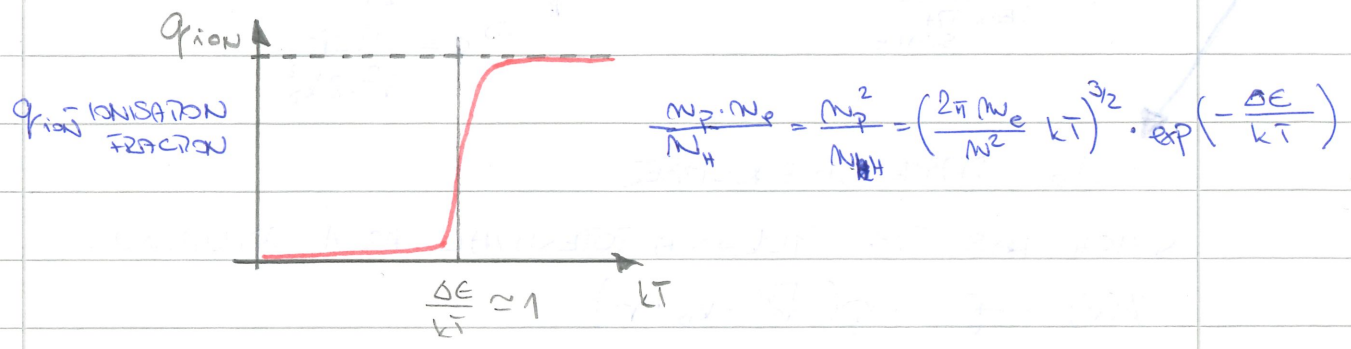


# PLASMA PHYSICS

## D1 PLASMA STATE - SEPARATION OF ELECTRONS AND IONS

IN A MEDIUM, RELEVANT IN ASTRO PHYSICS BECAUSE TEMPERATURES ARE VERY HIGH!!!

## SAHA - EQUATION: IONISATION EQUILIBRIUM OF A THERMAL PLASMA

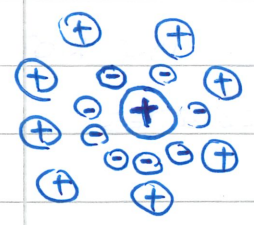


## DEBYE SHIELDING

SHIELDING OF A CHARGE WITHIN PLASMA

- A FREE CHARGE IN A PLASMA ATTRACTS OPPOSITE CHARGES WEAKENING THE ELECTRIC FIELD COUNTERACTED BY THERMAL DIFFUSION.

- MINIMISE POTENTIAL ENERGIES  $+e\phi$  AND  $-e\phi$  OF THE PARTICLES IN THERMAL EQUILIBRIUM.



$$n_i = N \cdot \exp\left(+\frac{e\phi}{kT}\right) \quad \text{"ATMOSPHERE OF } + \text{- CHARGE"}$$

$$n_e = N \cdot \exp\left(-\frac{e\phi}{kT}\right) \quad \text{CARRIES AROUND THE}$$

$N$  - MEAN CHARGE CARRIER DENSITY MAIN CHARGE "

$\phi$  IS DETERMINED BY THE POISSON-EQUATION

$$\Delta\phi = \underbrace{4\pi e(n_i - n_e)}_{\text{CLOUD OF CHARGES (POLARISED PLASMA)}} + \underbrace{4\pi q \delta_D(\vec{x})}_{\text{MAIN CHARGE}}$$

WITH THE BOUNDARY CONDITION  $\phi = 0$  AT INFINITY

LET'S ASSUME A HOT PLASMA ( $\frac{e^2}{kT} \ll 1$ ) WITH STRONG THERMAL DIFFUSION

$\Delta E \ll kT$  - BOLTZMAN-FACTOR IS CLOSE TO ONE  $\Rightarrow$  TAYLOR EXPANSION

WAVE NUMBER

$$\Delta \phi = \underbrace{\frac{8\pi e^2}{kT} n \phi}_{\text{INVERSE LENGTH SCALE}} - 4\pi q \cdot \delta_D \quad \parallel \quad \text{IN FOURIER SPACE}$$

$$(k^2 + \frac{8\pi e^2}{kT} n) \phi = 4\pi q$$

$$\Rightarrow \phi = \frac{4\pi q}{k^2 + 2k_D^2}$$

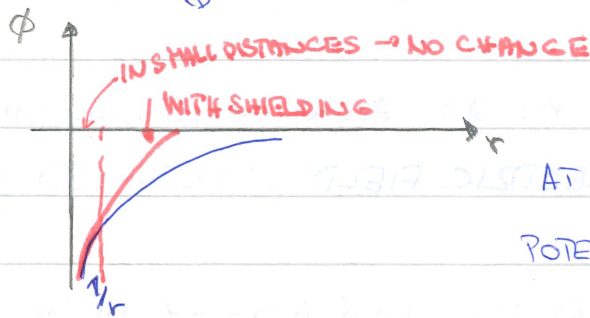
$= 2k_D$  DEBYE WAVE NUMBER

WHICH HAS THE YUKAWA POTENTIAL AS A SOLUTION:

$$\phi(r) = -\frac{q}{r} \cdot \exp(-\sqrt{2} \cdot k_D \cdot r)$$

DISTANCE  $\uparrow$  COULOMB

IF  $r \ll \lambda_D = \frac{2\pi}{k_D} \rightarrow \exp(k_D \cdot r) \approx 1$  NO SHIELDING EFFECT



AT  $r \approx \lambda_D \sim$  SHIELDING EFFECT  
POTENTIAL  $\rightarrow 0$  IN AN EXPONENTIAL WAVE

LENGTH SCALE  $\sim$  DEBYE LENGTH

$$\frac{\lambda_D}{2\pi} = \frac{1}{k_D} = \sqrt{\frac{kT}{4\pi n e^2}} \quad \lambda_D \text{ DEPENDS ON } n, e \text{ \& } T \text{ BUT NOT ON } q$$

PLASMA PROPERTIES

- THERMAL DIFFUSION VS. COULOMB FORCES

### PLASMA FREQUENCY

- WHAT'S THE TIME NEEDED FOR CHARGE CARRIERS TO COVER  $\lambda_D$  IF THEY MOVE AT THERMAL VELOCITIES.

