCOSITY

ζ - BULK VISCOSITY

$$\eta, \tau_{ij}' = \eta \cdot \frac{\partial v_i}{\partial x_j} \quad \zeta, \tau_{ii}' = \zeta \cdot \frac{\partial v_i}{\partial x_i} = \zeta \text{div}(\vec{v})$$

INCOMPRESSIBLE FLUIDS $\tau_{ij}' = 2\eta \epsilon_{ij}$, BECAUSE $\text{div}(\vec{v}) = 0$

η & ζ - LAME' VISCOSITY COEFFICIENTS, SPECIFIC TO ANY SUBSTANCE

IN INCOMPRESSIBLE FLUIDS, KINEMATICS IS NOT CHANGING A VOLUME OF FLUID ELEMENT (WATER).

NAVIER - STOKES EQUATION - APPLYING NEWTONIAN EQ. OF MOTION TO A FLUID ELEMENT WHICH TRANSPORTS THE MOMENTUM $\int_{\text{vol}} \rho \vec{v}$

$$\frac{d}{dt} \int_{\text{vol}} \rho \vec{v} = \int_{\text{vol}} dV \rho \vec{f} + \int_{\text{surf}} d\vec{A} \cdot \sigma$$

TOTAL MOMENTUM

$\rho \vec{v}$ - MOMENTUM DENSITY

2 TYPES OF FORCES $\left\{ \begin{array}{l} \vec{f} - \text{BODY FORCES ACTS ON A VOLUME} \\ \sigma - \text{SURFACE FORCES ACTS ON A SURFACE OF A VOLUME} \end{array} \right.$

- TIME DERIVATIVE $\frac{d}{dt}$ ACTS ONLY ON THE VELOCITY BECAUSE IN THE LAGRANGE - PICTURE THE FLUID ELEMENT ALWAYS CONTAINS THE SAME ATOMS/MOLECULES. ρ IS CONSTANT. THERE IS NO LOSS/GAIN IN THE FLUID.

$$\frac{d}{dt} \int_{\text{vol}} \rho \vec{v} \Rightarrow \int_{\text{vol}} dV \rho \frac{d\vec{v}}{dt}$$

- REWRITE THE SURFACE TERM AS A VOLUME INTEGRAL: GAUSS-THEOR...

$$\left(\int_{\text{vol}} d\vec{A} \cdot \sigma \right)_i = \int_{\text{vol}} dV \sigma_{ij} n_j = \int_{\text{vol}} dV \frac{\partial \tau_{ij}}{\partial x_j} = \left(\int_{\text{vol}} dV \nabla \cdot \sigma \right)_i$$

$$\int_V dV \rho \frac{d\vec{r}}{dt} = \int_V dV \rho \vec{f} + \int_V dV \nabla \cdot \tau \rightarrow \boxed{\rho \frac{d\vec{r}}{dt} = \rho \vec{f} + \nabla \cdot \tau}$$

COSMETIC CHANGES

∇ - CONTAINS PRESSURE AND VISCOSITY

- SEPARATING ∇ , ∇' - SIGMA PRIME; PRESSURE - p

$$(\nabla \cdot \tau)_i = (\nabla' \cdot \tau')_i - \frac{\partial}{\partial x_j} (p \cdot \delta_{ij}) = (\nabla' \tau'_i - \nabla p)_i$$

- SUBSTITUTE FOR THE ADVECTIVE DERIVATIVE

$$\frac{\partial}{\partial t} \vec{r} + (\vec{r} \cdot \nabla) \vec{r} = \vec{f} - \frac{\nabla p}{\rho} + \frac{1}{\rho} \nabla \cdot \tau$$

SOURCE OF NONLINEARITY BODY FORCE (GRAVITATIONAL) PRESSURE VISCOSITY

IN ASTROPHYSICS: $\vec{f} = -\nabla \phi$; ϕ - GRAVITATIONAL POTENTIAL

MAXWELL EQ - LINEAR, SCALLED SYS. IS STILL THE SAME

DEFINITION OF THE GRAV. POT. AND GRAV. FORCE PRESSURE ACT IN THE OPPOSITE DIRECTION OF THE SURFACE ELEMENT

ALL HYDROSTATIC EQUILIBRIUM

STATIONARY $\vec{r} = 0$
 STATIC $\frac{\partial}{\partial t} \vec{r} = 0$ } $-\nabla \phi - \frac{\nabla p}{\rho} = 0$ FROM NAVIER-STOKES EQUATION
 (EARTH ATMOSPHERE)

IDEAL GAS $pV = NkT \Rightarrow p \left\{ \left[\frac{m}{m} \right] \right\} \rightarrow$ FOR FIXED TEMPERATURE $p \propto \rho$

$$\frac{\nabla p}{\rho} = \frac{\nabla p}{p} = \nabla \ln p = -\nabla \phi \sim \rho \propto e^{(-\phi)} = e^{(-gh)}$$

BAROMETRIC FORMULA

IN A HOMOGENEOUS GRAVITATIONAL FIELD: $\phi = gh \rightarrow$

\rightarrow BAROMETRIC FORMULA $p \propto e^{(-gh)}$

MARS - THIN ATMOSPHERES - ^{ALMOST} NO PARTICLES

VENUS - SIMILAR TO EARTH BUT HUGE PRESSURE, THIS IS CONNECTED TO THE HEAVY MOLECULES IN VENUS ATMOSPHERE \Rightarrow HIGH PRESSURE AND DENSITY.

YOUNG EARTH - A LOT OF WATER IN THE ATMOSPHERE => HIGH PRESSURE, BUT THEN RAINED (FOR A LOT OF 10²⁰ YEARS).

A15 VISCOSITY

LAMÉ ANZAC

$$(\mathbf{D} \cdot \nabla)^i = \frac{\partial v_i}{\partial x_j} = \eta \left(\frac{\partial^2 v}{\partial x_i \partial x_j} + \left(\zeta + \frac{2}{3} \eta \right) \cdot \frac{\partial}{\partial x_i} \frac{\partial v_k}{\partial x_k} \right)$$

SUBSTITUTE A DEFINITION OF ϵ_{ij}

$$\mathbf{D} \cdot \nabla = \underbrace{\eta}_{\text{SHEAR VISCOSITY}} \cdot \Delta \vec{v} + \underbrace{\left(\zeta + \frac{2}{3} \eta \right)}_{\text{COMPRESIBILITY} = 0 \text{ IF } \text{div}(\vec{v}) = 0} \cdot \text{div}(\vec{v}) \vec{v}$$

$$\frac{\partial}{\partial t} \vec{v} + (\vec{v} \cdot \nabla) \vec{v} = -\nabla \frac{p}{\rho} + \frac{\eta}{\rho} \cdot \Delta \vec{v}$$

NAVIER-STOKES-EQUATION FOR INCOMPRESSIBLE FLUIDS

$\eta \sim$ VISCOSITY ; $\frac{\eta}{\rho} \sim$ KINETATIC VISCOSITY $\equiv \nu$

UP TO NOW THERE IS NO MICROSCOPIC THEORY OF VISCOSITY, IN EXAMPLE OF THE EFFECTIVE MOMENTUM TRANSPORT BETWEEN.

A16 FLUID MECHANICAL SIMILARITY

FLUID MECHANICAL IS CLASSICAL -> NO SCALE ASSOCIATED WITH IT (NO h, k, c, \dots)

THERE ARE CERTAIN DIFFERENT TYPES OF SOLUTIONS + REGIMES BUT CAN ONE FIND SIMILAR SOLUTIONS FOR DIFFERENT TYPES OF FLUIDS AND GEOMETRIES?

MAXWELL W - LINEARNI - VED

NAYER-STOKES - NEZINEARNI MERZE PRESALOVIAT, ALE
 UAPR, VE VERNEM RONEU-COOL AIR
 PSTOM LEPSI! ADJUSTING VISCOSITY OF AIR

DIMENSIONLESS FORM OF THE NAVIER-STOKES EQ
 (SCALING TRANSFORMATION)
 (REMOVING UNITS)

$x \rightarrow x^* = \frac{x}{L}$ WITH LENGTH SCALE L

$t \rightarrow t^* = \frac{t}{\tau}$ WITH TIME SCALE T

$u \rightarrow u^* = \frac{u}{V}$ WITH VELOCITY V

$p \rightarrow p^* = \frac{p}{P}$ WITH PRESSURE P

$g \rightarrow g^* = \frac{g}{G} = \frac{\nabla \Phi}{G}$ WITH GRAVITY G

DERIVATIVES BEHAVES LIKE THIS

$\frac{\partial}{\partial t} = \frac{\partial t^*}{\partial t} \frac{\partial}{\partial t^*} = \frac{1}{T} \frac{\partial}{\partial t^*}$

$\frac{\partial}{\partial x} = \frac{\partial x^*}{\partial x} \frac{\partial}{\partial x^*} = \frac{1}{L} \frac{\partial}{\partial x^*}$

RESCALED NAVIER-STOKES



RESCALE ALL VARIABLES

RESCALE ALL THE DERIVATIVES

$\frac{\rho V}{T} \frac{\partial \vec{u}^*}{\partial t^*} + \frac{\rho V^2}{L} (\vec{u}^* \cdot \nabla^*) \vec{u}^* = - \frac{P}{L} \nabla^* p^* - \frac{\rho G}{L} \nabla^* \Phi^* + \frac{\eta V}{L^2} \Delta^* \vec{u}^* \cdot \frac{L}{\rho V^2}$

$\frac{L}{TV} \frac{\partial \vec{u}^*}{\partial t^*} + (\vec{u}^* \cdot \nabla^*) \vec{u}^* = - \frac{P}{\rho V^2} \nabla^* p^* - \frac{G}{V^2} \nabla^* \Phi^* + \frac{\eta}{\rho L} \Delta^* \vec{u}^* \cdot \frac{L}{\rho V^2}$

STROUHAL

EULER NUMBER (EU¹)

FROUDE NUMBER

REYNOLDS

NUMBER (ST¹)

NUMBER [Fr²]

NUMBER

[Re⁻¹]

IF THESE NUMBERS ARE SUFFICIENTLY LARGE THAN THE NON LINEAR PART & WILL BE DIMINISHED.

FLUID MECHANICS IS CLASSICAL THEORY, THERE ARE NO SCALES ASSOCIATED WITH IT.

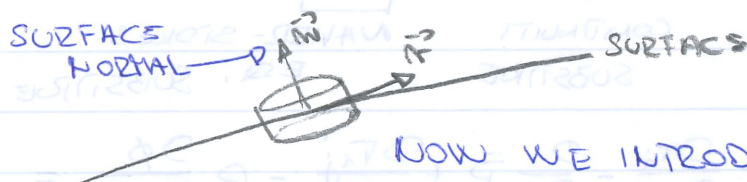
CAREFULLY ABOUT THESE, THEY BOTH ARE INVERSE.

$$St = \frac{L}{\nu} ; E_m = \frac{P}{\rho V^2} ; Fr = \frac{v}{\sqrt{g}} ; Re = \frac{\rho v L}{\mu}$$

2 FLOWS CAN BE DESCRIBED WITH THE SAME SET OF SCALING NUMBERS, THEY ARE SCALED VERSIONS OF EACH OTHER BECAUSE THEY FOLLOW THE SAME DIMENSIONLESS NAVIER-STOKES-EQUATION.

A17 BOUNDARY CONDITION

(NOT VERY RELEVANT IN ASTRO PHYSICS)



NOW WE INTRODUCE INTEGRATION OVER THIS SELECTED VOLUME:

$$\nabla \cdot \vec{n} : \int dV \operatorname{div} \vec{n} = \int dA \vec{n} = \int dA \cdot \underbrace{\vec{n} \cdot \vec{n}}_0 = 0 \text{ FOR INCOMPRESSIBLE FLUIDS}$$

$\vec{n} \cdot \vec{n} = 0$ IS BECAUSE THE FLOW CANNOT PENETRATE THE SURFACE (NO FLOW IN OR OUT OF THE SURFACE).

\vec{n}_n DEPEND ON A VISCOSITY $\vec{n} = 0$ (THERE WILL BE SHEAR FORCES AND VISCOS FORCES)
 NO-VISCOSITY, THEN \vec{n}_n IS UNCONSTRAINED

THIS APPLY FOR THE SURFACE, INSIDE THE FLOW THE \vec{n}_n IS AS WELL UNCONSTRAINED.

A18 CONSERVATION LAWS

CONSERVATION OF MASS (CONTINUITY EQUATION)

$$\frac{\partial}{\partial t} \rho + \text{div}(\rho \vec{v}) = 0$$

IN COMPONENTS: $\frac{\partial}{\partial t} \rho + \frac{\partial}{\partial x_i} (\rho v_i) = 0$

CONSERVATION OF MOMENTUM (NAVIER-STOKES EQUATION)

TIME DERIVATIVE OF MOMENTUM DENSITY:

$$\frac{\partial}{\partial t} (\rho \vec{v}) = \frac{\partial \rho}{\partial t} \vec{v} + \rho \frac{\partial \vec{v}}{\partial t}$$

IN COMPONENTS: $\frac{\partial}{\partial t} (\rho v_i) = \frac{\partial \rho}{\partial t} v_i + \rho \frac{\partial v_i}{\partial t} =$

CONTINUITY SUBSTITUTE NAVIER-STOKES EQ. SUBSTITUTE

$$= -v_i \frac{\partial}{\partial x_j} (\rho v_j) - \rho v_j \frac{\partial v_i}{\partial x_j} - \frac{\partial \rho}{\partial x_i} v_i + \frac{\partial \tau_{ij}}{\partial x_j} - \rho \frac{\partial \phi}{\partial x_i} =$$

$$= -\frac{\partial}{\partial x_j} [\rho v_i v_j + p \delta_{ij} - \tau_{ij}] - \rho \frac{\partial \phi}{\partial x_i} = \frac{\partial}{\partial t} (\rho v_i) **$$

WE DEFINE MOMENTUM FLUX TENSOR AS $\Pi_{ij} = \rho v_i v_j + p \delta_{ij} - \tau_{ij}$

IF NOTE ON MAGIC

USE $\frac{\partial}{\partial x_i} p = \frac{\partial p}{\partial x_j} \delta_{ij} = \frac{\partial}{\partial x_j} (p \delta_{ij})$

FLUX OF THE i -TH COMPONENT INTO THE j -DIRECTION

* $\frac{\partial}{\partial t} (\rho v_i) + \frac{\partial}{\partial x_j} \Pi_{ij} = -\frac{\partial \phi}{\partial x_i}$ (IN COORDINATE FREE REPRESENT.)

$$\frac{\partial}{\partial t} (\rho \vec{v}) + \text{div} \Pi_{ij} = -\nabla \phi$$

CORRESPONDS TO A TRANSPORT EQUATION WITH $\nabla \phi$ AS A SOURCE.

