

PMAS-PPAK integral-field spectroscopy of nearby Seyfert and normal spiral galaxies

Comparative study of nearby Seyfert and normal galaxies

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Almeria, 7th June 2010

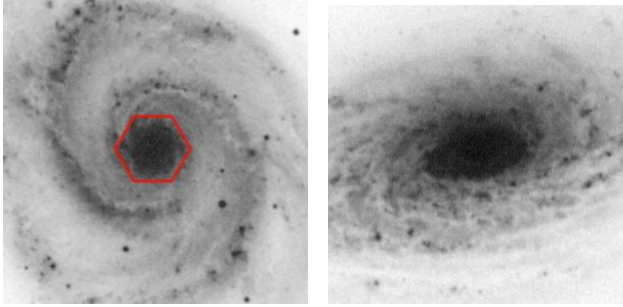
our project vs CALIFA

- PMAS-PPAK IFU
 - $0.001 < z < 0.005$
 - 8 galaxies, Seyfert vs normal spirals
 - $3600\text{\AA} - 7000\text{\AA}$, $R \sim 500$
 - $6200\text{\AA} - 6900\text{\AA}$, $R \sim 2360$
 - data obtained in 2005
- PMAS-PPAK IFU
 - $0.005 < z < 0.03$
 - 600 galaxies
 - $4300\text{\AA} - 7000\text{\AA}$, $R \sim 850$
 - $3700\text{\AA} - 5000\text{\AA}$, $R \sim 1650$
 - data being obtained (2010 - 2012)

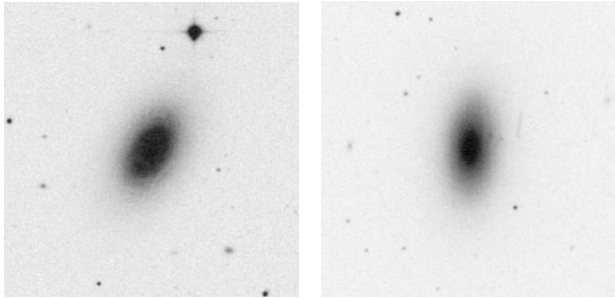
Our project is aimed to study differences in properties of ionized gas, gas & stellar kinematics and stellar population in central regions of nearby Seyfert and normal spiral galaxies.

The sample of 8 galaxies - meant as a pilot study.

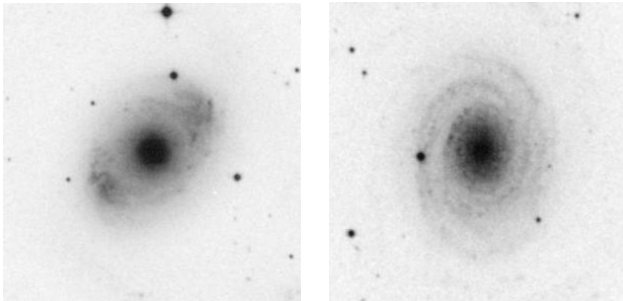
Seyfert & normal



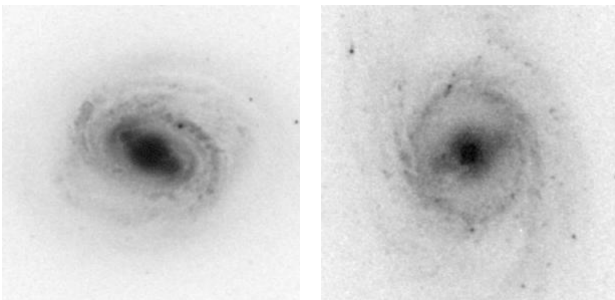
NGC 5194 (M51) & NGC 5055



NGC 4138 & NGC 3245



NGC 4151 & NGC 2985



NGC 4579 & NGC 3351

our project

← DSS images (**5' x 5'**) of 4 pairs of Seyfert & normal galaxies matched in Hubble type and distance

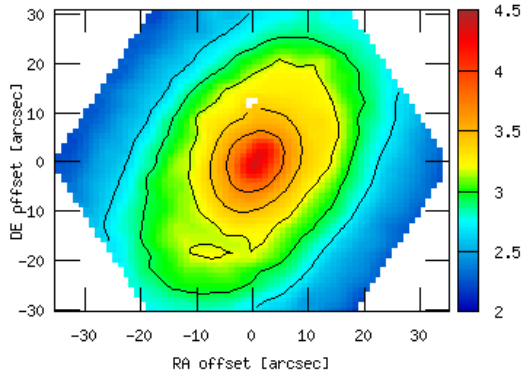
- data reduction processed with R3D
- analysis processing with FIT3D

stellar population modelling using the synthetic library of Bruzual & Charlot (2003)