- EQUIPARTITION  $\frac{1}{2} m_{TH} \langle v^2 \rangle = \frac{3}{2} kT \rightarrow v_{TH} = \sqrt{\frac{3kT}{m_{TH}}} \rightarrow t_D = \frac{\lambda_D}{v_{TH}} = \sqrt{\frac{m_e}{4\pi n e^2}}$  REACTION TIME OF THE PLASMA

$$\nu_D = \frac{1}{t_D} = \sqrt{\frac{4\pi n e^2}{m_e}} = 56 \text{ kHz } \sqrt{\frac{n}{\text{cm}^{-3}}}$$

# DL PLASMA MAGNETOHYDRODYNAMICS (MHD)

- ① TREAT IONISED PLASMA AS FLUIDS
- ② NO MACROSCOPIC CHARGE SEPARATION
- ③ NON-RELATIVISTIC MOTION

OHMS LAW  $\vec{j} = \sigma \cdot \vec{E}$

$\uparrow$  CONDUCTIVITY  $\sigma$        $\uparrow$  ELECTRIC FIELD

$\sigma$  IS IN PLASMA USUALLY VERY LARGE  $|\vec{j}| \gg |\vec{E}|$  CHARGES FOLLOW  $\vec{E}$  VERY FAST

$\sigma \sim \frac{1}{\mu_0 n e^2 \tau}$  →

IF  $\sigma$  IS TOO LARGE, ALL  $\vec{E}$  ARE SHORT CIRCUITED, THEREFORE  $\vec{E}$  ARE NOT THAT IMPORTANT IN ASTROPHYSICS.

THEN MAXWELL EQUATIONS LOOKS LIKE THIS:

$\text{div } \vec{B} = 0$  ;  $\text{div } \vec{E} = 4\pi \rho$  ;  $\text{rot } \vec{B} = \frac{4\pi}{c} \vec{j}$  →  $\partial_t \vec{E} \ll \vec{j}$  BECAUSE OF THE LARGE CONDUCTIVITY

$\text{rot } \vec{E} = \dots$  → INDUCTION EQUATION OHMS LAW

$\vec{E}$  IN THE PLASMA FRAME IS  $\vec{E} = \frac{1}{\sigma} \vec{j}$  → BOOST  $\vec{E}$  BY  $\vec{v}$

THEN WE GET  $\vec{E} = \frac{1}{\sigma} \vec{j} - \vec{v} \times \vec{B}$  IN THE LAB. FRAME

LET'S SUBSTITUTE INTO A ACTION EQUATION:

$\partial_{\mu} \vec{B} = \dots - \text{rot } \vec{E} = \text{rot} \left( \frac{1}{\sigma} \vec{j} - \vec{v} \times \vec{B} \right)$

NOW WE USE AMPÈRE'S LAW:  $\vec{j} = \frac{c}{4\pi} \text{rot } \vec{B}$

AND THE VECTOR IDENTITY  $\text{rot } \text{rot } \vec{B} = \nabla \text{div } \vec{B} - \Delta \vec{B}$   
 $= 0$

NOW WHEN WE PUT IT TOGETHER:

→  $\partial_{\mu} \vec{B} = \frac{c^2}{4\pi \sigma} \Delta \vec{B} - \text{rot} (\vec{v} \times \vec{B})$  INDUCTION EQUATION

$\underbrace{\hspace{10em}}_{\text{DIFFUSION TERM}}$        $\rightarrow$  DRIVING TERM,  $\propto \vec{v}$

DIFFUSIVITY  $\frac{c^2}{4\pi\sigma} \rightarrow \rho_{ce} \vec{B} = \frac{c^2}{4\pi\sigma} \Delta \vec{B}$  SD

$\sim \frac{\text{LENGTH}^2}{\text{TIME}}$

### BD 3 EULER-EQUATION FOR A MAGNETISED PLASMA

IDEA

- ① LORENTZ-FORCES ACT ON CHARGE CARRIERS
- ② CURRENTS ARE DRIVEN BY ELECTRIC FIELDS
- ③ MAGNETIC FIELDS ARE SOURCED BY CURRENTS

→ COUPLED PROCESSES IN THE EULER/CONTINUITY EQUATION

→ NEED TO INCORPORATE LORENTZ-FORCES

LORENTZ FORCE

$$\vec{F}_L = \frac{c}{4\pi} (\vec{\nabla} \times \vec{B}) \quad [\text{IN THE GAUSS. SYSTEM OF UNITS}]$$

LET'S MULTIPLY THIS WITH THE CHARGE CARRIER DENSITY  $\vec{n}$  AND IDENTIFY  $\vec{j} = \vec{n} \cdot e \cdot \vec{v}$  AS THE CURRENT DENSITY

$$\frac{e}{c} \cdot \vec{n} \times \vec{B} \cdot \vec{n} = \frac{1}{c} \cdot \vec{j} \times \vec{B} = \frac{1}{4\pi} (\text{rot } \vec{B}) \times \vec{B} \quad \begin{matrix} \text{SUBSTITUTE} \\ \text{WITH AMPERE'S LAW} \end{matrix}$$

$$\rightarrow \rho \frac{d\vec{r}}{dt} = -\nabla p - \rho \cdot \nabla \phi + \underbrace{\frac{1}{4\pi} \cdot (\text{rot } \vec{B}) \times \vec{B}}_{\text{LORENTZ FORCE}} \quad \begin{matrix} \text{EULER EQUATION} \\ \text{INCLUDING LORENTZ} \\ \text{FORCE} \end{matrix}$$

MATTER DENSITY

IDENTITY  $(\text{rot } \vec{B} \times \vec{B}) = (\vec{B} \cdot \nabla) \vec{B} - \frac{1}{2} \nabla (\vec{B}^2)$

(LECTURE ABOUT VORTICITY  $\vec{\omega} \times \vec{r} = (\vec{r} \cdot \nabla) \vec{r} - \frac{1}{2} \nabla (\vec{r}^2)$  VORTICITY)

WITH  $\vec{\omega} = \text{rot } \vec{v}$

$$\rightarrow \rho \frac{d\vec{r}}{dt} = -\nabla p - \rho \nabla \phi + \underbrace{\frac{1}{4\pi} (\vec{B} \cdot \nabla) \vec{B}}_{\text{①}} - \underbrace{\frac{1}{8\pi} \nabla (\vec{B}^2)}_{\text{②}}$$

MAGNET. PRESSURE      GRAD. POT.