• CONSERVATION OF (KINETIC) ENERGY DENSITY

LET'S IMAGINE: FLUID ELEMENT FLOWS AND CAN CHANGE ITS VELOCITY (KINETIC ENERGY), BUT AS IT MOVES ALONG ITS MASS IS NOT GOING TO CHANGE. $M = \rho dV$

$$\frac{\partial}{\partial t} \left(\frac{\rho}{2} \vec{v}^2 \right) = \rho \vec{v} \cdot \frac{\partial \vec{v}}{\partial t}$$

IN COMPONENTS:

$$\frac{\partial}{\partial t} \left(\frac{\rho}{2} v_i v_i \right) = \rho v_i \frac{\partial v_i}{\partial t}$$

KINETIC ENERGY DENSITY $\frac{\rho}{2} \vec{v}^2$ OF A FLUID ELEMENT IN THE LAGRANGE FRAME, SIMILAR TO THE DERIVATION OF THE NAVIER STOKES EQUATION, WE ASSUME NO LOSS OF MATTER OF THE FLUID ELEMENT SO $\frac{d}{dt} \int_V dV = 0$. NOW WE SUBSTITUTE THE NAVIER-STOKES EQUATION FOR $\frac{\partial v_i}{\partial t}$.

$$\begin{aligned} \frac{\partial}{\partial t} \left(\frac{\rho}{2} \vec{v}^2 \right) &= -\rho v_i v_j \frac{\partial v_i}{\partial x_j} - v_i \frac{\partial}{\partial x_i} p + v_i \frac{\partial}{\partial x_j} \tau'_{ij} - \rho v_i \frac{\partial \phi}{\partial x_i} \\ &= -v_i \frac{\partial}{\partial x_i} \left(\frac{\rho}{2} \vec{v}^2 \right) - v_i \frac{\partial}{\partial x_i} p + \frac{\partial}{\partial x_i} (v_i \tau'_{ij}) - \tau'_{ij} \frac{\partial v_i}{\partial x_j} - \rho v_i \frac{\partial \phi}{\partial x_i} \end{aligned}$$

$$\frac{\partial}{\partial t} (\rho \vec{v}^2) = -\vec{v} \cdot \nabla \cdot \left(\frac{\rho \vec{v}^2}{2} + p \right) + \nabla \cdot (\tau' \cdot \vec{v}) - \tau'_{ij} \frac{\partial v_i}{\partial x_j} - \rho \vec{v} \cdot \nabla \phi$$

IDENTITY (PRODUCT RATE)

$$\nabla \cdot [\vec{v} \cdot \left(\frac{\rho}{2} \vec{v}^2 + p \right)] = \left(\frac{\rho}{2} \vec{v}^2 + p \right) \text{div} \vec{v} + \vec{v} \cdot \nabla \left(\frac{\rho \vec{v}^2}{2} + p \right)$$

+ INCOMPRESSIBILITY $\text{div} \vec{v} = 0$ (MAKES SECOND TERM DISAPPEAR)

$$\frac{\partial}{\partial t} \left(\frac{\rho}{2} \vec{v}^2 \right) = -\nabla \cdot [\vec{v} \cdot \left(\frac{\rho \vec{v}^2}{2} + p \right) - \tau' \cdot \vec{v}] - \rho \vec{v} \cdot \nabla \phi - \tau'_{ij} \frac{\partial v_i}{\partial x_j}$$

APPLY AN INTEGRATION OVER THE VOLUME OF A FLUID ELEMENT + GAUSS THEOREM

$$\frac{d}{dt} \int_V dV \frac{\rho}{2} \vec{v}^2 = - \int_{\partial V} d\vec{A} \cdot \vec{n} \cdot \frac{\rho \vec{v}^2}{2} \quad \text{FLUX OF KIN ENERGY THROUGH SURFACE}$$

$$- \int_V dV \nabla \cdot \vec{n} \cdot p \quad \text{WORK DONE BY PRESSURE}$$

$$+ \int_{\partial V} d\vec{A} \cdot (\vec{\tau} \cdot \vec{n}) \quad \text{WORK DONE BY SURFACE FORCES}$$

$$- \int_V dV \cdot \rho \cdot \nabla \cdot \phi \cdot \vec{n} \quad \text{WORK DONE BY EXTERNAL FORCES (LIKE GRAVITY)}$$

$$- \int_V dV \tau_{ij} \frac{\partial v_i}{\partial x_j} \quad \text{DISSIPATION OF ENERGY BY FRICTION}$$

↳ VELOCITY TENSOR E_{ij}

PLEASE BE CAREFUL VISCOSITY APPEARS TWICE

- SURFACE FORCES ACTING ON A FLUID ELEMENT $\int d\vec{A} (\vec{\tau} \cdot \vec{n})$

- DISSIPATION INSIDE THE VOLUME \rightarrow GENERATION OF HEAT $\int_V dV \tau_{ij} \frac{\partial v_i}{\partial x_j} = \int_V dV \text{tr}(\vec{\tau} \cdot \vec{E})$ WITH SYM. VELOCITY TENSOR

$$\frac{\partial E}{\partial t} = - \frac{\mu}{2} \cdot \text{tr}(e^2) \quad \text{WITH THERMAL ENERGY DENSITY}$$

119 BERNOULLI EQUATION

- CONSERVATION OF KINETIC ENERGY FOR

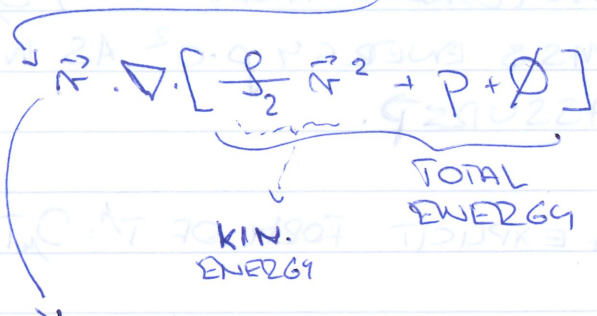
① IDEAL FLUIDS (VISCOSITY = 0 (INVISCID)) $\mu = \eta = 0$

② INCOMPRESSIBLE FLUIDS ($\text{div} \vec{v} = 0$) STATIONARY FLOWS

$$\frac{\partial E}{\partial t} = 0$$

$$\nabla \cdot \left[\vec{n} \left(\frac{\rho}{2} \vec{n}^2 + p \right) \right] + \rho \vec{n} \cdot \nabla \phi$$

THESE ARE THE RELEVANT TERMS



WITH TOTAL ENERGY $E = \frac{\rho}{2} \vec{n}^2 + p + \phi$

$\vec{n} \cdot \nabla E$ SUBSTITUTE INTO THE ENERGY CONSERVATION EQUATION

$$\frac{dE}{dt} = \vec{n} \cdot \nabla E$$

$\vec{n} \cdot \nabla \sim$ RATE OF CHANGE OF A QUANTITY WHEN DISPLACED ALONG A STREAM LINE, WHICH IS PARALLEL TO \vec{n} .

A20 RELATION TO RELATIVISTIC FLUID MECHANICS

LET'S CONSIDER DENSITY ρ

- ① LENGTH CONTRACTION - SAME DATA IN SMALLER VOLUME
- ② INCREASE IN MASS

ρ CAN'T BE SCALED, BUT MUST BE PART OF THE RELATIVISTIC ENERGY - MOMENTUM TENSOR $T^{00} \sim \rho \cdot c^2$

$$T^{\mu\nu} = \begin{pmatrix} \rho & p & 0 \\ 0 & p & p \end{pmatrix} \xrightarrow{\text{LORENTZ TRANSF}} \xrightarrow{\text{BOOST}} T^{\mu\nu} \rightarrow \Lambda^\mu_\rho \Lambda^\nu_\sigma T^{\rho\sigma}$$

$$T^{\mu\nu} = \left(\rho + \frac{p}{c^2} \right) u^\mu u^\nu - p \cdot \eta^{\mu\nu} \leftarrow T^{\mu\nu} \begin{pmatrix} \rho & p & 0 \\ 0 & p & p \end{pmatrix}$$

CONSERVATION LAW THE ENERGY - MOMENTUM TENSOR OBEYS THE POYNTING LAW (CONSERVATION LAW), $\partial_\mu T^{\mu\nu} = 0$ AND CONTAINS THE REST MASS ENERGY $\rho \cdot c^2$ AS WELL AS THE ISOTROPIC PRESSURE P .

COMPUTE DIVERGENCE WITH EXPLICIT FORM OF $T^{\mu\nu}$: $\partial_\mu T^{\mu\nu} = 0$

$$\partial_\mu \left(\rho + \frac{P}{c^2} \right) \cdot u^\mu \cdot u^\nu + \left(\rho + \frac{P}{c^2} \right) \cdot \left[\partial_\mu u^\mu \cdot u^\nu + u^\mu \partial_\mu u^\nu \right] - \partial_\mu P \zeta^{\mu\nu} = 0$$

TRICKS:

$$u^\nu u_\nu = c^2 \text{ FOR THE 4-VELOCITY: } \partial_\mu u^\nu \cdot u_\nu + u^\nu \partial_\mu u_\nu = 0$$

→ EACH TERM MUST BE ZERO INDIVIDUALLY:

$$u_\nu (\partial_\mu T^{\mu\nu}) = 0 ; \partial_\mu (\rho u^\mu) + \frac{P}{c^2} \partial_\mu u^\mu = 0 \text{ (CONTINUITY)}$$

RESUBSTITUTE:

$$\left(\rho + \frac{P}{c^2} \right) \underbrace{\partial_\mu u^\nu \cdot u_\nu}_{\downarrow \frac{\partial}{\partial t} \vec{v}^2 + (\vec{v} \cdot \nabla) \vec{v}^2} + \underbrace{\left(\frac{u^\mu u^\nu}{c^2} - \zeta^{\mu\nu} \right)}_{\downarrow \text{PROJECTION OF}} \underbrace{\partial_\mu P}_{\text{PERPENDICULAR TO } u^\mu} = 0$$

RELATIVISTIC EULER EQUAT.

- NON-RELATIVISTIC LIMIT: $\mu \approx 1$, $u^\mu (c; \vec{v})$; $P \ll \rho c^2$
- ↑ THREE VELOCITY
↑ FOUR VELOCITY

CONTINUITY:

$$\partial_\mu (\rho u^\mu) = 0 \text{ IF } P \ll \rho c^2 = \frac{\partial}{\partial t} \rho + \text{div}(\rho \vec{v}) = 0$$

$$\rho \partial_\mu u^\nu \cdot u_\nu + \left(\frac{u^\mu u^\nu}{c^2} - \zeta^{\mu\nu} \right) \partial_\mu P = 0 \text{ IF } P \ll \rho c^2$$

E-COMPONENT: IDENTICALLY FULFILLED $\partial_\mu u^0 = 0$

X-COMPONENTS: $\rho \partial_\mu u^i \cdot u^i = -\partial^i_j P$

$$\rho \cdot \left(\frac{\partial \vec{v}^2}{\partial t} + (\vec{v} \cdot \nabla) \vec{v}^2 \right) = -\nabla P$$

• CLASSICAL FLUID MECHANICS IS THE NORMAL LIMIT OF ENERGY - MOMENTUM CONSERVATION $\partial_\mu T^{\mu\nu} = 0$ PROJECTED \parallel AND \perp TO THE DIRECTION OF u^μ

• RELATIVISTIC FLUID MECHANICS INCLUDING GRAVITY, PHYSICS IN GRAVITY REQUIRES THE DEFINITION OF A CONNECTION Γ AND COVARIANT DERIVATIVE ∇_μ WITH THE PROPERTY $\nabla_\mu g^\mu = 0$ g^μ IS THE METRIC, REPLACING THE η^μ OF SPECIAL RELATIVITY.

$\partial_\beta v^\mu = \nabla_\beta v^\mu = \partial_\beta v^\mu - \Gamma_{\lambda\beta}^\mu v^\lambda$ COVARIANT DERIVATIVE

$\nabla_\mu T^{\mu\nu} = \partial_\mu T^{\mu\nu} - \Gamma_{\lambda\mu}^\mu T^{\lambda\nu} - \Gamma_{\lambda\beta}^\nu T^{\mu\lambda}$ FOR THE COVARIANT ENERGY CONSERVATION

IN A FREELY FALLING FRAME ONE REVEALS LOCALLY SPECIAL RELATIVITY.

$g \rightarrow \eta; \nabla_\mu \rightarrow \partial_\mu$ (AND $\partial g = 0$! IMPLIES $\Gamma = 0$)

$\rightarrow \nabla_\mu T^{\mu\nu} = 0$ WITH GRAVITY $\leftrightarrow \partial_\mu T^{\mu\nu}$ WITHOUT GRAVITY

WEAK FIELDS $ds^2 = g_{\mu\nu} dx^\mu dx^\nu = (1 + \frac{2\phi}{c^2}) c^2 dt^2 - (1 - \frac{2\phi}{c^2}) dx^2$
 IMPLY $\Gamma \sim \nabla\phi$, WHICH ADDS A $-\nabla\phi$ TO THE BULETS EQUATION

METRIC $g_{\mu\nu} = \begin{pmatrix} 1 + \frac{2\phi}{c^2} & \\ & -(1 - \frac{2\phi}{c^2}) \end{pmatrix}; \Gamma \sim \nabla\phi$

$\frac{\partial}{\partial t} \vec{n} + (\vec{n} \cdot \nabla) \vec{n} = -\frac{\nabla p}{\rho} - \nabla\phi$

A20 POTENTIAL FLOWS

- VELOCITY POTENTIAL φ ; VELOCITY FIELD $\vec{v} = \nabla\varphi$ FROM ELECTROSTATIC
- UNDER COMPRESSIBILITY CONDITION $\text{div}\vec{v} = \nabla\vec{v} = \nabla\nabla\varphi = \Delta\varphi = 0$
- $\text{rot}\vec{v} = 0$

$$\vec{v} \cdot \vec{n} = 0 = \frac{\partial\varphi}{\partial n} = \nabla\varphi \cdot \vec{n} \quad \text{ON A BOUNDARY (WALL)}$$

- UNIQUENESS OF THE VELOCITY POTENTIAL - GREEN THEOREM

2 SCALE FIELDS $\varphi; \psi \rightarrow$ VECTOR $\vec{M} = \varphi \nabla\psi$

$$\text{div}\vec{M} = \varphi \Delta\psi + \nabla\varphi \nabla\psi$$

GAUSS THEOREM

$$\int_V dV \text{div}\vec{M} = \int_{\partial V} d\vec{A} \cdot \vec{M} = \int_{\partial V} dA \cdot \vec{n} \cdot \vec{M}$$

$$\int_V dV \text{div}\vec{M} = \int_V dV [\varphi \Delta\psi + \nabla\varphi \nabla\psi] = \int_{\partial V} dA \cdot \underbrace{\varphi \nabla\psi \cdot \vec{n}}_{= \frac{\partial\psi}{\partial n}} =$$

GREEN THEOREM

$$= \int_{\partial V} dA \cdot \varphi \cdot \frac{\partial\psi}{\partial n}$$

NOW WE CAN INTERCHANGE φ AND ψ AND SUBTRACT FOR BOUNDARY PROBLEM*

$$\int_V dV [\varphi \Delta\psi - \psi \Delta\varphi] = \int_{\partial V} dA [\varphi \frac{\partial\psi}{\partial n} - \psi \frac{\partial\varphi}{\partial n}]$$

NOW LET'S ASSUME THAT THERE ARE TWO VALID SOLUTIONS φ AND ψ ; THIS IS A CONSEQUENCE OF INCOMPRESSIBILITY.

$$M = \varphi - \psi; \quad \Delta M = \Delta\varphi - \Delta\psi = 0 \quad (\text{INCOMPRESSIBILITY})$$

$$\int_V dV [\underbrace{M \Delta M}_0 - (\nabla M)^2] = \int_{\partial V} dA \cdot \underbrace{M \frac{\partial M}{\partial n}}_0 = 0$$

BECAUSE OF THE BOUNDARY: $\frac{\partial M}{\partial n} = 0$ NO FLOW ACROSS BOUNDARY
 M - NORMAL OF THE SURFACE

$\nabla M = 0 \Rightarrow \psi = \varphi + \text{CONST.}$ AND THE CONSTANT VANISHES IN DIFFERENTIAL VELOCITY FIELD $\nabla\psi = \nabla\varphi = \vec{v}$

• STREAM FUNCTION ψ (NO RELATION TO ϕ FROM PREVIOUS PARAGRAPH)

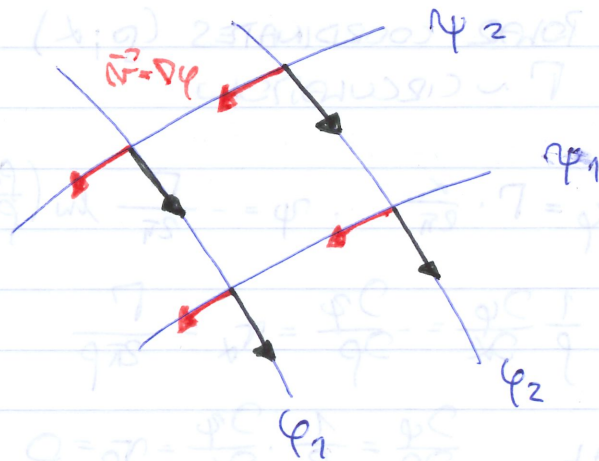
HOW CAN WE DESCRIBE FLOWS JUST USING THESE ψ & ϕ POTENTIAL WITHOUT ACTUAL VELOCITY, WELL HOW ψ WORK.