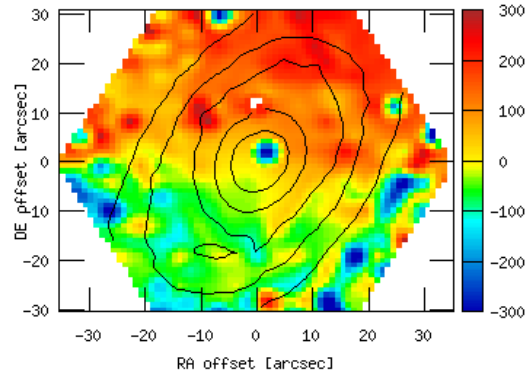
emission lines fitted by simple Gaussian functions

analysis -- NGC 4138

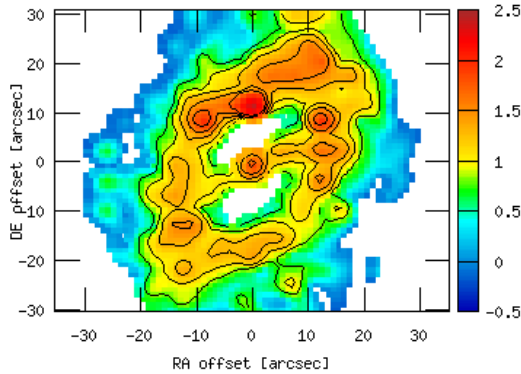
Intensity in stellar continuum (log scale) [$e^{-16} \text{erg/s/cm}^2/\text{arcsec}^2$]



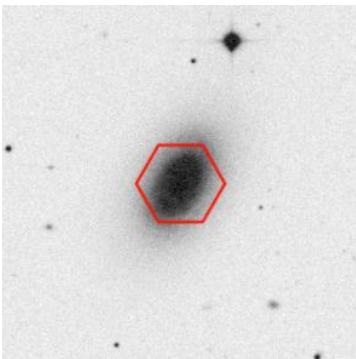
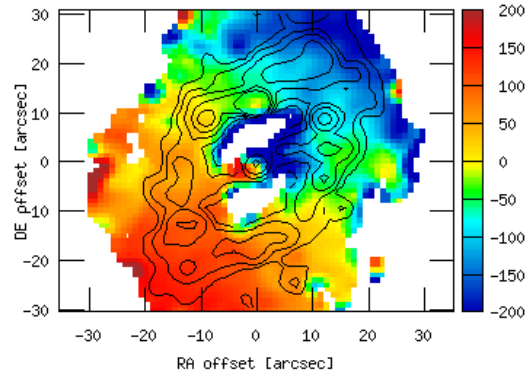
Stellar velocities $|-300/300|$ ($v_{\text{sys}}=888 \text{km/s}$) [km/s]



Intensity in H α (log scale) [$e^{-16} \text{erg/s/cm}^2/\text{arcsec}^2$]

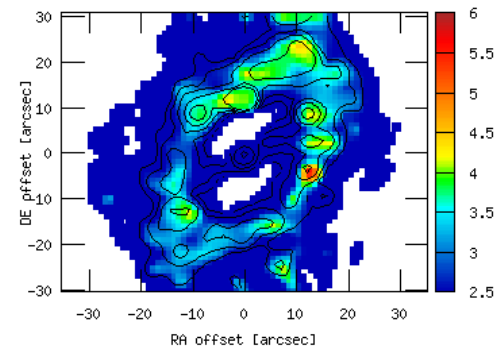


Velocity in H α $|-200/200|$ ($v_{\text{sys}}=888 \text{km/s}$) [km/s]



H α / H

Ratio H α /H β [2.5/*]