#### TWO NEW TERMS

① CHARGE OF  $\vec{B}$  ALONG  $\vec{B}$  → PROJECTION OF  $\nabla \vec{B}$  ONTO  $\vec{B}$

~ PLASMA TIES TO KEEP  $\vec{B}$ -FIELD LINES STRAIGHT

② GRADIENT  $\nabla (\vec{B}^2)$  OF THE FIELD ENERGY DENSITY ACTS LIKE

A PRESSURE TERM, COUNTERACTING A COMPRESSION OF THE FIELD LINES → MAGNETIC PRESSURE.

## D4 EQUATION OF MAGNETO-HYDRODYNAMICS

① INDUCTION EQUATION

$$\partial_t \vec{B} = \frac{c^2}{4\pi\sigma} \cdot \Delta \vec{B} + \text{rot}(\vec{v} \times \vec{B})$$

② EULER-EQUATION

$$\partial_t \vec{v} + (\vec{v} \cdot \nabla) \vec{v} = - \frac{\nabla p}{\rho} - \nabla \phi - \frac{1}{4\pi\sigma} (\text{rot} \vec{B}) \times \vec{B}$$

③ CONTINUITY (UNCHANGED)

$$\partial_t \rho + \text{div}(\rho \vec{v}) = 0$$

*\* GENERATION OF MAGNETIC FIELD + COLLISION BETWEEN TWO COLLAPSED STARS*

④ EQUATION OF STATE  $p = p(\rho, T)$

IN ELECTRODYNAMICS  $\text{div} \vec{B} = 0 \rightarrow$  LET'S LOOK AT INCOMPRESSIBLE FLUIDS

$$\frac{d}{dt} \vec{B} = \frac{\partial}{\partial t} \vec{B} + (\vec{v} \cdot \nabla) \vec{B} = (\vec{B} \cdot \nabla) \vec{v} + \frac{c^2}{4\pi\sigma} \Delta \vec{B}$$

$\text{div} \vec{v} = 0$

viscosity  
+  $(\mu \Delta \vec{v})$

$$\frac{d}{dt} \vec{v} = \frac{\partial}{\partial t} \vec{v} + (\vec{v} \cdot \nabla) \vec{v} = - \frac{1}{\rho} \cdot \nabla(p + \vec{B}^2) + \frac{1}{4\pi\sigma} (\vec{B} \cdot \nabla) \vec{v}$$

USING THE RELATION (REMEMBER VORTICITY LECTURE)

$$\text{rot}(\vec{A} \times \vec{B}) = \underbrace{\vec{A} \text{div} \vec{B}}_{=0} - \underbrace{\vec{B} \text{div} \vec{A}}_{=0} + (\vec{B} \cdot \nabla) \vec{A} - (\vec{A} \cdot \nabla) \vec{B}$$

$\Rightarrow$  THESE TERMS VANISHES DUE TO INCOMPRESSIBILITY AND ABSENCE OF MONOPLES.

2 COUPLED, NONLINEAR EQUATIONS  $\rightarrow$  MHD IS DIFFICULT

## D5 ADVECTION AND DIFFUSION SCALES IN MHD

IN PAST  $\rightarrow \sigma$  - LARGE

FLOX FREEZING: DIFFUSION VANISHES IF THE CONDUCTIVITY  $\sigma$  IS LARGE

$\rightarrow$  THE  $\vec{B}$ -FIELD MOVES ALONG WITH  $\vec{v}$

MAGNETIC REYNOLDS NUMBER (WHICH IS USUALLY LARGE!)

$$\text{rot } \vec{n} \times \vec{B} \sim \frac{n \cdot B}{L} \quad \text{AS } \frac{c^2}{4\pi\sigma} \Delta B \sim \frac{c^2}{4\pi\sigma} \cdot \frac{B}{L^2}$$

RATIO OF THESE EQUATIONS...

LARGE  $R$  ADVECTION DOMINATES

SMALL  $R$ , DIFFUSION DOMINATES

MAGNETIC REYNOLDS NUMBER  $R = 4\pi\sigma \frac{nL}{c^2}$

INDUCTION EQUATION  $\partial_t \vec{B} = \frac{c^2}{4\pi\sigma} \Delta \vec{B}$ ; DIFFUSION OF  $\vec{B}$ -FIELDS

IF THE CONDUCTIVITY IS LARGE DIFFUSION DOES NOT PLAY A ROLE 'FLUX FREEZING'

MAG. REYN. NUM.  $R$  - DESCRIBES RELATIVE EFFECTIVENESS OF ~~ADVECTION~~ & DIFFUSION ADVECTION

## D6 GENERATION OF MAGNETIC FIELDS → "BATTERY MECHANISM"

- HOW TO GENERATE MAGNETIC FIELDS → ELECTRIC CURRENT - SEPARATION OF CHARGES  
→ THE CHANGING ELECTRIC FIELDS

① IN ORDER A CURRENT THAT CAN INDUCE A MAGNETIC FIELD, ONE NEEDS A CHARGE SEPARATION

② BUT NEGATIVE AND POSITIVE CHARGES IN A PASTA ARE TIGHTLY COUPLED, DUE TO COULOMB FORCES.  
→ NEED TO RELAX THIS IS AN ASSUMPTION.

③ THE INDUCTION EQUATION IS A DIFFUSION-ADVECTION EQUATION WITH NO SOURCE FIELDS: IT CAN WORK ON EXISTING MAGNETIC-FIELDS.