IF $\text{div } \vec{v} = 0 \rightarrow$ THEN THERE MUST BE A POTENTIAL $\vec{A} = (0, 0, \psi)$ WITH $\vec{v} = \text{rot } \vec{A}$ $v_x = \frac{\partial \psi}{\partial y}$; $v_y = -\frac{\partial \psi}{\partial x}$

$\text{div } \vec{v} = 0$ IS IDENTICALLY FULLFIELD FOR INCOMPRESSIBLE FLUIDS BECAUSE $\text{div rot } \vec{A} = 0 = \text{div } \vec{v}$.

STREAM LINES, CURVES OF CONSTANT $\psi \rightarrow \vec{v} \perp \nabla \psi$
 $\vec{v} \cdot \nabla \psi = v_x \frac{\partial \psi}{\partial x} + v_y \frac{\partial \psi}{\partial y} = \frac{\partial \psi}{\partial y} \frac{\partial \psi}{\partial x} + \frac{\partial \psi}{\partial x} \frac{\partial \psi}{\partial y} = 0$ (PERPENDICULAR)

$\rightarrow \vec{v} \cdot \nabla \psi = 0 ; \vec{v} \perp \nabla \psi ; \nabla \psi \perp \nabla \phi$

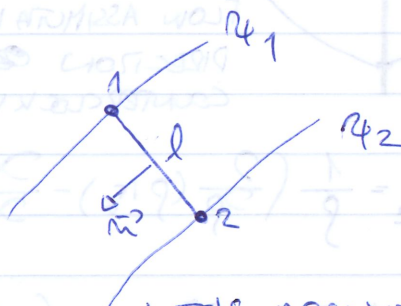


FLOW RATE $Q \sim \Delta \psi = \psi_2 - \psi_1$

~~FLUX~~ Q - FLUX FROM SURFACE

$Q = \int_1^2 \vec{v} \cdot \vec{m} dl$ $dl = (dx, dy, 0)$ $\vec{m} dl = (dy, dx, 0)$

$Q = \int_1^2 \vec{v} \cdot \vec{m} dl = \int_1^2 (\frac{\partial \psi}{\partial y} dy + \frac{\partial \psi}{\partial x} dx) = \int_1^2 d\psi = \psi_2 - \psi_1 = \Delta \psi$



LET'S ASSUME THAT LIQUID HAS A CONST. DENSITY.

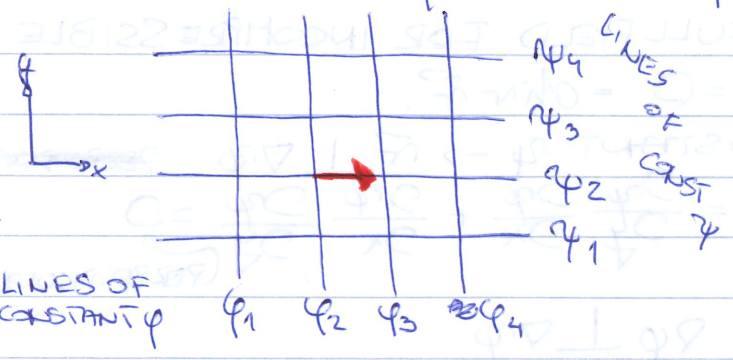
FLUX THROUGH THE SURFACE BETWEEN 2 LINES OF CONSTANT $\psi = \Delta\psi$.

EXAMPLES OF VELOCITY POTENTIALS + STREAM FUNCTIONS

1) UNIFORM PARALLEL FLOW $\vec{v} = \begin{pmatrix} u \\ 0 \end{pmatrix} = \begin{pmatrix} u_x \\ 0 \end{pmatrix}$ IN x-DIRECTION

$$\frac{\partial \phi}{\partial x} = \frac{\partial \psi}{\partial y} = v_x = u \quad \frac{\partial \psi}{\partial y} = -\frac{\partial \phi}{\partial x} = v_y = 0$$

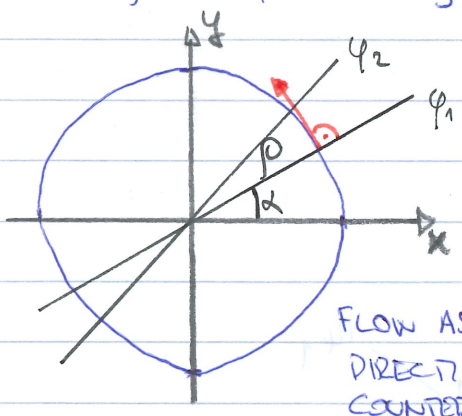
SOLVED BY $\phi = ux$ AND $\psi = u \cdot y$



IF PRIMITIVES ONE TO THE SUPERPOSITION YOU CAN ASSEMBLE MANY REALISTIC FLOWS

2) VORTEX FLOW \rightarrow POLAR COORDINATES $(\rho; \alpha)$

$$v_\rho = 0; \quad v_\alpha = \frac{\Gamma}{2\pi\rho} \quad \Gamma \sim \text{CIRCULATION}$$



$$\phi = \Gamma \cdot \frac{\alpha}{2\pi} \quad ; \quad \psi = -\frac{\Gamma}{2\pi} \ln\left(\frac{\rho}{\rho_0}\right)$$

$$\frac{1}{\rho} \frac{\partial \phi}{\partial \alpha} = -\frac{\partial \psi}{\partial \rho} = v_\alpha = \frac{\Gamma}{2\pi\rho}$$

$$\frac{\partial \phi}{\partial \rho} = \frac{1}{\rho} \cdot \frac{\partial \phi}{\partial \alpha} = v_\rho = 0$$

$$(\text{rot } \vec{v})_z = \frac{1}{\rho} \left(\frac{\partial}{\partial \rho} (\rho v_\alpha) - \frac{\partial v_\rho}{\partial \alpha} \right) = \frac{1}{\rho} \left(v_\alpha + \rho \frac{\partial v_\alpha}{\partial \rho} - \frac{\partial v_\rho}{\partial \alpha} \right) = 0$$

if $\rho \neq 0$ AT THE ORIGIN $(\text{rot } \vec{v})_z = \frac{\Gamma}{\rho}$ (BY APPLYING STOKES-THEOREM)

3) SOURCES AND SINKS - INTERCHANGED QUANTITIES COMPARED TO VORTEX

FLOW $\sigma_p = \frac{Q}{2\pi\rho}$; $\sigma_k = 0$; $Q \sim$ SOURCE STRENGTH

$$\varphi = \frac{Q}{2\pi} \cdot \ln\left(\frac{\rho}{\rho_0}\right) ; \psi = Q \cdot \frac{\alpha}{2\pi}$$

$$\frac{1}{\rho} \frac{\partial \varphi}{\partial \alpha} = \frac{\partial \psi}{\partial \rho} = 0$$

$$\frac{\partial \varphi}{\partial \rho} = \frac{1}{\rho} \frac{\partial \varphi}{\partial \ln \rho} = \frac{Q}{2\pi\rho} \quad \text{RADIAL FLOW}$$

$$\text{div } \vec{\pi} = \frac{1}{\rho} \left(\frac{\partial(\rho \sigma_p)}{\partial \rho} + \frac{\partial \pi_k}{\partial \alpha} \right) = \frac{1}{\rho} \left(\sigma_k + \rho \cdot \frac{\partial \sigma_k}{\partial \rho} + \frac{\partial \rho \sigma_p}{\partial \alpha} \right) = 0$$

If $\rho \neq 0$ AT THE ORIGIN $\text{div } \vec{\pi} = Q$ (BY APPLYING GAUSS-THEOR.)

FUNCTIONS

A21 CONFORMAL MAPPINGS - COMPLEX DIFFERENTIABILITY

COMPLEX DIFFERENTIABILITY HAS 4 ASPECTS:

- ① COMPLEX DERIVATIVE $\frac{dg}{dz}$ EXISTS AND IS UNIQUE
- ② ANALYTIC. CAUCHY-RIEMANN DIFFERENTIAL EQUATIONS APPLY
- ③ REGULAR. LOOP INTEGRALS $\oint d\{g(\zeta)\} = 0$
- ④ HOLOMORPHIC $g(z) = \oint \frac{d\zeta}{2\pi i} \cdot \frac{g(\zeta)}{\zeta - z}$
VALUES ~~at~~ A LOOP DEFINE EVERY VALUE IN THE INTERIOR

• HOLOMORPHISM

LET'S WRITE $g(\zeta) = g(z) + g(\zeta) - g(z)$

$$g(z) \stackrel{?}{=} \frac{1}{2\pi i} \oint d\zeta \frac{g(\zeta)}{\zeta - z} = \frac{1}{2\pi i} \oint d\zeta \frac{g(z)}{\zeta - z} + \frac{1}{2\pi i} \cdot \oint d\zeta \frac{g(\zeta) - g(z)}{\zeta - z} =$$

$$= \frac{g(z)}{2\pi i} \underbrace{\oint d\zeta \frac{1}{\zeta - z}}_{2\pi i} + \frac{1}{2\pi i} \underbrace{\oint d\zeta \frac{g(\zeta) - g(z)}{\zeta - z}}_{*} = g(z) = \text{---}$$

$$\leq \int d\xi \left(\frac{|g(\xi) - g(z)|}{\xi - z} \right) \leq \max |g(\xi) - g(z)| \cdot \int d\xi \frac{1}{\xi - z}$$

= $g(z)$ IF THE SECOND TERM IS $\leq \epsilon$ BY SELECTION OF THE INTEGRATION BOUNDARY.

THIS CAN BE EXTENDED BY TO n -FOLD DIFFERENTIATION

$$\frac{d^n}{dz^n} g(z) = \frac{n!}{2\pi i} \int d\xi \frac{g(\xi)}{(\xi - z)^{n+1}}$$

HOLomorphic FUNCTIONS HAVE HOLomorphic DERIVATIVES.

• COMPLEX DIFFERENTIABILITY

$$\left. \frac{dg}{dz} \right|_z = \lim_{z \rightarrow \xi} \frac{g(z) - g(\xi)}{z - \xi} \quad \text{EXISTS AND IS UNIQUE (THE DETAILS OF } z - \xi \text{ DON'T MATTER)}$$

• CAUCHY - RIEMANN DIFFERENTIAL EQUATIONS

$$\frac{dg}{dz} = \lim_{\Delta z \rightarrow 0} \frac{g(z + \Delta z) - g(z)}{\Delta z} \quad \rightarrow \text{SHOULD BE INDEPENDENT ON } \Delta z$$

$$\frac{\partial g}{\partial x} = \lim_{\Delta x \rightarrow 0} \frac{g(x + iy + \Delta x) - g(x + iy)}{\Delta x}$$

$$\frac{1}{i} \frac{\partial g}{\partial y} = \lim_{\Delta y \rightarrow 0} \frac{g(x + iy + i\Delta y) - g(x + iy)}{i\Delta y}$$

SHOULD BE EQUAL TO $\frac{dg}{dz}$

DECOMPOSE INTO REAL AND IMAGINARY PARTS $g(z) = u(x+iy) + iv(x+iy)$ BUT IF THE FUNCTION IS COMPLEX DIFFERENTIABLE, THE WAY IN WHICH THE DERIVATIVE IS COMPOSED, SHOULD NOT MATTER.

$$\frac{\partial u}{\partial x} = \frac{\partial v}{\partial y} \quad \text{AND} \quad \frac{\partial u}{\partial y} = -\frac{\partial v}{\partial x}$$

THAN $\frac{\partial g}{\partial x} = \frac{\partial(u+iv)}{\partial x} = \frac{\partial u}{\partial x} + i \frac{\partial v}{\partial x}$

$$\frac{1}{i} \frac{\partial g}{\partial y} = \frac{\partial(u+iv)}{\partial y} = \frac{1}{i} \frac{\partial u}{\partial y} + \frac{\partial v}{\partial y}, \quad \text{WITH } \frac{1}{i} = -i$$

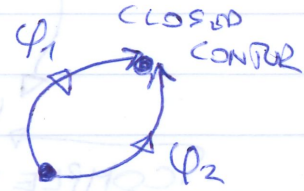
• REGULARITY CONDITION

CAUCHY THEOREM - CONTOUR INTEGRALS OF COMPLEX DIFFERENTIAL FUNCTIONS DON'T DEPEND ON INTEGRATION PATH.

LOOP INTEGRALS - CLOSED CONTOUR, IF $z(a) = z(b)$

~~CAUCHY THEOREM~~

$$\int_{\gamma_1} dz g(z) + \int_{\gamma_2} dz g(z) = \oint dz g(z) = 0$$



REWRITE CAUCHY THEOREM

$$\oint dz g(z) = \oint (dx + i dy) (u + iv) = \oint (u dx - v dy) + i \oint (v dx + u dy) = \oint \vec{R} \cdot \vec{R} + i \oint \vec{I} \cdot \vec{I}$$

REINTERPRET REAL AND IMAGINARY PATHS AS SCALED PRODUCTS.

$\vec{R} = (u, -v, 0); \vec{I} = (v, u, 0)$ WITH $d\vec{R} = (dx, dy, 0)$

THIS LEADS TO

APPLY STOKES THEOREM - TRANSFORM LOOP INTEGRAL INTO SURFACE INTEGRAL

$$\oint_{\gamma \rightarrow \partial A} d\vec{R} \cdot \vec{R} = \int_A d\vec{A} \text{rot} \vec{R} = \int_A dx dy \left(\frac{\partial v}{\partial x} + \frac{\partial u}{\partial y} \right) = 0$$

CR-1 \approx

$$\oint_{\partial A} d\vec{r} \cdot \vec{I} = \int_A dA \operatorname{rot} \vec{I} = \int dx dy \left(\frac{\partial u}{\partial x} - \frac{\partial v}{\partial y} \right) = 0$$

CR-2

• COMPLEX VELOCITY POTENTIAL

$$v_x = \frac{\partial \varphi}{\partial x} = \frac{\partial \psi}{\partial y} \quad \text{AND} \quad v_y = \frac{\partial \varphi}{\partial y} = -\frac{\partial \psi}{\partial x}$$

THIS CORRESPONDS TO THE CAUCHY-RIEMANN DIFFERENTIAL EQUAT.

if:

$$g(z) = \varphi(z) + i\psi(z) ; \quad z = x + iy \sim \text{COMPLEX VELOCITY POTENTIAL}$$

$$w(z) = \frac{dg}{dz} = \left\{ \frac{\partial g}{\partial x} = \frac{\partial \varphi}{\partial x} + i \frac{\partial \psi}{\partial x} = v_x - i v_y \right.$$

COMPLEX VELOCITY

$$\left. \frac{1}{i} \frac{\partial g}{\partial y} = \frac{\partial \varphi}{\partial y} - i \frac{\partial \psi}{\partial y} = v_x - i v_y \right\}$$

• PHYSICAL INTERPRETATION OF LOOP INTEGRALS

$$C(z) = \oint_{\gamma} dz w(z) = \oint_{\gamma} (dx + i dy)(v_x + i v_y) =$$

$$= \oint_{\gamma} (v_x dx + v_y dy) + i \oint_{\gamma} (v_x dy - v_y dx) =$$

$$= \oint_{\gamma} d\vec{l} \cdot \vec{v} + i \oint_{\gamma} d\vec{l} \cdot \vec{n} \cdot \vec{\omega} = \Gamma + iQ$$

CIRCULATION

SOURCE

• CONFORMAL MAPPINGS

- MAPPING OF COORDINATES $(x, y) \rightarrow (u, v)$ IS CALLED CONFORMAL IF THIS MAPPING IS AN ANALYTIC FUNCTION.