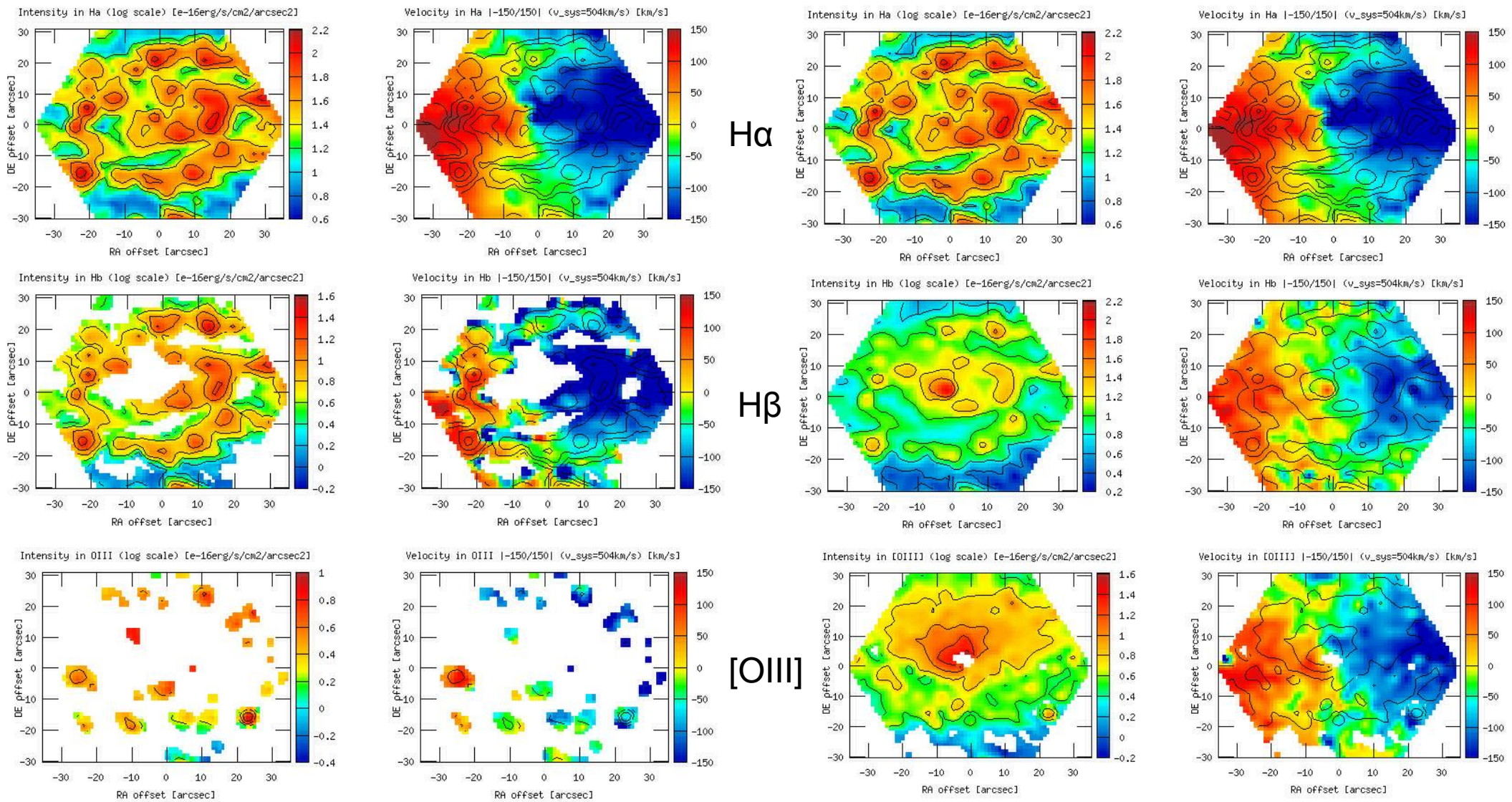
- Seyfert 1.9, SA(r)0+
- 1" ~ 80 pc
- two counterrotating stellar disks and gaseous disk
- (Jore et al., 1996)
- H α ring, no bar
- (Pogge & Eskridge, 1987)
- chemically distinct core
- (Afanasiev & Silchenko, 2002)
- merger, destroyed bar, both?