ONE WAY TO RELAX THIS IS AN ASSUMPTION

→ DROP THE ASSUMPTION ON TIGHT COUPLING TO MAKE POSITIVE AND NEGATIVE CHARGES AT DIFFERENT SPEEDS.  
→ THE EFFECTIVE CURRENT SOURCES A MAGNETIC  $\vec{B}$ -FIELD

STARTING POINT - TWO EULER EQUATIONS FOR + (POSITIVE) AND - (NEGATIVE) CHARGES

- ASSUME MACROSCOPIC CHARGE NEUTRALITY  $n_+ = n = n_-$
- DENSITIES  $\rho_+ = n_+ \cdot m_+ = n \cdot m_+$  WITH PARTICLE MASSES

$$\underbrace{n \cdot m_+}_{\rho_+} \cdot \frac{d\vec{v}_+}{dt} = -\underbrace{\nabla P_+}_{\text{GRADIENT PRESSURE}} + ne \cdot (\vec{E} + \frac{1}{c} \cdot \vec{v}_+ \times \vec{B}) - n m_+ \cdot \nabla \phi \quad / \cdot \frac{m_-}{m}$$

$$\underbrace{n \cdot m_-}_{\rho_-} \cdot \frac{d\vec{v}_-}{dt} = -\nabla P_- - ne \cdot (\vec{E} + \frac{1}{c} \cdot \vec{v}_- \times \vec{B}) - n m_- \cdot \nabla \phi \quad / \cdot \frac{m_+}{m}$$

→ NOW, LET'S CONSTRUCT A DIFFERENCE, SO WE ARE AFTER RELATIVE ACCELERATION:

$$\frac{d}{dt} (\vec{v}_- - \vec{v}_+) = -\frac{m_+}{m} \nabla P_- + \frac{m_-}{m} \nabla P_+ - \frac{e}{m_-} (\vec{E} + \frac{1}{c} \vec{v}_+ \times \vec{B}) + \frac{e}{m_+} (\vec{E} + \frac{1}{c} \vec{v}_- \times \vec{B})$$

IF  $m_+ \gg m_-$  (AS IN THE CASE OF IONS AND ELECTRONS)

$$\frac{d}{dt} (\vec{v}_- - \vec{v}_+) = -\frac{\nabla P_-}{m \cdot m_-} - \frac{e}{m_-} (\vec{E} + \frac{1}{c} \vec{v}_+ \times \vec{B})$$

- BECAUSE ACCELERATION ON THE POSITIVE CHARGE CARRIES ARE MUCH SMALLER DUE TO THEIR HIGH MASS.

LET'S INTRODUCE A COLLISION TERM SOURCING THE VELOCITY DIFFERENCE

$$\frac{d}{dt} (\vec{v}_- - \vec{v}_+) = -\frac{\nabla P_-}{m \cdot m_-} - \frac{e}{m_-} \cdot (\vec{E} + \frac{1}{c} \vec{v}_+ \times \vec{B}) - \frac{\vec{v}_- - \vec{v}_+}{\tau}$$

↳ TIME SCALE  $\tau$

NET CURRENT DENSITY  $\vec{j} = ne (\vec{v}_+ - \vec{v}_-)$

- RELATIVE ACCELERATION = 0 IN A STEADY STATE

~~LET'S INTRODUCE A COLLISION TERM SOURCING THE VELOCITY DIFFERENCE~~

$$\frac{d}{dt} (\vec{v}_- - \vec{v}_+) = 0 \rightarrow \vec{E} = -\frac{\nabla P_-}{ne} - \frac{\vec{v}_+}{c} \times \vec{B} + \frac{m_-}{m e^2} \cdot \vec{j}$$

MAKE THIS  $\vec{E}$  FIELD INDUCE A MAGNETIC FIELD

$$\partial_t \vec{B} = -\text{rot} \vec{E} \quad (\text{FARADAY'S LAW})$$

$$\rightarrow \partial_t \vec{B} = \frac{c}{e} \text{rot} \left( \frac{\nabla p}{\rho} \right) + (\vec{\Omega} \times \vec{B}) - \frac{m \cdot c}{e^2 \cdot \rho} \cdot \text{rot} \left( \frac{\vec{j}}{\rho} \right)$$

QUITE SIMILARLY TO THE DISCUSSION OF VORTICITY  
(TIME DERIVATIVE  $\partial_t \omega$ )

GENERATION OF MAGNETIC FIELDS: 2 RELEVANT TERMS:

$$\textcircled{1} \quad \text{rot} \left( \frac{\nabla p}{\rho} \right) = - \frac{\nabla p \times \nabla \rho}{\rho^2} \sim \nabla p \cdot \nabla \rho \quad \text{"BAROCLINIC TERM"}$$

$$\textcircled{2} \quad \text{rot} \left( \frac{\vec{j}}{\rho} \right) = \frac{c}{4\pi \rho} \text{rot rot} \vec{B} + \underbrace{\vec{j} \times \frac{\nabla \rho}{\rho^2}}_{=0 \text{ BECAUSE } \vec{j} \parallel \nabla \rho} = \frac{c}{4\pi \rho} \cdot \text{rot rot} \vec{B}$$

$$\rightarrow \partial_t \vec{B} = \frac{c^2}{4\pi \rho} \Delta \vec{B} + \text{rot} (\vec{\Omega} \times \vec{B}) - \frac{c}{e \rho^2} (\nabla p \times \nabla \rho)$$

COMPARE TO THE VORTICITY EQUATION FROM FLUID MECHANICS

$$\partial_t \omega = \nu \cdot \Delta \vec{\omega} + \text{rot} (\vec{\Omega} \times \vec{\omega}) - \frac{\nabla p \times \nabla \rho}{\rho^2} \sim \text{FORMAL ANALOGY WITH VORTICITY GENERATION OF FLUID MECHAN.$$