CONFORMAL MAPPINGS CONSERVE THE HARMONIC PROPERTY $\Delta_{xy} g = 0$ IF $(x,y) \rightarrow (u,v)$ WITH A NEW $G(u,v)$, THEN $\Delta_{uv} G = 0$.

$$\Delta_{xy} g = 0 \xrightarrow{(x,y) \rightarrow (u,v)} \Delta_{uv} G = 0$$

PHYSICAL CONSEQUENCES, ANY SOLUTION g OF A 2D INCOMPRESSIBLE FLOW CAN BE MAPPED INTO A ~~NEW~~ MORE COMPLICATED GEOMETRY.

$$G(u,v) \rightarrow g(x,y) = G(u(x,y); v(x,y))$$

$$\frac{\partial g}{\partial x} = \frac{\partial u}{\partial x} \frac{\partial G}{\partial u} + \frac{\partial v}{\partial x} \frac{\partial G}{\partial v}$$

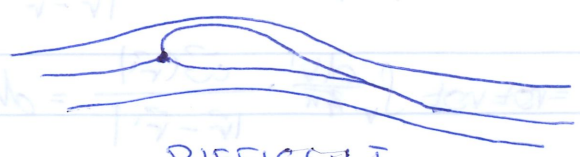
$$\frac{\partial^2 g}{\partial x^2} = \left(\frac{\partial u}{\partial x}\right)^2 \frac{\partial^2 G}{\partial u^2} + 2 \frac{\partial u}{\partial x} \frac{\partial v}{\partial x} \frac{\partial^2 G}{\partial u \partial v} + \left(\frac{\partial v}{\partial x}\right)^2 \frac{\partial^2 G}{\partial v^2} + \frac{\partial^2 u}{\partial x^2} \frac{\partial G}{\partial u} + \frac{\partial^2 v}{\partial x^2} \frac{\partial G}{\partial v}$$

$$\frac{\partial g}{\partial y} = \dots \quad \left| \quad \frac{\partial^2 g}{\partial y^2} = \dots \right.$$

$$\Delta g = \underbrace{\frac{\partial^2 g}{\partial x^2}}_0 + \underbrace{\frac{\partial^2 g}{\partial y^2}}_0 = \dots = \underbrace{\left(\frac{\partial^2 G}{\partial u^2} + \frac{\partial^2 G}{\partial v^2}\right)}_0 \cdot \underbrace{\left[\left(\frac{\partial u}{\partial x}\right)^2 + \left(\frac{\partial v}{\partial x}\right)^2\right]}_{\text{CORRELATION FACTOR}}$$

CORRELATION FACTOR

$\Delta g \propto \Delta G$ WITH A POSITION DEPENDENT FACTOR
FLOW ACROSS AIRPLANE WING



EASY: 1 SINK 1 SOURCE
1 VORTEX
1 UNIFORM FLOW

DIFFICULT
GEOMETRY

A22 VORTICITY - GENERAL PROPERTIES

- VORTICITY + CIRCULATION

VELOCITY TENSOR $\frac{\partial v_i}{\partial x_j}$

SYMMETRY PART $\epsilon_{ij} = \frac{1}{2} \left(\frac{\partial v_i}{\partial x_j} + \frac{\partial v_j}{\partial x_i} \right)$

ANISYMM. PART $\omega_{ij} = \frac{1}{2} \left(\frac{\partial v_i}{\partial x_j} - \frac{\partial v_j}{\partial x_i} \right)$

ω_{ij} IS TRACELESS + ANISYMMETRIC

$$\omega_i = \epsilon_{ijk} \omega_{jk} \rightarrow \vec{\omega} = \text{rot } \vec{v} \text{ (AXIAL VECTOR)}$$

FOCUS ON INCOMPRESSIBLE FLUIDS $\text{div } \vec{v} = 0 \rightarrow$

ANALOGY TO ELECTRODYNAMICS $\vec{B} \sim \vec{v}$; $\vec{\omega} \sim \vec{j}$

EXISTENCE OF A VECTOR POTENTIAL \sim STREAM FUNCTION ψ IN 2D

CIRCULATION

$$\Gamma = \int_{\partial A} d\vec{r} \cdot \vec{v} = \int_A d\vec{A} \cdot \text{rot } \vec{v} = \int_A d\vec{A} \cdot \vec{\omega} \quad \text{WITH STOKES THEOREM}$$

$$\rightarrow \int_{\partial A} d\vec{r} \cdot \vec{B} = \int_A d\vec{A} \cdot \text{rot } \vec{B} = \int_A d\vec{A} \cdot \vec{j} \quad \text{ENCLOSED EL. CURRENT}$$

- BIOT-SAVART LAW: RECOVER \vec{v} FROM $\vec{\omega}$

$$\vec{v}(\vec{r}) = \int \frac{dV'}{4\pi} \omega(\vec{r}') \times \frac{\vec{r} - \vec{r}'}{|\vec{r} - \vec{r}'|^3}$$

$$\hookrightarrow \nabla \frac{1}{|\vec{r} - \vec{r}'|} = -\nabla' \frac{1}{|\vec{r} - \vec{r}'|}$$

$$\text{rot } \vec{v} = \text{rot} \int \frac{dV'}{4\pi} \omega(\vec{r}') \times \nabla \frac{1}{|\vec{r} - \vec{r}'|}$$

$$= \text{rot} \int \frac{dV'}{4\pi} \cdot \nabla \times \left(\frac{\omega(\vec{r}')}{|\vec{r} - \vec{r}'|} \right) = \text{BECAUSE } \omega \text{ DOES NOT DEPEND ON } \vec{r}$$

$$= \text{rot rot} \int \frac{dV'}{4\pi} \cdot \frac{\omega(\vec{r}')}{|\vec{r} - \vec{r}'|} = \text{div} \int \frac{dV'}{4\pi} \cdot \omega(\vec{r}') \cdot \nabla \frac{1}{|\vec{r} - \vec{r}'|}$$

$$- \int \frac{dV'}{4\pi} \omega(\vec{r}') \cdot \Delta \frac{1}{|\vec{r} - \vec{r}'|} = *$$

$$= \text{div} \int_V \frac{dv'}{4\pi} \underbrace{\nabla' \vec{\omega}(\vec{r}')}_{=0} \cdot \frac{1}{|\vec{r}-\vec{r}'|} + \int_V dv' \vec{\omega}(\vec{r}') - 4\pi \int_V d\vec{r}' \delta(\vec{r}-\vec{r}') = \vec{\omega}$$

A23 VORTICITY EQUATION \rightarrow TIME EVOLUTION OF $\vec{\omega}$

STARTING POINT NAVIER-STOKES EQUATION FOR INCOMPRESSIBLE FLUIDS

$$\frac{\partial \vec{v}}{\partial t} + (\vec{v} \cdot \nabla) \vec{v} = -\frac{\nabla p}{\rho} - \nabla \phi + \mu \Delta \vec{v}$$

USE $(\vec{v} \cdot \nabla) \vec{v} = \nabla \frac{\vec{v}^2}{2} - \underbrace{\vec{v} \times \nabla \times \vec{v}}_{\vec{\omega} = \text{rot} \vec{v}}$, WITH $\vec{\omega} = \text{rot} \vec{v}$

$$\vec{v} \times \nabla \times \vec{v} = \vec{v} \times \vec{\omega} = -\vec{\omega} \times \vec{v} = -(\nabla \times \vec{v}) \times \vec{v} = -\frac{\nabla \vec{v}^2}{2} + (\vec{v} \cdot \nabla) \vec{v}$$

BUILD ROTATION OF THE NAVIER-STOKES-EQ

$$\frac{\partial \vec{v}}{\partial t} + \nabla \frac{\vec{v}^2}{2} - \vec{v} \times \vec{\omega} = -\frac{\nabla p}{\rho} - \nabla \phi + \mu \Delta \vec{v} \quad / \text{rot}$$

$$\frac{\partial \text{rot} \vec{v}}{\partial t} - \text{rot}(\vec{v} \times \vec{\omega}) = \mu \text{rot}(\Delta \vec{v})$$

- $\text{rot} \nabla(\vec{v}^2) = 0$
- $\text{rot} \nabla \phi = 0$

ROTOR OF A GRADIENT

- IFF $p = p(\rho)$, AND NOTHING ELSE

$$\nabla p = \frac{\partial p}{\partial \rho} \cdot \nabla \rho \rightarrow \nabla p \times \nabla \rho = 0!$$

$$\text{rot} \left(\frac{\nabla p}{\rho} \right) = \frac{\text{rot} \nabla p}{\rho} - \frac{\nabla p \times \nabla \rho}{\rho^2} = -\frac{\nabla p \times \nabla \rho}{\rho^2} = 0$$

"BARIONIC TERM" ONLY ACTIVE FOR EOS $p = p(\rho)$

SIMPLIFY THE TWO REMAINING TERMS

- $\text{rot}(\vec{A} \times \vec{B}) = \vec{A} \cdot \text{div} \vec{B} - \vec{B} \cdot \text{div} \vec{A} + (\vec{B} \cdot \nabla) \vec{A} - (\vec{A} \cdot \nabla) \vec{B}$

$$\rightarrow \text{rot}(\vec{v} \times \vec{\omega}) = (\vec{\omega} \cdot \nabla) \vec{v} - (\vec{v} \cdot \nabla) \vec{\omega} + \underbrace{\vec{\omega} \cdot \text{div} \vec{v}}_{=0} + \underbrace{\vec{v} \cdot \text{div} \vec{\omega}}_{=0} =$$

$$= (\vec{\omega} \cdot \nabla) \vec{v} - (\vec{v} \cdot \nabla) \vec{\omega}$$

INCOMPRESSIBLE

DIVERGENCE OF

rot

• $\text{rot rot } (\vec{r}) = \nabla \text{div } \vec{r} - \Delta \vec{r}$

$\rightarrow (\Delta \vec{r}) = \text{rot } (\nabla \cdot \text{div } \vec{r} - \text{rot rot } \vec{r}) = -\text{rot rot } \vec{r} = 0$

$= -\text{rot rot } \vec{\omega} = \Delta \vec{\omega} - \nabla \text{div } \vec{\omega} = \Delta \vec{\omega}$

$\rightarrow \text{rot } (\Delta \vec{r}) = \Delta \vec{\omega}$ if $\text{div } \vec{r} = 0$

LET'S COLLECT ALL RESULTS FOR THE VORTICITY EQUATION

$\frac{\partial}{\partial t} \vec{\omega} + (\vec{r} \cdot \nabla) \vec{\omega} = (\vec{\omega} \cdot \nabla) \vec{r} + \mu \Delta \vec{\omega}$

ADVECTIVE DERIVATIVE VISCOSITY

ADVECTION: GIVEN BY $\vec{r} = \sqrt{\frac{\rho l^3}{4\pi}} \vec{\omega} \times \nabla \frac{1}{|\vec{r} - \vec{r}'|} \sim \text{NONLOCAL}$

VISCOSITY GENERALIZES DIFFUSION: $\frac{\partial \vec{\omega}}{\partial t} - \mu \Delta \vec{\omega} = 0$

GREEN FUNCTION $G(\vec{r}, t) = \frac{1}{(2\pi t)^{3/2}} \cdot e^{-\frac{r^2}{2\mu t}}$

DIFFUSION $\propto \sqrt{t} \sim \text{RANDOM WALK}$

VORTICITY EQUATION IS NONLINEAR $(\vec{r} \cdot \nabla) \vec{\omega}$ AND $(\vec{\omega} \cdot \nabla) \vec{r}$ + NONLOCAL GENERATION ONLY DUE TO BARIONIC TERM $\rho \times \rho$

GRAVITY CAN'T DO IT: $\text{rot } \nabla \phi = 0$ (WHY DO GALAXIES ROTATE)

A24 ENSTROPY AND PALINOSTROPY

• ENSTROPY $e = \frac{\vec{\omega}^2}{2}$

$\frac{\partial}{\partial t} e = \frac{1}{2} \frac{\partial}{\partial t} \vec{\omega}^2 = \frac{1}{2} \vec{\omega} \cdot \frac{\partial \vec{\omega}}{\partial t} = \frac{\vec{\omega}}{2} \cdot ((\vec{r} \cdot \nabla) \vec{\omega}) = \sum_{ij} \omega_i \cdot \frac{\partial \omega_i}{\partial x_j} \sim \mathcal{O}(\omega^3)$

$\rightarrow \frac{\partial}{\partial t} \omega^2 \propto \omega^3$

SOLVED BY $\omega \sim \frac{1}{\frac{1}{\omega_0} - t}$ DIVERGES WHEN $t \rightarrow \frac{1}{\omega_0}$
 ω_0 - INITIAL VALUE

PALINOSTROPHY $\vec{p} = \text{rot } \vec{\omega} = \text{rot rot } \vec{n} = \text{grad div } \vec{n} - \Delta \vec{n}$
 $\vec{p} = -\Delta \vec{n}$ FOR INCOMPRESSIBLE FLUIDS
 (ONLY NEEDED FOR IMPRESSING PEOPLE)

A25 DYNAMICS OF THE CIRCULATION → KELVIN-THEOREM

(IN THE GERMAN LITERATURE, THEY ARE CALLED HELMHOLTZ-THEOREM)

$\Gamma = \int_{\partial A} d\vec{r} \times \vec{n} = \int_A dA \cdot \vec{\omega}$ STOKES-THEOREM WITH $\vec{\omega} = \text{rot } \vec{n}$

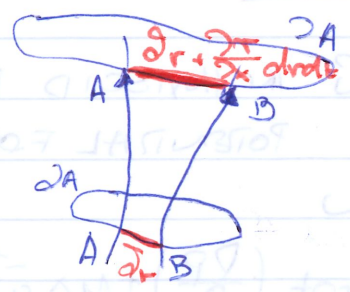
- ASSUME {
 IDEAL FLUIDS, VISCOSITY = 0
 POTENTIAL FORCE $D\phi$, NO OTHER FORCES
 CONSTANT DENSITY (OR AT MOST A FUNCTION OF p)

→ IDENTICAL TO THE ASSUMPTIONS OF THE BERNOULLI-THEOR

KELVIN-THEOREM: THE ADVECTED CIRCULATION IS CONSERVED

$\frac{d\Gamma}{dt} = \frac{d}{dt} \int_A d\vec{r} \times \vec{n} = \frac{d}{dt} \int_A dA \cdot \vec{\omega} = 0$

WITH ADVECTIVE DERIVATIVES $\frac{d}{dt} () = \frac{\partial}{\partial t} () + \vec{n} \cdot \nabla ()$



$\frac{d}{dt} \int_{\partial A} d\vec{r} \cdot \vec{n} = \int_{\partial A} \frac{d}{dt} d\vec{r} \cdot \vec{n} + \int_{\partial A} d\vec{r} \cdot \frac{d\vec{n}}{dt}$

WITH $\int_{\partial A} \frac{d}{dt} d\vec{r} \cdot \vec{n} = \int_{\partial A} \frac{d}{dt} \frac{\partial r_i}{\partial x_j} n_i = \int_{\partial A} \frac{\partial n_i}{\partial x_j} \frac{\partial r_j}{\partial x_i} n_i =$

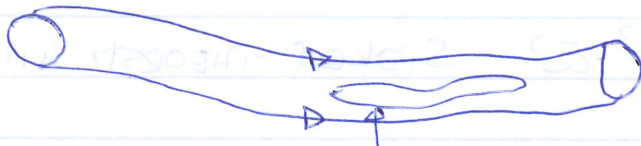
$= \int_{\partial A} \frac{\partial}{\partial x_j} \left(\frac{n_i^2}{2} \right) = \int_{\partial A} d\vec{r} \cdot \nabla \left(\frac{n^2}{2} \right)$

$$\text{AND } \int_{\partial A} \vec{v} \cdot \frac{d\vec{n}}{dt} = \int_{\partial A} \vec{v} \cdot \left(\frac{\partial \vec{n}}{\partial t} + (\vec{n} \cdot \nabla) \vec{n} \right) = \int_{\partial A} \vec{v} \cdot \left(-\frac{\nabla p}{\rho} - \nabla \phi \right)$$

LOOK AT BOTH TERMS: THEY ARE GRADIENTS OF SCALARS, SO THEIR CORRESPONDING CIRCULATION NEEDS TO BE ZERO!