analysis -- stellar populations influence

without stellar population subtraction

with stellar population subtraction

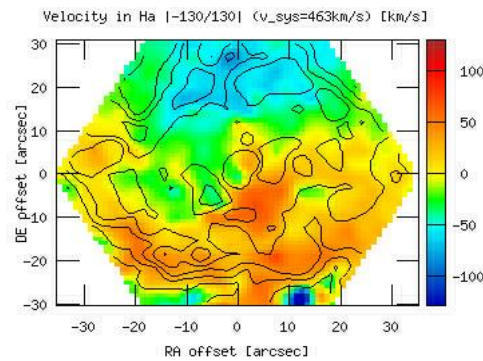
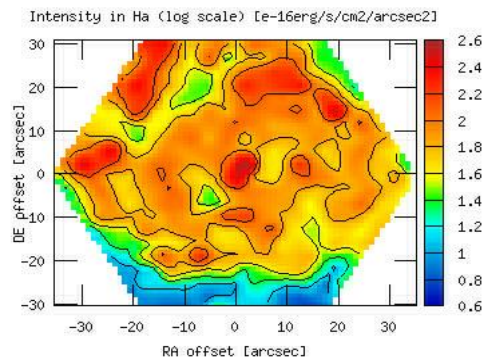
NGC 5055



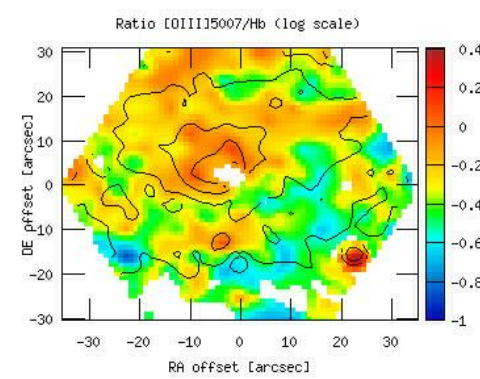
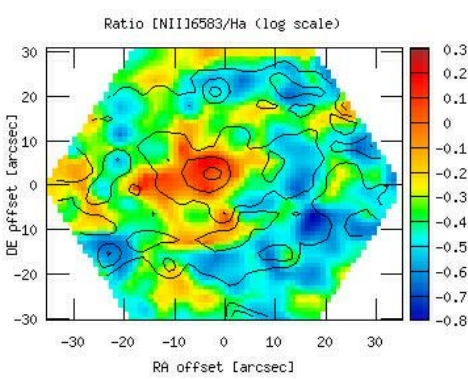
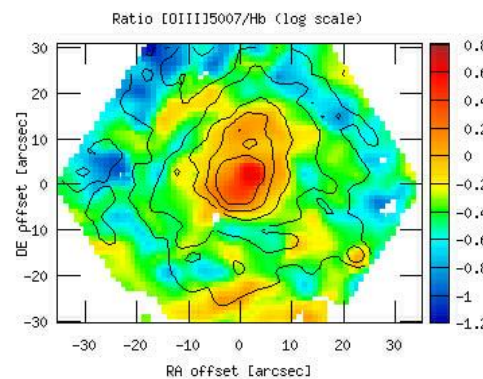
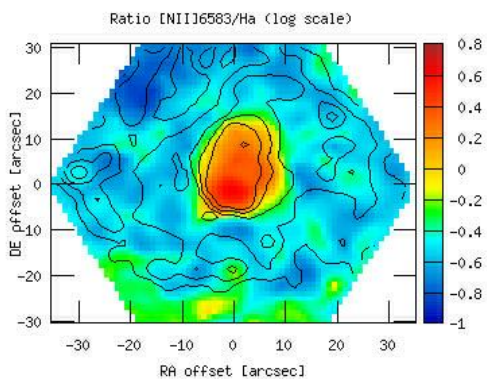
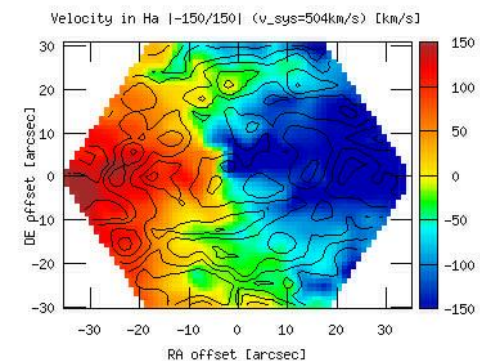
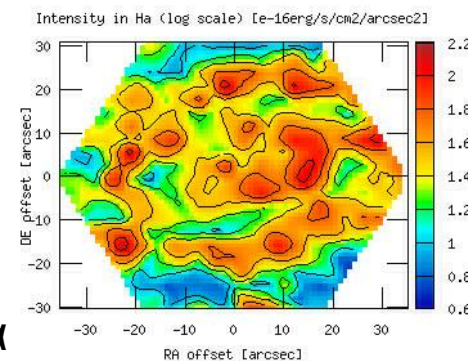
analysis -- active vs normal galaxies