• BAROCLINIC TERM SOURCES  $\vec{B}$

• DIFFUSION OF THE MAGNETIC FIELD, VORTICITY  $\sim$  CONDUCTIVITY

• CORIOLIS OR CENTRIFUGAL FORCES CAN SOURCE  $\vec{B}$

- BIERMANN BATTERY

• SIMPLE EQUATIONS OF STATE  $p = p(\rho)$  WILL NOT SOURCE  $\vec{B}$

$$p(\rho) \rightarrow \nabla p \times \nabla \rho = \frac{\partial p}{\partial \rho} \underbrace{\nabla \rho \times \nabla \rho}_{=0} = 0$$



## D7 HYDROMAGNETIC WAVES

- PLASMA CAN CARRY WAVE LIKE EXCITATIONS
- LINEARISATION (TO ENSURE SUPERPOSITION) → SIMPLE DIFF. STRUCTURE
- NEW RESTORING FORCES <sup>DUE</sup> TO MAGNETIC FIELDS

LET'S SET CONDUCTIVITY  $\sigma \rightarrow \infty$ : NO DIFFUSION, PERFECT FREEZING

- ①  $\text{div } \vec{B} = 0$
- ②  $\nabla_{\perp} \vec{B} = \text{rot}(\vec{n} \times \vec{B})$
- ③  $\nabla_{\perp} \rho = -\text{div}(\rho \cdot \vec{n})$  [CONTINUITY EQUATION]
- ④  $\nabla_{\perp} \vec{n} + (\vec{n} \cdot \nabla) \vec{n} = -\frac{1}{\rho} \nabla p + \frac{1}{4\pi \rho} (\text{rot} \vec{B}) \times \vec{B}$  [EULER-EQUATION]

LET'S INTRODUCE PERTURBATION + LINEARISE:

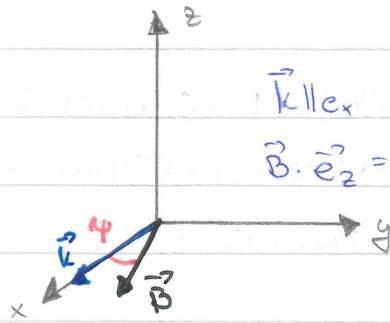
$$\rho = \rho_0 + \delta\rho; \quad \vec{B} = \vec{B}_0 + \delta\vec{B}; \quad p = p_0 + \delta p; \quad \vec{n} = \delta\vec{n} \quad [\text{PERTURBATION BY CHOICE OF FRAME}]$$

- ①  $\text{div } \delta\vec{B} = 0$
  - ②  $\nabla_{\perp} \delta\vec{B} = \text{rot}(\delta\vec{n} \times \vec{B}_0)$
  - ③  $\nabla_{\perp}(\delta\rho) = -\rho_0 \text{div}(\delta\vec{n})$
  - ④  $\nabla_{\perp}(\delta\vec{n}) = -\frac{1}{\rho_0} \cdot \nabla \delta p + \frac{1}{4\pi \rho_0} \text{rot}(\delta\vec{B}) \times \vec{B}_0$
- $\swarrow$   $\delta p = c_s^2 \cdot \delta\rho$   
 WITH THE ADIABATIC SOUND SPEED

NOW LET'S SUBSTITUTE A PLANE-WAVE (ANZATZ FOR ALL QUANTITIES)

- ①  $\vec{k} \cdot \delta\vec{B} = 0$
- ②  $-\omega \delta\vec{B} = \vec{k} \times \delta\vec{n} \times \vec{B}_0$
- ③  $-\omega \delta\rho = -\rho_0 \cdot \vec{k} \cdot \delta\vec{n}$
- ④  $-\omega \delta\vec{n} = -\frac{c_s^2}{\rho_0} \cdot \delta\rho \cdot \vec{k} - \frac{1}{4\pi \rho_0} (\vec{k} \times \delta\vec{B}) \times \vec{B}_0$

## COORDINATE FRAME



$$\left. \begin{aligned} \vec{k} \parallel \vec{e}_z \\ \vec{B} \cdot \vec{e}_z = 0 \end{aligned} \right\} \vec{k} \cdot \vec{B}_0 = k \cdot B_0 \cos \gamma$$

$\vec{B}$  IS IN THE  $xy$  PLANE

IN COMPONENTS WE HAVE NOW THE RELATIONS

EULER EQ.  $\rightarrow \mathcal{J} \pi_x = \frac{k^2}{\omega^2} \cdot c_s^2 \cdot \mathcal{J} \pi_x = \frac{1}{4\pi \rho_0} \cdot \frac{k}{\omega} B_0 \mathcal{J} B_y \cdot \sin \gamma$

$\mathcal{J} \pi_{yz} = -\frac{1}{4\pi \rho_0} \frac{k}{\omega} B_0 \cdot \mathcal{J} B_{yz} \cdot \cos \gamma$

$\mathcal{J} p = \rho_0 \cdot \frac{k}{\omega} \cdot \mathcal{J} \pi_x$

HERE IS  
COUPLE  
VELOCITIES

VELOCITIES

$c_s^2 = \frac{\mathcal{J} p}{\mathcal{J} \rho}$

ADIABATIC SOUND SPEED

$c_k = \frac{\omega}{k}$

PHASE VELOCITY OF THE WAVES

$c_A^2 = \frac{B_0^2}{4\pi \rho_0}$

ALFVEN VELOCITY, FOR WAVES  
DRIVEN BY MAG. PRESSURE.

$$\begin{pmatrix} c_k^2 - c_s^2 - c_A^2 \sin^2 \gamma & c_A^2 \sin \gamma \cos \gamma & 0 \\ c_A^2 \sin \gamma \cos \gamma & c_k^2 - c_A^2 \cos^2 \gamma & 0 \\ 0 & 0 & c_k^2 - c_A^2 \cos^2 \gamma \end{pmatrix} \mathcal{J} \vec{\pi} = \vec{0}$$

$\det(\dots) = 0$  FOR NONTRIVIAL SOLUTIONS

THE POLYNOMIAL ESTABLISHES A DISPERSION RELATION BY SETTING  $c_k$  IN RELATION WITH  $c_s$  AND  $c_A$ . PROPAGATION IS DIRECTION DEPENDENT.