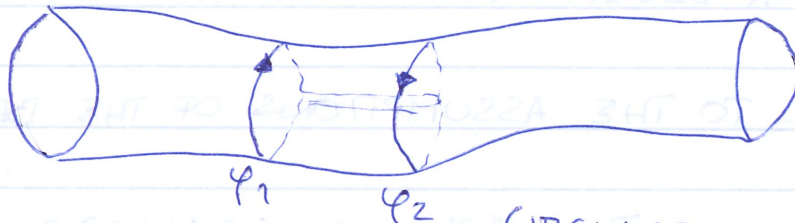
A26 CONCLUSIONS FROM THE KELVIN-THEOREM

- IF THE CIRCULATION $\Gamma = 0$, IT REMAINS ZERO $\rightarrow \vec{\omega} = 0$
- VORTICITY FOLLOWS THE FLOW BY ADVECTION



∂A ON SURFACE NO FLUX OF $\vec{\omega}$ OUT OR IN

- CIRCULATION IS CONSERVED ALONG A FLOW



CIRCULATION THROUGH φ_1 AND φ_2 ARE OPPOSITE AND EQUAL
 \rightarrow CONTRACT LOOP

A27 GENERATION OF VORTICITY

CONSERVATION OF $\Gamma = \int_A d\vec{A} \cdot \vec{\omega}$ IN INVISCID FLOWS WITH POTENTIAL FORCES

FULL VORTICITY EQUATION

$$\frac{\partial \vec{\omega}}{\partial t} + (\vec{n} \cdot \nabla) \vec{\omega} = (\vec{\omega} \cdot \nabla) \vec{\omega} - \text{rot} \left(\frac{\nabla p}{\rho} \right) + \mu \Delta \vec{\omega}$$

BAROMETRIC TERM

$\text{rot} \left(\frac{\nabla p}{\rho} \right) = -\frac{\nabla p \times \nabla p}{\rho} \neq 0$ IF $p = p(\rho, T)$, SUCH THAT PRESSURE AND DENSITY GRADIENTS ARE NOT EQUAL

NON-POTENTIAL FORCES → CORIOLIS FORCE
 ↘ LORENTZ FORCE

• CORIOLIS FORCE → $2\vec{\Omega} \times \vec{v}$ APPEARS AS A TERM IN THE NS-EQ

$$\frac{D\vec{v}}{Dt} + (\vec{v} \cdot \nabla) \vec{v} = -\frac{\nabla p}{\rho} - \nabla\phi + \mu \Delta \vec{v} - 2\vec{\Omega} \times \vec{v}$$

(SIMILARLY TO ϕ , AS AN INERTIAL (GRAVITATIONAL FORCE))

$$\text{rot}(\vec{\Omega} \times \vec{v}) = \underbrace{(\vec{v} \cdot \nabla) \vec{\Omega}}_{=0 \text{ if } \vec{\Omega} = \text{konst}} - \underbrace{(\vec{\Omega} \cdot \nabla) \vec{v}}_{=0 \text{ INCOMPRESSIBLE}} - \underbrace{\vec{\Omega} \text{div} \vec{v}}_{=0} + \underbrace{\vec{v} \cdot \text{div} \vec{\Omega}}_{=0 \text{ if } \vec{\Omega} = \text{konst}}$$

→ $\text{rot}(\vec{\Omega} \times \vec{v}) = (\vec{\Omega} \cdot \nabla) \vec{v}$

$$\frac{\partial \vec{\omega}}{\partial t} + (\vec{v} \cdot \nabla) \vec{\omega} = + \frac{\nabla p \times \nabla p}{\rho^2} + \mu \Delta \vec{\omega} - 2(\vec{\Omega} \cdot \nabla) \vec{v}$$

IMPORTANCE: QUANTITY WITH A DIMENSIONLESS NUMBER

$$\left| \frac{(\vec{v} \cdot \nabla) \vec{v}}{\vec{\Omega} \times \vec{v}} \right| \sim \frac{v^2/L}{\Omega \cdot v} = \frac{v}{\Omega L} = Ro \quad \text{ROSSBY NUMBER}$$

• LORENTZ FORCE (NON-RELATIVISTIC LIMIT)

$$\vec{f} = \frac{q}{\rho} (\vec{E} + \vec{v} \times \vec{B}), \quad \text{rot} \vec{f} = \frac{q}{\rho} \cdot \text{rot}(\vec{v} \times \vec{B}) = -\frac{q}{\rho} (\vec{B} \cdot \nabla) \vec{v}$$

IMPORTANCE

$$\left| \frac{(\vec{v} \cdot \nabla) \vec{v}}{q \cdot (\vec{B} \cdot \nabla) \vec{v}} \rho \right| = \frac{v \cdot \rho}{q \cdot B \cdot L} = \frac{1}{Re \cdot Ha} \quad \text{HARTMANN-NUMBER}$$

WE'LL GO THROUGH PLASMA PHYSICS (IN DETAIL)

A28 STOKES EQUATION - FLOW AT LOW REYNOLDS NUMBER

- VORTICITY IMPORTANT IN TURBULENT FLOWS

$(\vec{v} \cdot \nabla) \vec{v}$ DOES NOT PLAY AN IMPORTANT ROLE → LINEAR THEORY (FLUID)
 VISCOSITY DOMINATES

WHAT IS LEFT IN NAVIER-STOKES EQ:

$$-\frac{\nabla p}{\rho} - \nabla \phi + \mu \Delta \vec{n} = 0$$

UNDER CONDITION, THAT THE FLOW IS STATIC

$$\frac{\partial \vec{n}}{\partial t} = 0 \text{ HONEY ON A TOAST}$$

TWO GRADIENTS

$$\nabla p' = \nabla p + \rho \nabla \phi$$

$$\nabla p' = \mu \Delta \vec{n}$$

Absorbing $\nabla \phi$ INTO ∇p

EQUIVALENT REPRESENTATION OF STOKES EQUATION

- $\frac{\partial \sigma_{ij}'}{\partial x_i} = \mu \Delta n_j = \mu \rho \Delta n_j$ - STRESS TENSOR WHERE TRACS OF THIS TENSOR IS EQUAL TO PRESSURE p

$$\frac{\partial \sigma_{ij}'}{\partial x_j} = 0 \text{ WITH } \sigma_{ij}' = \sigma_{ij}' - p \delta_{ij}'$$

STRESS TENSOR DIVERGENCE IS EQUAL TO ZERO

- $\text{rot rot } \vec{n} = \underbrace{\nabla \text{div } \vec{n}}_{=0} - \Delta \vec{n} \rightarrow -\frac{\nabla p}{\rho} = \mu \text{rot } \vec{\omega}$

INCOMPRESSIBLE FLUIDS

LAPLACE TYPE POTENTIAL PROBLEM

- $\text{rot rot } \vec{\omega} = \underbrace{\nabla \text{div } \vec{\omega}}_0 - \Delta \vec{\omega} \rightarrow -\frac{1}{\mu} \text{rot } \nabla p = 0 = \Delta \vec{\omega}$
FOR CONSTANT p

30 SOLUTIONS TO THE STOKES EQUATION

- UNIQUENESS - ASSUME THAT THERE ARE 2 SOLUTIONS, IDENTICAL BOUNDARY CONDITIONS

$$\frac{\partial \pi_i}{\partial x_j} = \frac{\partial \pi_i'}{\partial x_j}$$

IDENTICAL VELOCITY TENSORS?

EQUALITY OF \vec{n} THEN FOLLOWS FROM INTEGRATION WITH FIXED BOUNDARY CONDITION

$$\int_V dV \left(\frac{\partial \pi_i}{\partial x_j} - \frac{\partial \pi_i'}{\partial x_j} \right)^2 = \int_V dV \frac{\partial}{\partial x_j} \left[(\pi_i - \pi_i') \cdot \left(\frac{\partial \pi_i}{\partial x_j} - \frac{\partial \pi_i'}{\partial x_j} \right) \right] - \int_V dV (\pi_i - \pi_i') (\Delta \pi_i - \Delta \pi_i')$$

① = 0

②

① SURFACE INTEGRAL = 0 (GAUSS THEOREM) DUE TO IDENTICAL BOUNDARY CONDITIONS

$$\begin{aligned} \textcircled{2} &= -\frac{1}{2} \int_V dV (\pi_i - \pi_i') \cdot \left(\frac{\partial p}{\partial x_i} - \frac{\partial p'}{\partial x_i} \right) = \\ &= -\frac{1}{2} \int_V dV \frac{\partial}{\partial x_i} (\pi_i - \pi_i') (p - p') + \frac{1}{2} \int_V dV \\ &\quad \underbrace{\left(\frac{\partial \pi_i}{\partial x_i} - \frac{\partial \pi_i'}{\partial x_i} \right) \cdot (p_i - p_i')}_{\substack{\text{D=GAUSS TH.} \\ =0 \text{ INCOMPRESSIBLE}}} = 0 \end{aligned}$$

→ VELOCITY TENSORS ARE EQUAL! IF $\text{div } \vec{v} = 0$

→ (IDENTITY OF \vec{v}) FOLLOWS FROM INTEGRATION

AD1 ENERGY DISSIPATION ^{IN} ~~A~~ STOKES FLOW

$$E = \int_V dV \epsilon'_{ij} \frac{\partial v_i}{\partial x_j} = \frac{\eta}{2} \int_V dV \left(\frac{\partial v_i}{\partial x_j} + \frac{\partial v_j}{\partial x_i} \right)^2 \Rightarrow = 2\eta \int_V dV e^2$$

VELOCITY TENSOR

ENERGY DISSIPATION → MINIMAL FOR STOKES FLOW $\delta P = \delta \sigma \vec{v}$

NOW ASSUME THAT THERE IS A STRAIN TENSOR ϵ' , WHICH DISSIPATES MORE ENERGY THAN ϵ

WHERE ϵ IS THE SOLUTION TO THE STOKES EQUATION & ϵ & ϵ' IS BOTH INCOMPRESSIBLE!

$$2\eta \int_V dV e'^2 = \underbrace{2\eta \int_V dV e^2}_{\textcircled{1}} + \underbrace{2\eta \int_V dV (\epsilon' - \epsilon)^2}_{\textcircled{2}}$$

$$+4\eta \cdot \int_V dV (e' - e) e \quad \textcircled{3}$$

INTEGRAL ② IS NECESSARY POSITIVE $\rightarrow (\)^2 \rightarrow +$
 IF WE CAN SHOW THAT ③ IS ZERO, THEN THE DISSIPATION WITH e' IS MORE EFFECTIVE THAN WITH e .
 OUR GOAL WILL BE TO SHOW THAT $\int_V dV e'^2 > \int_V dV e^2$ BY SHOWING THAT ③ $\equiv 0$

$$\int_V dV (e' - e) e = \frac{1}{2} \left[\int_V dV \left(\frac{\partial r_i'}{\partial x_j} - \frac{\partial r_i}{\partial x_j} \right) \frac{\partial r_i}{\partial x_j} + \int_V dV \left(\frac{\partial r_i'}{\partial x_j} - \frac{\partial r_i}{\partial x_j} \right) \frac{\partial r_j}{\partial x_i} \right] \quad \textcircled{A} \quad \textcircled{B}$$

$$\textcircled{A} = \int_V dV \frac{\partial}{\partial x_j} \left[(r_i' - r_i) \frac{\partial r_i}{\partial x_j} \right] - \int_V dV (r_i' - r_i) \frac{\partial^2 r_i}{\partial x_j^2} =$$

SURFACE INTEGRAL = 0

$$= \frac{1}{\eta} \int_V dV (r_i' - r_i) \cdot \frac{\partial p}{\partial x_i} = \frac{1}{\eta} \left(\int_V dV \frac{\partial}{\partial x_i} \left[(r_i' - r_i) p \right] - \int_V dV \frac{\partial}{\partial x_i} (r_i' - r_i) \cdot p \right)$$

$\int_V dV \frac{\partial}{\partial x_i} \left[(r_i' - r_i) p \right] = 0$ GAUSS. TH.

$$- \int_V dV \frac{\partial}{\partial x_i} (r_i' - r_i) \cdot p = 0$$

$\int_V dV \frac{\partial}{\partial x_i} (r_i' - r_i) \cdot p = 0$ INCOMPRESSIBLE

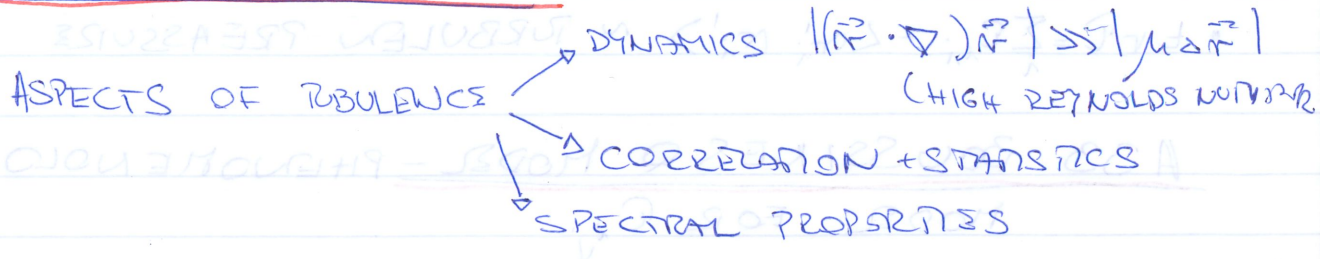
$$\textcircled{B} = \int_V dV \left(\frac{\partial r_i'}{\partial x_j} - \frac{\partial r_i}{\partial x_j} \right) \cdot \frac{\partial r_j}{\partial x_i} = \int_V dV \frac{\partial}{\partial x_i} \left[(r_i' - r_i) \cdot \frac{\partial r_j}{\partial x_i} \right] -$$

$\int_V dV \frac{\partial}{\partial x_i} (r_i' - r_i) \cdot \frac{\partial r_j}{\partial x_i} = 0$ SURFACES

$$- \int_V dV (r_i' - r_i) \frac{\partial^2 r_j}{\partial x_i^2} = 0$$

$\int_V dV (r_i' - r_i) \frac{\partial^2 r_j}{\partial x_i^2} = 0$ INCOMPRES = 0

REYNOLDS EQUATION



- PROPERTIES OF TURBULENT FLOW:
- MANY TIME SCALES AND LENGTH SCALES
 - UNORDERED "RANDOM" STRUCTURE
 - NONLINEAR $(\vec{v} \cdot \nabla) \vec{v}$ IS DOMINANT, HIGH Re
 - STRONG DEPENDENCE ON INITIAL CONDITIONS

LET'S INTRODUCE AN IDEA ABOUT AVERAGING SCALE AND SEPARATE MACROSCOPIC FROM MICROSCOPIC MOTION \rightarrow MACROSCOPIC FLOW PATH \rightarrow MICROSCOPIC FLUCTUATIONS

AVERAGE OF THE VELOC. FLUCTUATION = 0

$\vec{v} = \langle \vec{v} \rangle + \vec{v}'$; $p = \langle p \rangle + p'$ WITH $\langle \vec{v}' \rangle = \langle p' \rangle = 0$

AVERAGE OF THE PRESSURE FLUCTU = 0

WHERE $\langle \dots \rangle$ ARE ENSEMBLE AVERAGES. SUBSTITUTE INTO NAVIER-STOKE EQUATION

$$\rightarrow \rho \cdot \frac{\partial}{\partial t} \langle v_i \rangle + \rho \frac{\partial}{\partial x_j} \langle v_i v_j \rangle = \rho \frac{\partial}{\partial x_i} \phi - \frac{\partial}{\partial x_i} p + \mu \frac{\partial^2}{\partial x_j^2} \dots$$

THIS LOOKS LIKE KIN. EN. PRESSURE

$$\cdot \left(\frac{\partial \langle v_i \rangle}{\partial x_j} + \frac{\partial \langle v_j \rangle}{\partial x_i} - \frac{2}{3} \frac{\partial}{\partial x_k} \langle v_k \rangle \cdot \delta_{ij} \right) + \rho \frac{\partial}{\partial x_j} \langle v_i' v_j' \rangle$$

τ_{ij}

WE DEFINE REYNOLDS TENSOR $\tau_{ij} = \langle v_i' v_j' \rangle$ AS IT IS SORT OF A VARIANCE OF COMPONENTS OF THE VELOCITY (ALL THE FLUCTUATIONS OF THE VELOCITY), MACROSCOPIC MOTION DEPENDS ON MICROSCOPIC VELOCITY VARIANCE.