NGC 5194 (M51) - Seyfert 2

NGC 5055 - normal galaxy



H α



[NII] / H α

[OIII] / H β

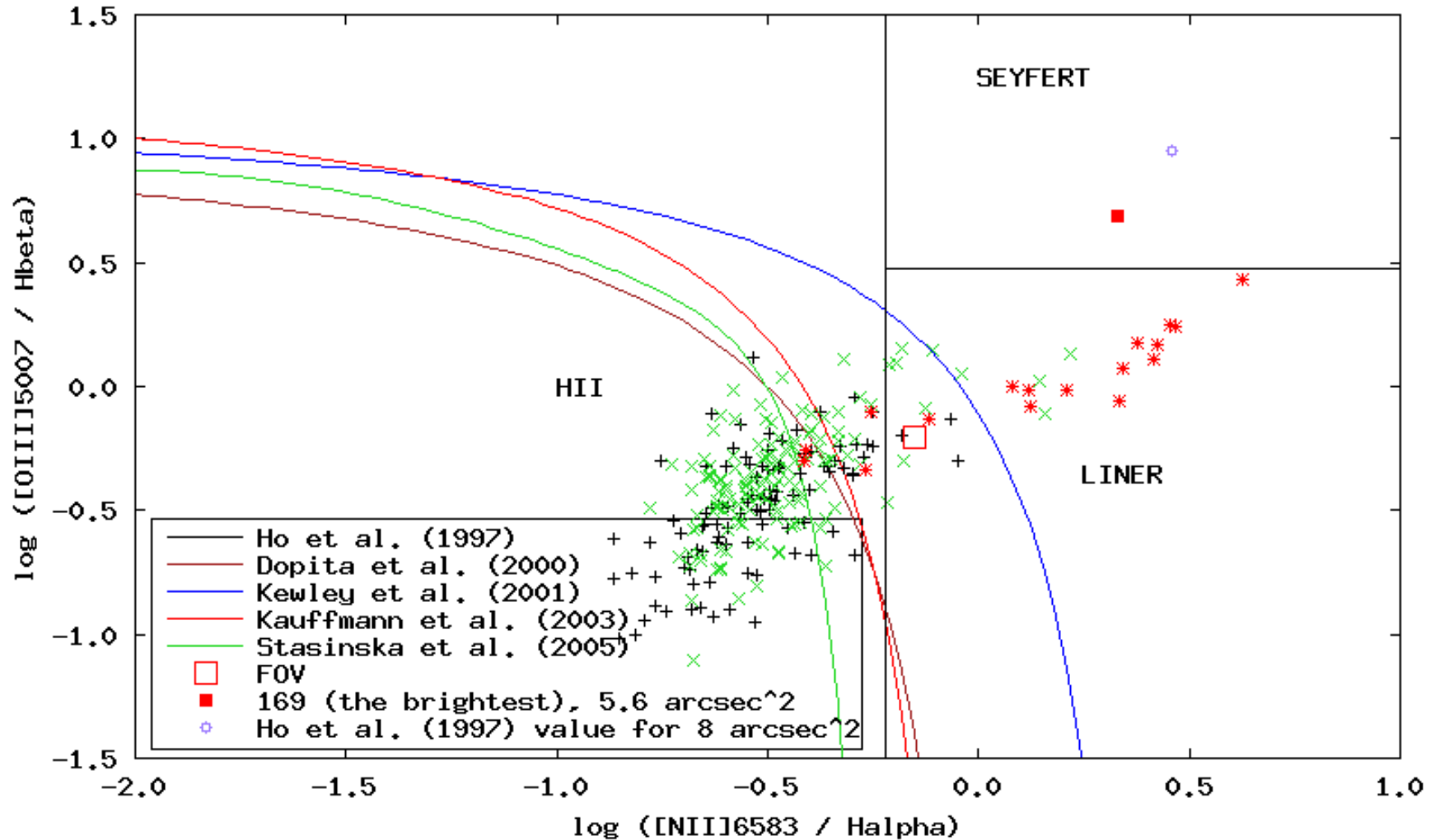
[NII] / H α

[OIII] / H β

spatially resolved BPT diagrams

NGC 5194 (M51) - Seyfert

NGC5194 Diagnostic diagram [OIII]5007 / Hbeta vs. [NII]6583 / Halpha, S/N > 3



red points: spaxels with $r \leq 10''$, **green points:** spaxels with $10'' < r \leq 25''$,
black points: spaxels with $r > 25''$