## DB PLASMA WAVES

$\det(\dots) = 0$  GIVES THE DISPERSION RELATION

$c_k^4 - c_k^2 \cdot (c_A^2 + c_s^2) + c_A^2 c_s^2 \cdot \cos^2 \gamma = 0$

$\Delta c_k^2 = \frac{1}{2} [(c_A^2 + c_s^2) \pm \sqrt{(c_A^2 + c_s^2)^2 - 4 \cdot c_A^2 c_s^2 \cdot \cos^2 \gamma}]$

EXTREME CASES

①  $\cos \varphi = 0$

$c_e^2 = c_A^2 + c_s^2$

②  $\cos \varphi = 1$

$c_e^2 = c_s^2$  OR  $c_A^2$ , DEPENDING ON ± SIGN

PROPAGATION IS ANISOTROPIC

D9 ELECTRODYNAMICS OF A WAVE IN PLASMA

→ PLASMA REACTS TO THE WAVE BY BEING POLARISED →

PLASMA REACTS ———→ MOTION OF THE CHARGE

CARRIERS; THIS LEADS TO AN ELECTRIC FIELD  $\vec{D} = \epsilon \cdot \vec{E}$

AND A MAGNETIC FIELD  $\vec{B} = \mu \vec{H}$

↑ DIELECTRIC CONST PERMITTIVITY

↑ PERMEABILITY

→ DIELECTRIC  $\epsilon$  AND PERMEABILITY CONSTANT ARE DETERMINED BY THE MEDIUM.

→ ANISOTROPY:  $\epsilon$  AND  $\mu$  ARE TENSORS, NOT ONLY SCALAR NUMBERS. (OF MEDIUM)

MAXWELL'S EQUATIONS IN MEDIUM

$\text{div } \vec{D} = 0$  (SET  $\rho = 0$ , OTHERWISE  $\text{div } \vec{D} = 4\pi \rho$ )

$\text{div } \vec{B} = 0$

$\partial_t \vec{D} = c \cdot \text{rot } \vec{H} + 4\pi \vec{j}$

$\partial_t \vec{B} = -c \cdot \text{rot } \vec{E}$

WITH  $\vec{D} = \epsilon \vec{E}$ ;  $\vec{j} = \sigma \cdot \vec{E}$  AND ASSUMING  $\mu = 1 \rightarrow \vec{B} = \vec{H}$

↑ CONDUCTIVITY OF MEDIUM

DIELECTRIC TENSOR:  $\vec{D} = \epsilon \cdot \vec{E}$ ;  $D_i = \epsilon_{ij} \cdot E_j$

- ANSATZ  $\epsilon_{ij} = A \delta_{ij} + B \cdot \frac{k_i k_j}{k^2}$  WITH 2 FUNCTIONS A AND B (FREQUENCY)

OR OTHER PROP. OF THE WAVE

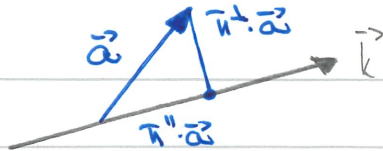
(VELOCITY TENSORS → HYDRODYN. ...)

PROJECTION OPERATORS (PROJECTION IS A LINEAR OPERATION)

$\pi_{ij}^{\parallel} = \frac{k_i k_j}{k^2}$ ;  $\pi_{ij}^{\perp} = \delta_{ij} - \frac{k_i k_j}{k^2} = \delta_{ij} - \pi_{ij}^{\parallel}$

WE HAVE A VECTOR  $\vec{E}$  AND WE CAN PROJECT OTHER VECTORS INTO THEIR COMPONENTS PARALLEL AND PERPENDICULAR.

$n^{\perp} = \overline{n^{\perp}}$   
 $n^{\parallel} = \overline{n^{\parallel}}$  (SAME PT)  
 $n = \overline{n}$  (SAME)



VECTOR  $\vec{k}$  AND ANOTHER VECTOR  $\vec{a}$

WITH THE PROPERTIES  $\vec{n}^{\parallel} \cdot \vec{n}^{\parallel} = \vec{n}^{\parallel}$ ;  $\vec{n}^{\perp} \cdot \vec{n}^{\perp} = \vec{n}^{\perp}$  [IDENTITY]  
 $\vec{n}^{\parallel} \cdot \vec{n}^{\perp} = \vec{n}^{\perp} \cdot \vec{n}^{\parallel} = 0$  [PERPENDICULARITY]

BASIS OF  $\epsilon$ -TENSOR  $\epsilon_{ij} = \epsilon_{\perp} \cdot \vec{n}^{\perp}_i \vec{n}^{\perp}_j + \epsilon_{\parallel} \cdot \vec{n}^{\parallel}_i \vec{n}^{\parallel}_j$  — LET'S SEPARATE THE AMPLITUDES

$\text{tr}(\epsilon \cdot \vec{n}^{\perp}) = \text{tr}(\epsilon_{\perp} \cdot \vec{n}^{\perp} \cdot \vec{n}^{\perp} + \epsilon_{\parallel} \cdot \vec{n}^{\parallel} \cdot \vec{n}^{\perp}) = \text{tr}(\epsilon_{\perp} \vec{n}^{\perp}) = \epsilon_{\perp} \text{tr}(\vec{n}^{\perp}) = \epsilon_{\perp}$   
 $\text{tr}(\epsilon \cdot \vec{n}^{\parallel}) = \text{tr}(\epsilon_{\perp} \cdot \vec{n}^{\perp} \cdot \vec{n}^{\parallel} + \epsilon_{\parallel} \cdot \vec{n}^{\parallel} \cdot \vec{n}^{\parallel}) = \text{tr}(\epsilon_{\parallel} \vec{n}^{\parallel}) = \epsilon_{\parallel} \text{tr}(\vec{n}^{\parallel}) = \epsilon_{\parallel}$