• $\text{tr} \underline{C} = \sum_i C_{ii} = \langle \pi_i' \pi_i' \rangle \sim \text{TURBULEN PRESSURE}$

A33 BOUSSINESQ MODEL - PHENOMENOLOGICAL MODEL FOR \underline{C}_{ij}

TOHLECO JE NA PREDCHOZI STRANE

$$\rho \cdot \underline{C}_{ij} = \mu_e \cdot \left(\frac{\partial \langle \pi_i' \rangle}{\partial x_j} + \frac{\partial \langle \pi_j' \rangle}{\partial x_i} - \frac{2}{3} \frac{\partial \langle \pi_k' \rangle}{\partial x_k} \delta_{ij} \right) + \frac{2}{3} \rho \cdot k \cdot \delta_{ij}$$

↑
KONST.

WITH $k = \frac{1}{2} \text{tr} \langle \pi_i' \pi_i' \rangle$ AND FIXED μ_e

μ_e & k ^{THEY} SHOULD SOMEHOW MODEL HOW THE FLOW MACROSCOPICALLY BEHAVE.

A34 CORRELATION FUNCTIONS OF THE VELOCITY FIELD

IDEA: FULLY DEVELOPED TURBULANCE IS CHARACTERISED BY RANDOM FIELDS, BUT THERE WILL BE CORRELATION IN THE DENSITY FIELD

PROBABILITY DISTRIBUTION FOR VELOCITY AT THE POINT \vec{x}

$$P(\vec{v}(\vec{x}); \vec{v}(\vec{y})) \neq P(\vec{v}(\vec{x})) \cdot P(\vec{v}(\vec{y})) \quad \text{FOR THE VELOCITY FIELD } \vec{v}(\vec{x})$$

IF THERE ARE CORRELATIONS ASSUME GAUSSIAN STATISTICS (EVEN THOUGH THEY ARE NOT TRUE IN THIS CASE)

$$P(\vec{v}(\vec{x}); \vec{v}(\vec{y})) = \frac{1}{\sqrt{(2\pi)^6 \cdot \det(\underline{C})}} e^{-\frac{1}{2} (\vec{v}(\vec{x}); \vec{v}(\vec{y})) \cdot \underline{C}^{-1} \begin{pmatrix} \vec{v}(\vec{x}) \\ \vec{v}(\vec{y}) \end{pmatrix}}$$

↑
NORMALIZATION FACTOR

↑
C-MATRIX

$(\vec{v}(\vec{x}); \vec{v}(\vec{y}))$ - FORMS SIX DIMENSIONAL VECTOR

COVARIANCE MATRIX

$$C = \begin{pmatrix} \langle \vec{v}(x) \otimes \vec{v}(x) \rangle & \langle \vec{v}(x) \otimes \vec{v}(y) \rangle \\ \langle \vec{v}(x) \otimes \vec{v}(y) \rangle & \langle \vec{v}(y) \otimes \vec{v}(y) \rangle \end{pmatrix}$$

3x3 MATRIX

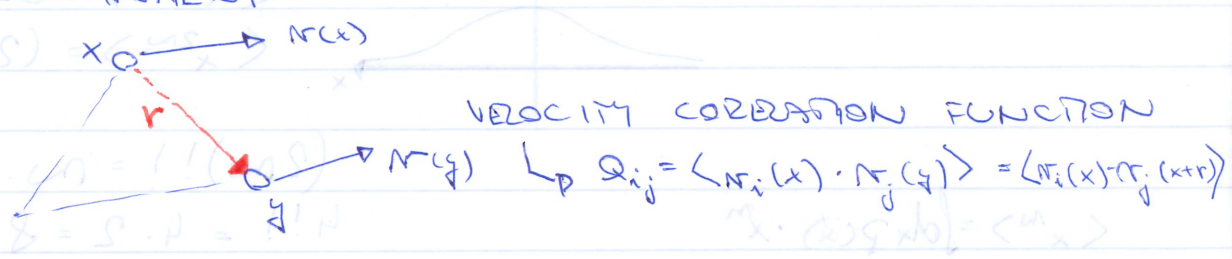
LETS DEFINE VELOCITY CORRELATION FUNCTION

$$Q = \langle \vec{v}(y) \otimes \vec{v}(x) \rangle ; Q_{ij} = \langle v_i(x) \cdot v_j(y) \rangle$$

$Q_{ij}(\vec{r}) = Q_{ij}(-\vec{r})$ FOR ISOTROPIC TURBULENCE

$Q_{ij}(\vec{r}) = Q_{ij}(\vec{r})$ HOMOGENEOUS TURBULENCE

DESCRIBES HOW i -COMPONENT OF THE VELOCITY IS RELATED TO THE j -COMPONENT



PROPERTIES OF THE VELOCITY CORRELATION

① $\frac{\rho}{2} Q_{ij}(0) = \frac{\rho}{2} \langle \vec{v}^2 \rangle \sim$ KINETIC ENERGY DENSITY

② $|Q_{ij}(r)| \leq |Q_{ij}(0)|$ CAUCHY-SWARZ INEQUALITY

③ $\rho \cdot Q_{ij}(0) = \epsilon_{ij}$ REYNOLDS STRESS

④ $\frac{\partial}{\partial r_i} Q_{ij} = \frac{\partial}{\partial r_j} Q_{ij} = 0$ FOR INCOMPRESSIBLE FLUIDS

PARAMETRISATION OF Q

$$\mathcal{R}(r) = \frac{1}{2} \text{tr} Q = \frac{1}{2} \langle \vec{v}(x) \cdot \vec{v}(y) \rangle$$

$$\left. \begin{aligned} \mathcal{R}^2 \cdot \delta = Q_{xx} (r \cdot \vec{e}_x) \\ \mathcal{R}^2 \cdot \gamma = Q_{yy} (r \cdot \vec{e}_x) \end{aligned} \right\} \text{ WITH } \mathcal{R}^2 = \sqrt{\frac{1}{3} \langle \vec{v}^2 \rangle}$$

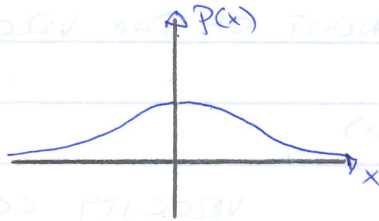
A35 NON-GAUSSIAN STATISTICS

TURBULENCE IS NOT CHARACTERIZED BY GAUSSIAN DISTRIBUTION. TURBULENT FLOWS DON'T FOLLOW GAUSSIAN STATISTICS.

DISTRIBUTION $p(x) dx$ IF IT IS A GAUSSIAN ALL YOU NEED TO KNOW IS $\langle x^2 \rangle \rightarrow p(x) = \frac{1}{\sqrt{2\pi \langle x^2 \rangle}} \cdot e^{-\frac{x^2}{2\langle x^2 \rangle}}$

$$\langle x^2 \rangle = \int dx p(x) \cdot x^2$$

IF WE MEASURE HIGHER ORDER MOMENTS IN GAUSSIAN:



$$\langle x^{2m+1} \rangle = 0$$

$$\langle x^{2m} \rangle = (2m-1)!! \langle x^2 \rangle^m$$

$$\langle x^m \rangle = \int dx p(x) \cdot x^m$$

$$(2m)!! = m \cdot (m-2)!!$$

$$4!! = 4 \cdot 2 = 8$$

$$4! = 4 \cdot 3 \cdot 2 = 24$$

$$\begin{aligned} \Phi(k) &= \int dx p(x) \cdot \exp(itx) = \int dx p(x) \cdot \sum_n \frac{(itx)^n}{n!} = \sum_n \frac{(it)^n}{n!} \cdot \int dx p(x) \cdot x^n \\ &= \sum_n \frac{(it)^n}{n!} \langle x^n \rangle \end{aligned}$$

SERIES EXPANSION

CHARACTERIZE NON-GAUSSIAN STATISTICS BY HIGHER ORDER MOMENTS:

$$S_{ijk}(\vec{r}) = \langle n_i(x); n_j(x); n_k(x+r) \rangle$$

DEVIDE OUT SCALING

$$N^3 k(r) = \langle n_x^2(x) \cdot n_x(x + e_x \cdot r) \rangle$$

A36 GENERAL SLOPES OF VELOCITY TENSOR

(HOMOGENOUS & ISOTROPIC)

$$\tau_{ij} = A(r) \cdot r_i r_j$$

$$Q_{ij} = A(r) \cdot r_i r_j + B(r) \cdot \delta_{ij}$$

$$S_{ijk} = A(r) \cdot r_i r_j r_k + B(r) r_i \delta_{jk} + C(r) \cdot r_j \delta_{ik} + D(r) \cdot r_k \delta_{ij}$$

$Q_{ij} = \underline{A}$ AND \underline{B} MUST BE RELATED TO f AND g

$$B = r^2 \cdot g \quad A = \frac{r^2}{r^2} \cdot (f-g) \rightarrow Q_{ij} = r^2 \cdot \left(\frac{f-g}{r^2} \cdot r_i r_j + g \delta_{ij} \right)$$

if $\text{div } \tau = 0 \cdot \frac{\partial}{\partial r_i} Q_{ij} = 0 = \underbrace{(r A' + 4A + \frac{B'}{r})}_{=0} \cdot r_j = 0$

SUBSTITUTE f AND $g \quad g = f + \frac{r}{2} f'$

$$\rightarrow Q_{ij} = \frac{r^2}{2r^2} \cdot [(r^2 f)' \cdot \delta_{ij} - f' r_i r_j]$$

IN ISOTROPIC INCOMPRESSIBLE TURBULENCES $R = \text{tr } Q$ DETERMINE

~~$R = \frac{r^2}{2r} \cdot (r^3 f')$ DEFINE $e = \int dr f$~~

THE CORRELATIONS COMPLETELY.

$R = \frac{r^2}{2r} (r^3 f')$ DEFINE $e = \int dr f$ (INTEGRAL SCALE OF TURBULENCE)

$$r^2 \cdot e = r^2 \int dr f(r) = \int dr R(r)$$

A37 DYNAMICS OF TURBULENCE \rightarrow KARMAN - HOWARTH EQ.

- USES NAVIER-STOKES-EQUATION FOR PREDICTING THE TIME EVOLUTION OF Q_{ij}

$$\frac{\partial}{\partial t} \tau_{ij} = - \frac{\partial}{\partial x_k} (\tau_{ij} \tau_k) - \frac{1}{\rho} \frac{\partial}{\partial x_i} p + \mu \Delta \tau_{ij} \quad \text{FLOW AT } @ \vec{x} \quad | \cdot \tau_j^i$$

$$\frac{\partial}{\partial t} \tau_j^i = - \frac{\partial}{\partial x_k} (\tau_j^i \tau_k^i) - \frac{1}{\rho} \frac{\partial p^i}{\partial x_j^i} + \mu \Delta^i \tau_j^i \quad \text{@ } \vec{x}^i \quad | \cdot \tau_i$$

NOW WE ADD BOTH EQUATIONS TOGETHER

$$\begin{aligned} \frac{\partial}{\partial t} \langle \pi_i \pi_j' \rangle = & - \langle \pi_i \frac{\partial}{\partial x_i'} (\pi_j' \pi_k') + \pi_j' \frac{\partial}{\partial x_k} (\pi_i \pi_k) \rangle - \\ & - \frac{1}{\rho} \langle \pi_i \frac{\partial}{\partial x_i} p' + \pi_j' \frac{\partial p'}{\partial x_i} \rangle + \\ & + \mu \cdot \langle \pi_i \Delta \pi_j' + \pi_j' \Delta \pi_i' \rangle \end{aligned}$$

LET'S SIMPLIFY THIS A LITTLE BIT $\frac{\partial}{\partial t} \langle \dots \rangle$ AND $\frac{\partial}{\partial t}$ COMPUTARE

WITH $\langle \dots \rangle$ (BOTH LINEAR) $\frac{\partial}{\partial x_i} \langle \dots \rangle = \frac{\partial}{\partial x_i} \langle \dots \rangle$

π_i DOES NOT DEPEND ON x' AND π_j' NOT ON x

$$\rightarrow \frac{\partial}{\partial t} Q_{ij} = \frac{\partial}{\partial r_k} (S_{ikj} + S_{jki}) + 2\mu \cdot \Delta Q_{ij}$$

BECAUSE $\langle \pi_i p' \rangle = \langle \pi_j' p \rangle = 0$

NOW LET'S SUBSTITUTE Q_{ij} AND S_{ijk} IN TERMS OF f AND k :

$$\frac{\partial}{\partial t} (r^2 r^4 f) = r^3 \frac{\partial}{\partial r} (r^4 k) + 2\mu r^2 \cdot \frac{\partial}{\partial r} (r^4 f')$$

THIS IS KARMAN-HOWARTH EQUATION

KARMAN-HOWARTH EQUATION DESCRIBES EVOLUTION OF VELOCITY CORRELATIONS

- GENERALISM TO THE n -TH MOMENT
- EVOLUTION OF $\langle r^n \rangle$ DEPENDS ON $\langle r^{n+1} \rangle$
- GENERATION OF SKEWNESS $\langle r^3 \rangle$ FROM $\langle r^4 \rangle \approx 3 \langle r^2 \rangle^2$
- $(\frac{\partial}{\partial t} + \Delta)$ DIFFUSION

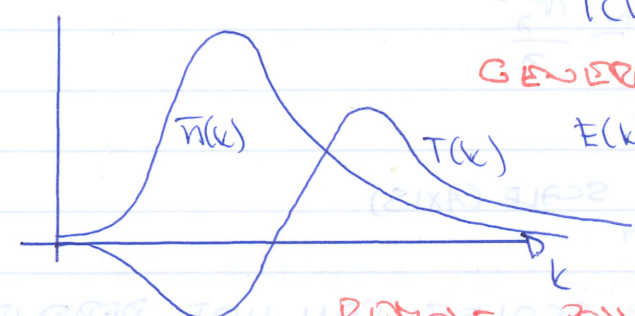
KARMA-HOWARTH EQUATION IN TERM OF R

$$\frac{\partial}{\partial t} R = \Gamma(r) + 2\mu \Delta R \text{ WITH } \Gamma = \frac{1}{2r^2} \frac{\partial}{\partial r} \cdot \frac{1}{r} \frac{\partial}{\partial r} (r^2 r^3 k)$$

EXPANSION

INTRODUCE ENERGY SPECTRUM: $E(k) = \frac{2}{\pi} \int dr R(r) \cdot (kr) \sin(kr) \Rightarrow R(r) = \int dk \cdot E(k) \cdot \frac{\sin(kr)}{kr}$