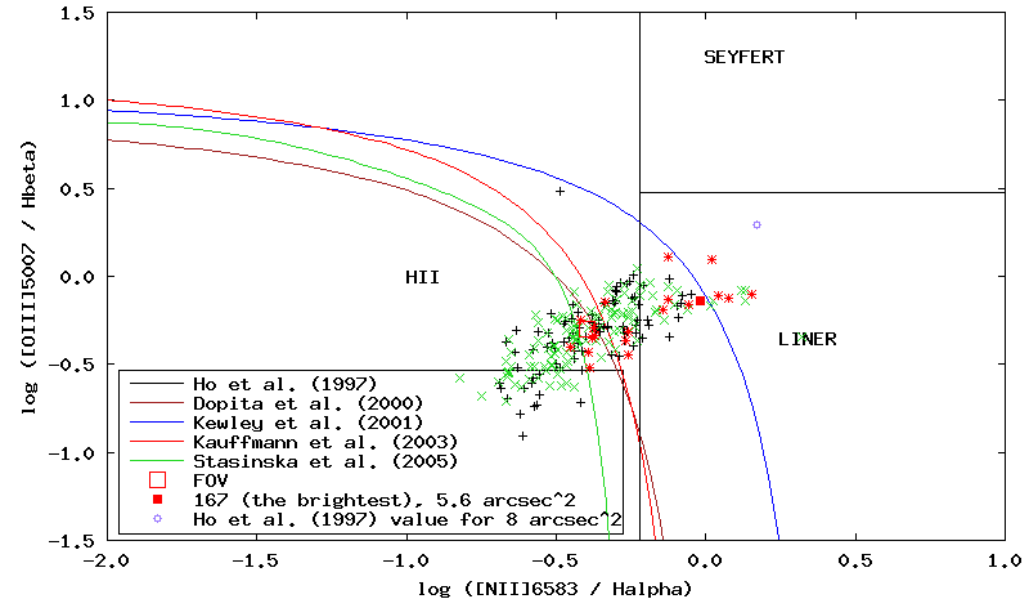
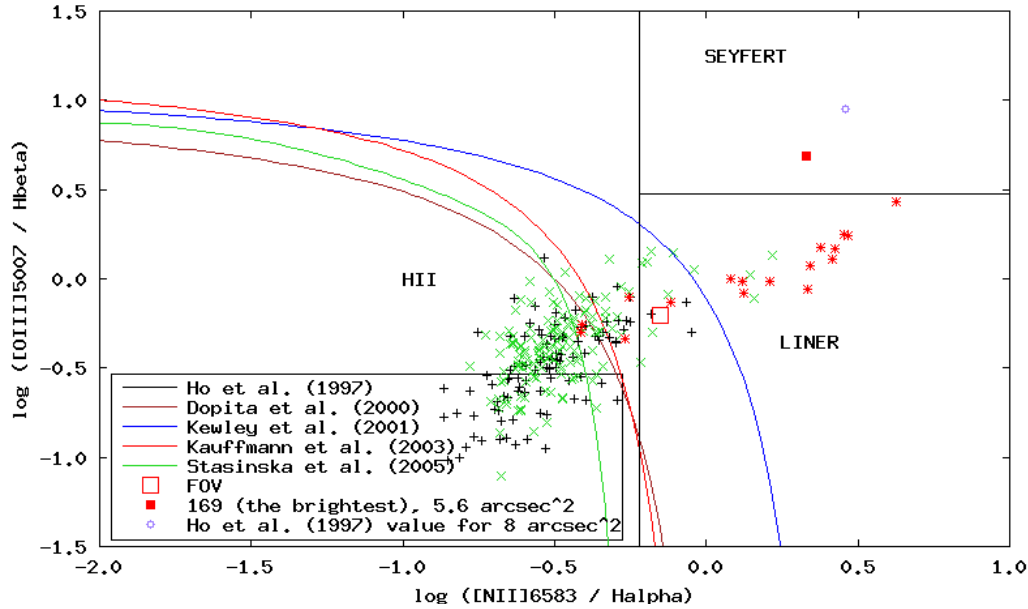
spatially resolved BPT diagrams