TRY A WAVE  $\sim \exp(i \cdot (\vec{k}\vec{x} - \omega t))$

$\left\{ \begin{array}{l} \frac{\omega}{c} \vec{B} = \vec{k} \times \vec{E} \quad ; \quad \vec{k} \cdot \vec{B} = 0 \\ \frac{\omega}{c} \vec{D} = -\vec{k} \times \vec{E} \quad ; \quad \vec{k} \cdot \vec{D} = 0 \end{array} \right.$  IN VACUUM  
 $\rightarrow \vec{k} \times (\vec{k} \times \vec{E}) = \frac{\omega}{c} \vec{k} \times \vec{B} = -\frac{\omega^2}{c^2} \vec{D}$   
 $\rightarrow \vec{k} \cdot (\vec{k} \cdot \vec{E}) - k^2 \cdot \vec{E} = -\frac{\omega^2}{c^2} \vec{D}$

WITH DIELECTRIC TENSOR  $D_i = \epsilon_{ij} E_j$   
 $\rightarrow \frac{\omega^2}{c^2} \cdot \epsilon_{ij} E_j = k^i \cdot k_j E_i - k_i k^j E_j = k^2 \left( \delta_{ij} - \frac{k_i k_j}{k^2} \right) E_j$

$\rightarrow \left( \delta_{ij} - \frac{k_i k_j}{k^2} - \left( \frac{\omega}{ck} \right)^2 \cdot \epsilon_{ij} \right) \cdot E_j = 0$   
 $= \vec{n}^{\perp}_{ij}$

INSERT DIELECTRIC TENSOR WITH THE COMPONENTS  $\epsilon_{\parallel}$  AND  $\epsilon_{\perp}$   
 + CONSTRUCT DISPERSION RELATION FROM THE DETERMINANT.

$\det \left[ \left( 1 - \left( \frac{\omega}{ck} \right)^2 \cdot \epsilon_{\perp} \right) \vec{n}^{\perp} - \left( \frac{\omega}{ck} \right)^2 \cdot \epsilon_{\parallel} \vec{n}^{\parallel} \right] = 0$

NOW TRANSVERSE AND LONGITUDINAL WAVES.

• TRANSVERSE WAVES:  $\vec{n}^{\parallel} \cdot \vec{E} = 0 \rightarrow \det \left[ \left( 1 - \left( \frac{\omega}{ck} \right)^2 \epsilon_{\perp} \right) \vec{n}^{\perp} \right] = 0$   
 $\rightarrow \omega = \frac{ck}{\sqrt{\epsilon_{\perp}}}$ , BECAUSE  $\det(\vec{n}) = 1$

LONGITUDINAL WAVES:  $\vec{n} \perp \vec{E} = 0 \rightarrow \text{div} \left[ \left( \frac{c}{\omega} \right)^2 \cdot \epsilon_{ij} \vec{n} \right] = 0$

ABE TO B410 KULA  $\xrightarrow{\text{NUSI}} \epsilon_{ij} = 0$  NO LONGITUDINAL WAVE PROPAGATION

### DIO SHADES OF SPECTRA LINES

- WE CAN IDENTIFY CERTAIN ELEMENTS FROM THEIR SPECTRAL SIGNATURE, AND NOW WE CAN EVEN SAY SOMETHING ABOUT STATE OF MOTION OF THE SYSTEM, GRAVITATIONAL FIELDS ETC.

IDEA: RADIATION IS PRODUCED BECAUSE ATOMS ARE COUPLED TO THE EM-FIELD (RADIATION FIELD)

TRANSFER BETWEEN ATOMS/PLASMA AND THE RADIATION FIELD  $\xrightarrow{\text{OR VICE VERSA}}$  EMISSION OR ABSORPTION, DEPENDING ON THE DIRECTION OF THE ENERGY FLUX. EITHER  $\text{PHOTONS} \rightarrow \text{MATTER}$  OR  $\text{MATTER} \rightarrow \text{PHOTONS}$

● NATURAL LINE SHAPE - IDEALISED CASE OF AN ISOLATED ATOM  
 $\rightarrow$  TRANSITION BETWEEN TWO STATES OCCURS SPONTANEOUSLY.

$\Delta E \cdot \Delta t \sim \hbar$        $\Delta t = \Gamma^{-1}$  LIFE TIME OF AN EXCITED STATE  
 $\Delta E =$  ENERGY UNCERTAINTY  $\rightarrow$  TYPICAL WIDTH OF A SPECTRAL LINE

$$E = \frac{p^2}{2m} ; \Delta E = \frac{p}{m} \Delta p$$

$$p = m \cdot \frac{\Delta x}{\Delta t}$$
$$\Delta t = \frac{m \cdot \Delta x}{p}$$

$$\Delta E \Delta t = \frac{p}{m} \Delta p \cdot \frac{m}{p} \Delta x = \Delta p \Delta x \sim \hbar$$

RADIATIVE TRANSITION (OSCILLATION) BETWEEN TWO STATES  $E_1 \rightleftharpoons E_2$

$$i \hbar \partial_t \rho_{nm} = \sum_{m'} \langle m | H | m' \rangle \cdot \rho_{m'n} \cdot \exp\left(\frac{i}{\hbar} \cdot (E_m - E_{m'}) \cdot t\right)$$

## SCHROD. EQUATION

FOR THE AMPLITUDE  $a_n$  OF THE STATE  $n$ .  $\rightarrow$

$$i\hbar \partial_t a_1(t) = \langle 1 | H | 2 \rangle \cdot a_2(t) \cdot \exp\left(\frac{i}{\hbar} (E_2 - E_1) \cdot t\right)$$

WITH THE ANSATZ:  $a_2(t) = \exp\left(-\frac{\Gamma \cdot t}{2} \cdot \exp(+i\omega t)\right)$

*LOW PRESSURE TO LOW COLLISIONS*

$$\rightarrow i\hbar \partial_t a_1 = \dots = \langle 1 | H | 2 \rangle \cdot \exp\left[i(\omega - \omega_{12})t - \frac{\Gamma}{2}t\right]$$