$\rightarrow \frac{\partial}{\partial t} E = \frac{T(k)}{k} - 2\mu k^2 E(k)$ $T \sim$ SPECTRAL ENERGY TRANSFER FUNCTION



$T(k) = -\frac{\partial}{\partial k} \Pi$
 GENERATE POWER ON SMALL SCALES
 $E(k) \sim k^{-5/3}$
 KOLMOGOROV - SPECTRUM
 REMOVE POWER ON LARGE SCALES

A39 TURBULENCE SUMMARY

	DYNAMICS	SPECTRUM	CORRELATION
QUANTITY	REYNOLDS TENSOR $\langle u_i' u_j' \rangle$	ENERGY SPECTRUM $E(k)$	FOURIER TRANSFORM VELOCITY CORRELATION $\langle u_i(x) u_j(x+r) \rangle$
PHENOMENON	TURBULENT PRESSURE $\frac{1}{3} \text{tr}(\mathbf{D} \cdot \mathbf{e})$ BOUSSINESQ APPROXIMATION (MODEL)	KOLMOGOROV CASCADING	GENERATION OF NON-GAUSSIAN STATISTICS
EVOLUTION	REYNOLDS AVERAGE NAVIER-STOKES EQUATION	SPECTRAL ENERGY FLUX Π	KARMAN HOWARTH EQUATION

A40 TURBULENCE: KOLMOGOROV CASCADING

- VORTICITIES ARE UNSTABLE NOW THE KELVIN-HELMHOLTZ-INSTABILITY
 \rightarrow BREAKUP INTO SECONDARY VORTICIES

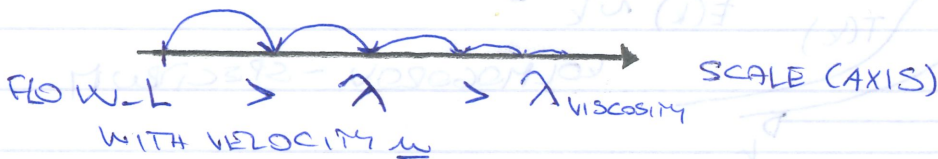
55
 - PROCESS TO MOST AND A CASCADE OF VORTICITIES AT EVEN SMALLER SCALES IS ESTABLISHED

→ KOLMOGOROV CASCADE

- ON THE SMALLEST SCALES THE REYNOLDS NUMBER BECOMES SMALL AND VISCOSITY DISSIPATES THE VORTEX

LOOK AT VORTICITIES OF SIZE λ , WITH VELOCITY $v_\lambda \approx \omega \cdot \lambda$, KINETIC ENERGY $E \sim \frac{1}{2} v_\lambda^2$ & ENERGY DISSIPATION:

$$\dot{E} = \frac{dE}{dt} = \frac{\Delta E}{\Delta t} = \frac{v_\lambda^2}{2} \cdot \frac{v_\lambda}{\lambda} \approx \frac{v_\lambda^3}{\lambda}$$



ENERGY FLOW THROUGH THE SCALES CAN NOT DEPEND ON λ .

→ OTHERWISE THERE WOULD BE AN ACCUMULATION

→ ENERGY AS A FUNCTION OF SCALE SHOULD BE A

POWER LAW.

$\dot{E} \approx \frac{v_\lambda^3}{\lambda}$ IF \dot{E} IS CONSTANT AND NOT A FUNCTION

OF SCALE THEN $v_\lambda \sim \lambda^{1/3}$. $v_\lambda(\lambda) = \omega \left(\frac{\lambda}{L}\right)^{1/3}$

$$\omega = \frac{v_\lambda}{\lambda} = \frac{v_\lambda}{\lambda} \cdot \left(\frac{\lambda}{L}\right)^{3/3} = \omega \cdot \left(\frac{1}{(\lambda^2 \cdot L)^{1/3}}\right)$$

WHILE THE VELOCITY DECREASES TOWARDS λ_{visc} , THE VORTICITY INCREASES.

VISCOUS FORCES DISSIPATE ENERGY AND GENERATE HEAT

$$\eta \cdot \left(\frac{v_\lambda}{\lambda}\right)^2 \sim \eta \cdot \left(\frac{v_\lambda^3}{\lambda}\right)^{2/3} \cdot \lambda^{-1/3} = \eta \cdot \dot{E}^{2/3} \cdot \lambda^{-1/3} = \rho \cdot \dot{E}_{visc}$$

STEADY SITUATION ENERGY FLOW = DISSIPATION RATE

SET $\lambda = \lambda_{visc}$ $\lambda_{visc} = \left(\frac{\eta \cdot L^{1/3}}{\rho \cdot \omega}\right)^{3/4} = L \cdot \left(\frac{\eta}{\rho \cdot L}\right)^{3/4} = \frac{L}{Re^{3/4}}$

KOLMOGOROV SPECTRUM

$\langle v_\lambda^2 \rangle \sim \lambda^3 \cdot \lambda^{-2/3} = \lambda^{11/3} = k^{-11/3}$ WITH WAVE/NUMBER
FROM FOURIER $v_\lambda \sim \lambda^{1/3}$ $k = \frac{2\pi}{\lambda}$

FOR ENERGY

$E(k) = k^2 \cdot k^{-11/3} = k^{-5/3}$

KOLMOGOROV SCALING

A41 REVERSIBILITY OF THE EULER-EQUATION

$t \rightarrow -t ; \vec{v} \rightarrow -\vec{v} ; \rho \rightarrow -\rho$

$\frac{\partial}{\partial t} \vec{v} + (\vec{v} \cdot \nabla) \vec{v} = -\frac{\nabla p}{\rho} - \nabla \phi + \mu \Delta \vec{v}$

TIME REVERSAL INVARIANT BUT A VISCOUS TERM WOULD BREAK IT.

VISCOSITY \rightarrow DISSIPATION OF ENERGY, HEAT CAN NOT BE CONVERTED INTO KINETIC ENERGY

VORTICITY + MOMENTUM DIFFUSION IS NOT EVOLVING BACKWARDS

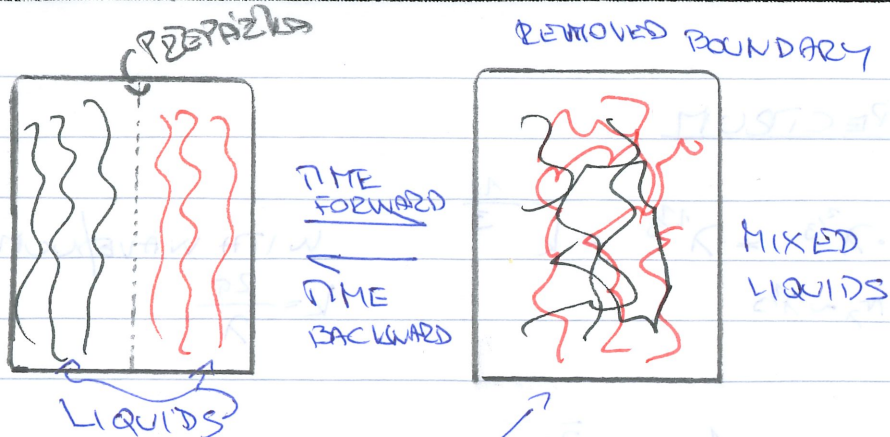
WHAT HAPPENS IN TURBULENCE?

- KOLMOGOROV - CASCADING
- LARGE VORTICITIES TRANSFER POWER TO SMALL SCALES
- SPECTRAL ENERGY FLUX
- KARMAN-HOWARTH - GENERATION OF NON-GAUSSIAN
- TURBULENT PRESSURE FROM MICROSCOPIC MOTION
- KEVIN-HELMHOLTZ - INSTABILITY

FEATURES

PROBABILISTIC ARGUMENT FOR MACROSCOPIC TIME-NON-REVERSIBILITY THROUGH EXP.

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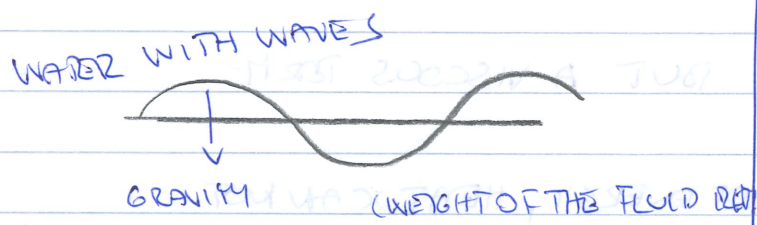


THESE LIQUIDS DON'T UNMIX, BUT THERE IS A SPECIAL INITIAL CONDITION FOR UNMIXING!

ALZ WAVES & INSTABILITY

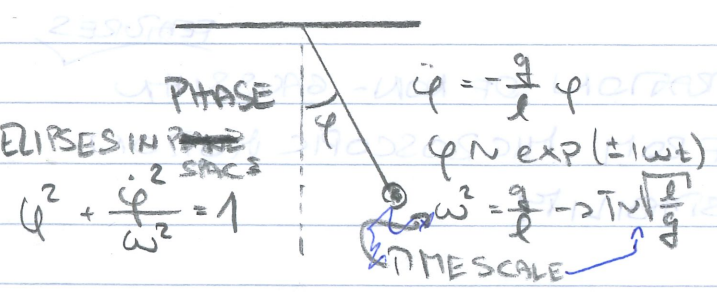
- TWO TYPES OF WAVES

GRAVITY WAVES



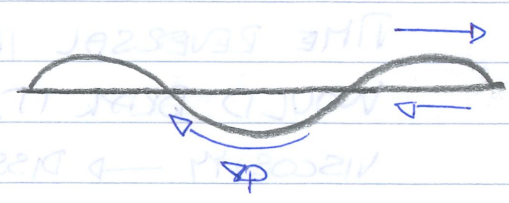
RESTORING FORCES GRAVITY & BUOYANCY (WHICH IS IN FACT GRAVITY)
 RESTORE $\Delta \phi \downarrow$ DISPLACEMENT
 STABLE, OSCILLATING

PENDULUM



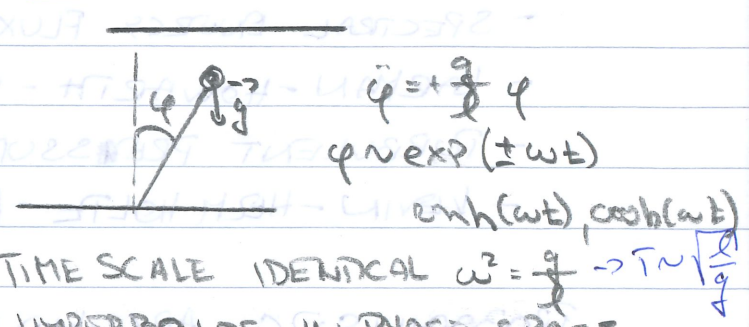
CONVEX POTENTIAL $\propto \phi^2$
 BECAUSE $\sin^2(\omega t) + \cos^2(\omega t) = 1$
 INITIALLY CLOSE POINTS STAY CLOSE

KELVIN-HELMHOLTZ INSTABILITY



PRESSURE GRADIENT Δp
 AMPLIFIES WAVE $\Delta p \perp \Delta$ DISPLACEMENT
 UNSTABLE, EXPONENTIALLY GROWING.

INVERTED PENDULUM



HYPERBOLAE IN PHASE SPACE
 BECAUSE $\cosh^2(\omega t) - \sinh^2(\omega t) = 1$
 SEPARATION $\propto \exp(\omega t)$ OF INITIAL CLOSE POINTS SEPARATES

A43 WAVES - TAXONOMY OF WAVES → DEPENDS ON RESTORING FORCE

- TYPES OF WAVES CAN BE CHARACTERIZED BY THE TYPES OF RESTORING FORCES

$$\frac{\partial}{\partial t} \vec{v} + (\vec{v} \cdot \nabla) \vec{v} = -\frac{\nabla p}{\rho} - \nabla \phi + \mu \Delta \vec{v} + 2 \vec{\omega} \times \vec{v} + \text{CAPILLARY} + \frac{g}{\rho} \cdot \vec{b} \times \vec{v}$$

SOUND
GRAVITY
NON-NEWTONIAN ELASTIC
CORIOLIS FORCE (ROSSBY WAVES)
SURFACE TENSION
LORENZ FORCE (PLASMA)

HARMONIC WAVES → LINEAR SUPERPOSITION MUST HOLD
 → RESTORING FORCE PROPORTIONAL TO AMPLITUDE.

VELOCITY OF THE WAVE = $\frac{\text{MAGNITUDE OF RESTORING FORCE}}{\text{INERTIA OF THE SYSTEM}}$

SOUND WAVES - RESTORING FORCE = PRESSURE

PRESSURE IS LINKED TO ρ THROUGH THE EQUATION OF STATE $p = p(\rho)$.

LINEARIZATION:

$p = p_0 + \delta p$; $\rho = \rho_0 + \delta \rho$; $\vec{v} = \delta \vec{v}$ WITH $v_0 = 0$
 WITH CONSTANT AND HOMOGENEOUS p_0 ; ρ_0 ; $\vec{v}_0 = 0$

- CONTINUITY $\frac{\partial}{\partial t} \delta \rho + \rho_0 \text{div} \delta \vec{v} = 0$

BECOMES THIS $\text{div}(\rho \vec{v}) = \nabla p \cdot \vec{v} + \rho \Delta \vec{v} = 0 =$

$= \nabla \delta p \cdot \delta \vec{v} + (\rho_0 + \delta \rho) \text{div} \delta \vec{v} \approx \rho_0 \text{div} \delta \vec{v}$

EULER EQUATION : $\frac{\partial}{\partial t} \vec{v} + \frac{1}{\rho} \nabla \partial p = 0$ (if $\mu=0, \nabla \phi=0$)

BECAUSE $\frac{1}{\rho_0 + \partial \rho} \cdot \nabla (p_0 + \partial p) = \frac{1}{\rho_0} \nabla \partial p$

IF $\nabla p_0 = 0$ AND $\rho_0 \gg |\partial \rho|$

RECAST $\nabla \partial p$ IN TERMS OF $\nabla \partial p$ USE AN EQUATION OF STATE $\partial p = \left. \frac{\partial p}{\partial \rho} \right|_{p_0} \partial \rho$ AS A FUNCTION OF ρ

DERIVATIVE, EVALUATED AT ρ_0

$\rightarrow \frac{\partial^2}{\partial t^2} \partial p - c^2 \Delta \partial p = 0$ WITH $c = \sqrt{\left. \frac{\partial p}{\partial \rho} \right|_{p_0}}$ BY

ELIMINATING $\frac{\partial}{\partial t} \text{div } \vec{v}$ FROM BOTH EQUATIONS RESTORING FORCE IS THE ADIABATIC COMPRESSIBILITY (THERE ARE TEMPERATURE CHANGES IN A SOUND WAVE, BUT THERE IS NO LOSS OF THERMAL ENERGY $\partial Q = 0$)

$\left(\frac{\partial p}{\partial \rho} \right)_s = \frac{c_p}{c_v} \left(\frac{\partial p}{\partial \rho} \right)_T$ WITH $\kappa = \frac{c_p}{c_v}$ ADIABATIC COEFFICIENT

@ CONST. ENTROPY \equiv @ AT CONST. TEMPERATURES
ADIABATIC ISOTHERMAL

ADIABATIC SOUND SPEED

SOUND SPEED FOR AN IDEAL GAS $c = \sqrt{\kappa \frac{RT}{M}}$ BECAUSE OF $p = \frac{RT}{M} \rho$
 M - MOLECULAR WEIGHT

WAVE EQUATION $\frac{\partial^2}{\partial t^2} p - c^2 \Delta p = 0$ c - SPEED OF SOUND
 $c = \sqrt{\left. \frac{\partial p}{\partial \rho} \right|_{p_0}}$
SOUND WAVE PROPAGATE

ISOTROPICALLY, TIME IS REVERSABLE

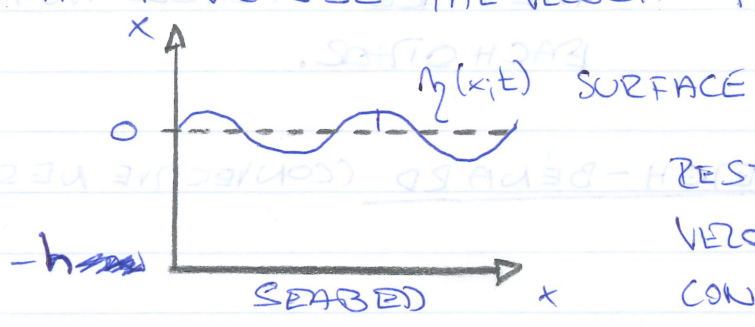
RELATIVISTIC WAVE - c - SPEED OF LIGHT

SHALLOW AIRY-LAPLACE WAVES

RESTORING FORCE - GRAVITY

- THESE ARE SURFACE WAVES IN WATER - RESTORING FORCE IS SIMPLY THE WEIGHT OF THE FLUID.

AGAIN WE USE THE VELOCITY POTENTIAL ϕ WITH $\vec{v} = \nabla\phi$



RESTORING FORCE - GRAVITY (BUOYANCY)
 VELOCITY POTENTIAL $\nabla\phi = \vec{v}$
 CONSTANT DENSITY, $\text{div} \vec{v} = 0$
 $\rho = \text{konst.}$ (INCOMPRESSIBLE FLUIDS)
 $\Delta\phi = 0$

BOUNDARY CONDITIONS $\vec{v}_z = 0$ AT $z = -h \rightarrow \frac{\partial\phi}{\partial z} = 0$
 LET'S LOOK AT THE SURFACE (THAT'S MOVING) - THAT MOVES UP & DOWN

FREE SURFACE: $\frac{\partial}{\partial t} \eta = v_z = \frac{\partial\phi}{\partial z}$ AT $z = \eta(x,t)$

DYNAMICS OF THE SURFACE $\frac{\partial\phi}{\partial t} = -gz$ COMES FROM CA
 BERNOULLI EQUATION $\frac{\partial\phi}{\partial t} + \frac{v^2}{2} + \frac{p}{\rho} + gz = \text{konst.}$

DYNAMIC WAVE Ansatz $\eta = \eta_0 \cdot \cos(kx - \omega t) \rightarrow$

$\rightarrow \phi = \frac{\omega}{k} \cdot \eta_0 \cdot \frac{\cosh[k(z+h)]}{\sinh(kh)} \cdot \sin(kx - \omega t)$

FROM A BOUNDARY CONDITION, DISPERSION $\omega^2 = gk \cdot \tanh(kh)$
 THIS IS WHAT WE GET WHEN WE SUBSTITUTE THIS TO THE WAVE EQUATION. BASED ON DEPTH OF THE WATER h WE GET WAVES FROM DIFFERENT REGIMES:

DEEP WATER: $\frac{h}{\lambda} = kh \gg 1 \rightarrow \omega^2 = gk$

SHALLOW WATER $\frac{h}{\lambda} = kh \ll 1 \rightarrow \omega^2 = ghk^2$

THEREFORE GROUP VELOCITY $\frac{\partial\omega}{\partial k} \neq \frac{\omega}{k}$ IS UNEQUAL TO THE PHASE VELOCITY.

Tsunami - DEEP WATER SMALL AMPLITUDES / SHALLOW WATER LARGE AMPLITUDES.

KELVIN-HELMHOLTZ INSTABILITY - TWO LAYERS OF LIQUID AND THEY ARE SHEARED AGAINST EACH OTHER.

A45 INSTABILITIES - RAYLEIGH-BÉNARD (CONVECTIVE INSTABILITY).

HOW THERMAL CONVECTION WORKS:

$$\frac{\partial T}{\partial t} + (\vec{v} \cdot \nabla) T = \kappa \nabla^2 T$$

$\frac{\partial T}{\partial t}$ TEMPERATURE CHANGE κ THERMAL CONDUCTIVITY $\nabla^2 T$ DIFFUSION

ADVECTION - DIFFUSION EQUATION FOR THE TEMPERATURE T WITH THERMAL CONDUCTIVITY κ

IN ANALOGY TO THE MOMENTUM TRANSPORT: (BOLE EQ.)

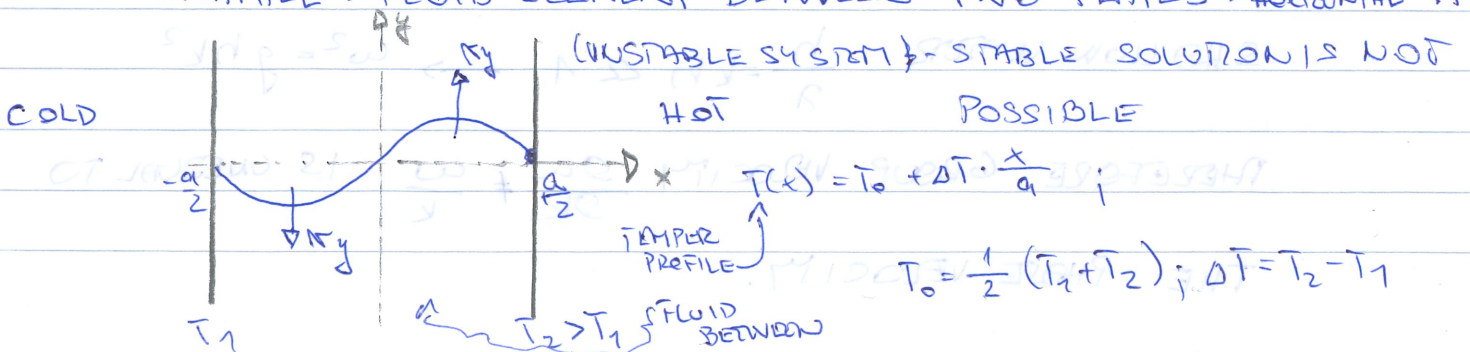
$$\frac{\partial \vec{v}}{\partial t} + (\vec{v} \cdot \nabla) \vec{v} = -\frac{\nabla p}{\rho} - \nabla \phi + \nu \nabla^2 \vec{v}$$

TEMPERATURE VARIANCES \rightarrow CHANGES IN VOLUME.

$\rho = \rho(p, T) \rightarrow$ DEPENDS ON TEMPERATURES AND REACTS BACK ON THE FLOW

THERMAL CONVECTION, T CHANGES ρ , GIVING RISE TO THE BUOYANT FORCES

EXAMPLE: FLUID ELEMENT BETWEEN TWO PLATES \rightarrow HORIZONTAL ∇T



INCOMPRESSIBLE FLUIDS. ρ DEPENDS ONLY ON TEMPERATURE

$$\rho = \rho_0 + \Delta\rho = \rho_0 \cdot [1 - \alpha(T - T_0)] = \rho_0 [1 - \alpha \cdot \Delta T \cdot \frac{x}{a}]$$

TEMP INCREASES FROM LEFT TO THE RIGHT, DENSITY INCREASES FROM LEFT TO RIGHT.

α - THERMAL EXPANSION COEFFICIENT

- VERTICAL LINES OF CONSTANT ρ

FLUID INITIALLY AT REST

$$\nabla p = \rho \cdot \nabla \phi \quad (\text{HYDROSTATIC EQUILIBRIUM})$$

→ LINES OF CONSTANT ρ COINCIDE WITH EQUIPOTENTIAL LINES (HORIZONTAL); $\nabla p \times \nabla \phi$ ~~PERPENDICULAR~~ PERPENDICULAR

GRADIENT OF PRESSURE IS IN THE SAME DIRECTION AS THE GRADIENT OF GRAV. POTENTIAL

→ ASSUME $\nabla \phi = \vec{g} = \text{CONST}$ ρ DEPENDS ONLY y (AXIS)
 ρ DEPENDS ONLY ON x (AXIS)
HYDROSTATIC EQUILIBRIUM IS NOT POSSIBLE

TEMPERATURE GRADIENT GENERATES DENSITY GRADIENT WHICH NOT PARALLEL TO THE PRESSURE GRADIENT SYSTEM WILL PICK UP SOME KIND OF ACCELERATION AND WILL BE UNSTABLE.

• $\nabla p = \rho \cdot \nabla \phi$ ∇p DEPENDS ON x , ρ DEPENDS ON y
→ EQUATION CANNOT BE FULLFIELD

• VORTICITY GENERATION $\nabla p \times \nabla \rho \neq 0$ BECAUSE ∇p IS NOT PARALLEL TO THE $\nabla \rho$ IN FACT IS PERPENDICULAR

NAVIER-STOKES EQUATION

- ASSUME STATIONARY CONDITION (ACCELERATION BALANCED BY VISCOUS FORCES)

$$-\frac{1}{\rho(x)} \cdot \frac{\partial p}{\partial y} - g + \mu \frac{\partial^2}{\partial x^2} \eta_y = 0 \quad \text{WITH } g = \nabla \phi = \nabla_y \phi$$

FOR THE y-COMPONENT, $x+z$ IN TRANSLATION INVARIANCE
ASSUMPTION (PERMEATION ANGLE)

NOW WE WILL INTRODUCE VARIATION OF ρ : $\rho(x) = \rho_0 + \rho(x)$

$$-\frac{1}{\rho} \frac{\partial p}{\partial y} - g + \mu \frac{\partial^2}{\partial x^2} \eta_y - \frac{\partial \rho}{\partial x} g = 0$$

BUSYANT FORCES - APPEAR WHEN THERE IS VARIATION IN ρ

$\rightarrow \frac{\partial \rho}{\partial x} \cdot g = -\alpha \cdot \rho_0 \cdot \frac{\partial T}{\partial x} \cdot g$, WITH THERMAL EXPANSION

BY SUBSTITUTING

$$\mu \frac{\partial^2}{\partial x^2} \eta_y + g \alpha \frac{\Delta T}{a} \cdot x = 0$$

THIS IS SOLVED BY $\eta_y(x) = -\frac{g \alpha}{\mu a} \cdot \frac{\Delta T}{6} \cdot x^3 + A x + B$

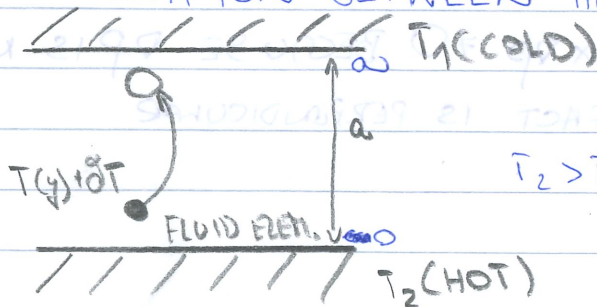
BOUNDARY CONDITION $\eta_y = 0$ AT $x = \pm \frac{a}{2}$

$$\rightarrow \eta_y(x) = -\frac{g \alpha}{\mu a} \cdot \frac{\Delta T}{6} \left(x^2 - \frac{a^2}{4} \right)$$

RAYLEIGH-BÉNARD INSTABILITY

- IS A FLOW IF IN NON-EQ?

- COMPETITION BETWEEN HEAT CONDUCTION ΔT AND ADVECTION



$T_2 > T_1$ SPHERICAL FLUID ELEMENT OF RADIUS r

DIFFUSIVE TRANSPORT OF HEAT - TRANSPORT OF HEAT & MATTER
PHYSICAL CONDITION OF INSTABILITY - IS THE VELOCITY
PERTURBATION \uparrow OR DAMPED OR WOULD IT GROW?

$$\frac{\partial}{\partial t} T + (\vec{v} \cdot \nabla) T = \kappa \nabla^2 T$$

FLUID ELEMENT TAKES THE , IT WANTS TO INCREASES
IT'S SIZE AND DECREASE ITS DENSITY IN THE PROCESS

$$\rho T = \frac{\partial T}{\partial y} \rho y = A \cdot \pi \frac{r^2}{\mu} \cdot \frac{\Delta T}{\omega} \rightarrow \rho = -\rho_0 \cdot \alpha \cdot \Delta T = *$$

THE SCALE

TEMP GRADIENT DISTANCE BY WHICH THE ELEMENTS MOVE

$$\rho = \pi \cdot \Omega ; \Omega = A \cdot \frac{r^2}{\mu}$$

↑
THE SCALE FOR CHANGING TEMPERATURE

$$* = -A \cdot \rho_0 \cdot \alpha \cdot \pi \cdot \frac{r^2}{\mu} \cdot \frac{\Delta T}{\omega}$$

BUOYANT FORCE: $F_b = -\frac{4}{3} \pi \cdot r^3 \cdot \rho \cdot g = \frac{4}{3} \pi \cdot A \cdot \rho_0 \cdot \alpha \cdot \frac{r^5}{\mu} \cdot \frac{\Delta T}{\omega}$

↑
VARIATION IN ρ

VISCOUS FORCES: $F_v = -6 \pi \eta \cdot r \cdot v$

ACCELERATION IF $F_b > F_v$

$$\frac{4\pi}{3} A \cdot \rho_0 \cdot \alpha \cdot g \cdot \pi \frac{r^5}{\mu} \frac{\Delta T}{\omega} > 6 \pi \eta \cdot r \cdot v$$

THE INSTABILITY CONDITION IS A STRONG FUNCTION OF r (BUT DOES NOT DEPEND ON μ), THEREFORE LARGEST FLUID ELEMENT TO GET AN UPPER ESTIMATE OF $F_b \Rightarrow r = \frac{g}{2}$ (MAX POSSIBLE)

$$Ra = \frac{\alpha g \cdot a^3}{\mu \kappa} \Delta T > Ra_c = \frac{72}{A}$$

RAYLEIGH NUMBER RAYLEIGH CRITICAL NUMBER

SMALL ELEMENTS EFFECTED BY SMALL BUOYANT FORCES (SMALL VOLUME)
LARGE ———— 4 ———— ACCELERATED BY BUOYANT FORCES .

FLOW BECOMES UNSTABLE IF THE VOLUME FORCE IS THE WINNING FORCE OVER THE VISCOUS FORCE.

$$\sigma_T + (\mu \nabla^2) T = \rho \sigma$$

IF σ RISES AND DECREASES ITS DIMENSION IN THE PROCESS FLUID ELEMENT TAKES THE TENDENCY TO INCREASE

$$\sigma_T = \rho \sigma \frac{dT}{Dx} = \rho \sigma \frac{v}{r} \pi \cdot A = \rho \sigma \frac{DT}{Dx} = \rho \sigma \frac{v}{r} \pi \cdot A$$

$$F_v = \mu \frac{dv}{dy} \cdot A = \mu \frac{v}{r} \cdot A$$

$$F_v = \mu \frac{v}{r} \cdot A \quad F_\sigma = \rho \sigma \frac{v}{r} \cdot A \cdot \pi \cdot r$$

viscous forces: $F_v = \mu \frac{dv}{dy} \cdot A$

ACCELERATION IF $F_\sigma > F_v$

$$\rho \sigma \frac{v}{r} \cdot A \cdot \pi \cdot r > \mu \frac{v}{r} \cdot A$$

THE INSTABILITY CONDITION IS A STRONG FUNCTION OF r (BUT DOES NOT DEPEND ON μ), THEREFORE LARGEST FLUID ELEMENT TO BE AN OPTIC ELEMENT OF $F_\sigma = \frac{\rho}{\mu} \sigma v r$ (MAX POSSIBLE)

$$\frac{\rho}{\mu} \sigma v r > \frac{\rho}{\mu} \sigma v r$$

THE CRITICAL VELOCITY IS THAT WHICH IS NOT DEPENDENT ON r (BUT DOES NOT DEPEND ON μ)