$\hookrightarrow \frac{E_1 - E_2}{\hbar}$

$\rightarrow$  INTEGRATION FROM 0 TO  $\infty$ :  $a_1(t=0) = 0$

$$a_1(t) = \frac{i}{\hbar} \langle 1 | H | 2 \rangle \cdot \frac{1 - \exp\left[i(\omega - \omega_{12}) \cdot t - \frac{\Gamma}{2}t\right]}{\omega - \omega_{12} + i \cdot \frac{\Gamma}{2}}$$

AT LATE TIMES  $\Gamma \cdot t \gg 1 \cdot \exp\left(-\frac{\Gamma}{2}t\right) \rightarrow 1$

$$\rightarrow |a_1(t)|^2 = \left| \frac{\langle 1 | H | 2 \rangle}{\hbar} \right|^2 \cdot \frac{1}{|\omega - \omega_{12} + i \frac{\Gamma}{2}|^2} = \underbrace{\left| \frac{\langle 1 | H | 2 \rangle}{\hbar} \right|^2 \cdot \frac{1}{(\omega - \omega_{12})^2 + \frac{\Gamma^2}{4}}}_{\text{LORENTZ LINE PROFILE}}$$

PROFILE FUNCTION  $\phi(\omega) = \frac{\Gamma}{(\omega - \omega_{12})^2 + \frac{\Gamma^2}{4}}$  ... WITH  $\int_{-\infty}^{\infty} d\omega \phi(\omega) = 1$

## COLLISIONAL BROADENING

NATURAL LINE WIDTH  $\sim$  DETERMINED BY THE LIFE TIME

OF STATES, COLLISIONS CAN REDUCE

THE LIFE TIME  $\rightarrow$  LINE GETS BROADER

LOW PRESSURE  $\rightarrow$  LOW COLLISION  $\rightarrow$  LESS BROADENING

COLLISIONS CHANGE THE SCHRODINGER EQUATION BY A RANDOM

PHASE:

$$i\hbar \cdot \partial_t a_1(t) = \langle 1 | H | 2 \rangle \cdot \exp\left[i(\omega - \omega_{12})t - \frac{\Gamma}{2}t\right] \cdot \exp(i\phi(t))$$

$$\langle \exp(i\phi(t)) \rangle = \begin{cases} 1 & \text{IF THERE WAS NO COLLISION UNTIL } t \\ 0 & \text{IF THERE WAS A COLLISION} \end{cases}$$

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INTRODUCE MANY PARTICLES: EXPECTED VALUE SHOULD FOLLOW  
a POISSON DISTRIBUTION

$$\langle \exp(i\phi(t)) \rangle \rightarrow \exp(-\frac{\Gamma_c}{2} t) \sim \text{CLASSICAL PROBABILITY}$$

$$\Delta a_1 \sim \int dt \cdot \exp[i(\omega - \omega_{12})t - \frac{\Gamma}{2} t + i\phi], \text{ AND AFTER AVERAGING}$$

$$\langle \Delta a_1 \rangle \sim \int dt \cdot \exp[i(\omega - \omega_{12})t - \frac{\Gamma}{2} t - \frac{\Gamma_c}{2} t] \text{ WITH } \exp(\frac{\Gamma_c}{2} t)$$

$\Gamma_c$  (INCREASING THIS WITH  
WITH COLLIS.)

$$\text{EFFECTIVELY } \frac{1}{\Gamma} \rightarrow \frac{1}{\Gamma} + \frac{1}{\Gamma_c} = \frac{1}{\Gamma} \cdot \frac{1}{1 + \frac{\Gamma_c}{\Gamma}}$$

$\frac{\Gamma_c}{\Gamma} > 0$

~ LINES GET BROADENED  
BY COLLISIONS

### • VELOCITY BROADENING

ATOMS IN MOTION  $\rightarrow$  OBSERVE A DOPPLER EFFECT

$$\frac{\nu}{\nu_0} \sim 1 + \beta \quad \text{WITH } \beta = \frac{v}{c}$$

IF THE ATOMS HAVE VELOCITIES FOLLOWING A MAXWELL-DISTRIBUTION  
 $\rightarrow$  CONVOLUTION BETWEEN THE PROFILE SLOPE AND THE DISTRIBUTION

### • VOIGT - PROFILE

NATURAL LINE PROFILE + COLLISIONAL BROADENING  $\rightarrow$  LORENTZ

VELOCITY BROADENING  $\rightarrow$  CONVOLUTION BETWEEN LORENTZ + MAXWELL

$$\phi(\omega) = \int dv \frac{\Gamma}{(\omega - \omega_{12})^2 + \frac{\Gamma^2}{4}} \cdot \frac{1}{\sqrt{2\pi}v} \cdot \exp(-\frac{v^2}{2v_0^2})$$

$$\text{BUT } \omega_{12} \rightarrow \omega_{12} \cdot (1 + \beta) \text{ WITH } \beta = \frac{v}{c}$$

$$\rightarrow \phi(\omega) = \int dV \frac{\Gamma}{(\omega - \omega_{12} - \omega_{12} \cdot \frac{v}{c})^2 + \frac{\Gamma^2}{4}} \cdot \frac{1}{\sqrt{2\pi} \sqrt{v}} \cdot \exp\left(-\frac{v^2}{2\Gamma^2}\right)$$

$$\phi(\omega) = \sqrt{\frac{2}{\pi}} \cdot \frac{ac}{\sqrt{v} \cdot \omega_{12}} \cdot \int_{-\infty}^{+\infty} dq \frac{\exp(-q^2)}{(\omega - q)^2 + a^2}$$

$$\text{WITH } a = \frac{\Gamma}{\omega_{12}} \cdot \frac{c}{2\sqrt{2}v} \quad ; \quad q = \frac{v}{\sqrt{2} \sqrt{v}} \quad ; \quad \mu = \frac{\omega - \omega_{12}}{\omega_{12}} \cdot \frac{c}{\sqrt{2} \sqrt{v}}$$

$$\text{SMALL } \Delta\omega: \frac{1}{(\omega - q)^2 + a^2} \sim 1 \rightarrow \text{GAUSSIAN SLOPE}$$

$$\text{LARGE } \Delta\omega: \exp(-q^2) \sim 1 \rightarrow \text{LORENTZIAN SLOPE}$$