NGC 5194 (M51) - Seyfert

NGC 5055 - normal

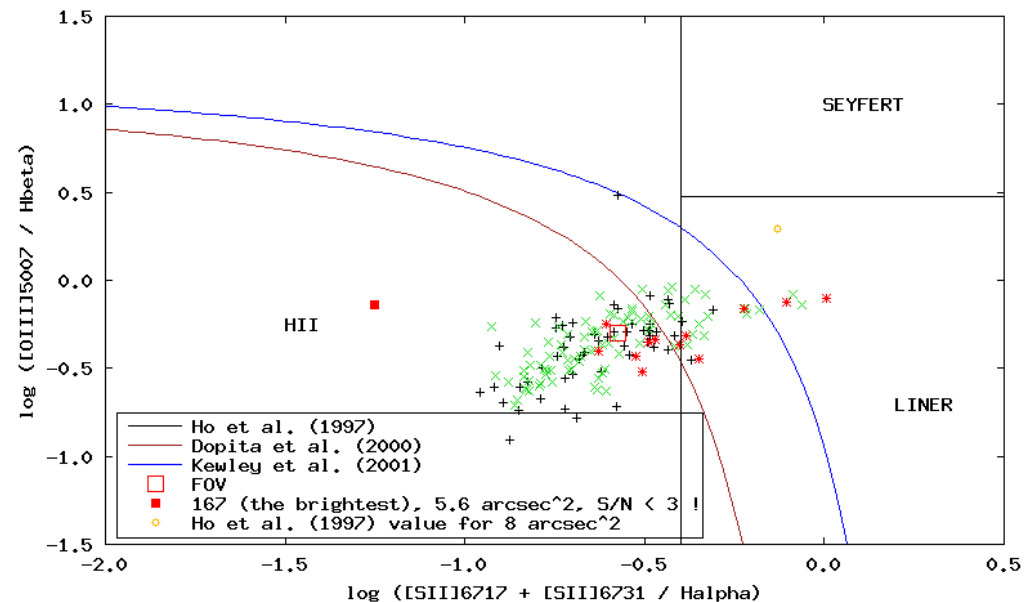
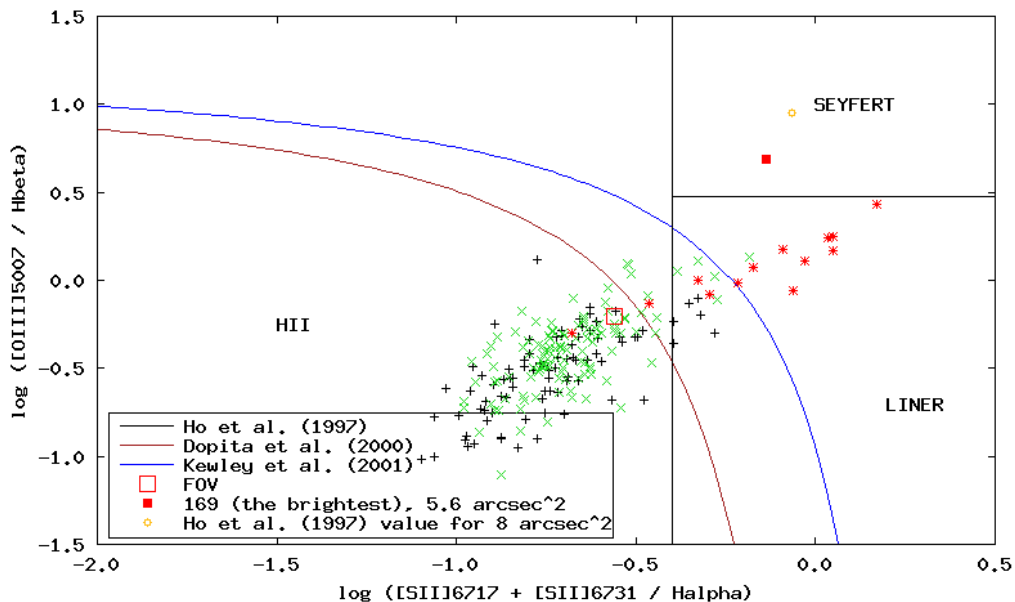
NGC5194 Diagnostic diagram [OIII]5007 / Hbeta vs. [NII]6583 / Halpha, S/N > 3

NGC5055 Diagnostic diagram [OIII]5007 / Hbeta vs. [NII]6583 / Halpha, S/N > 3



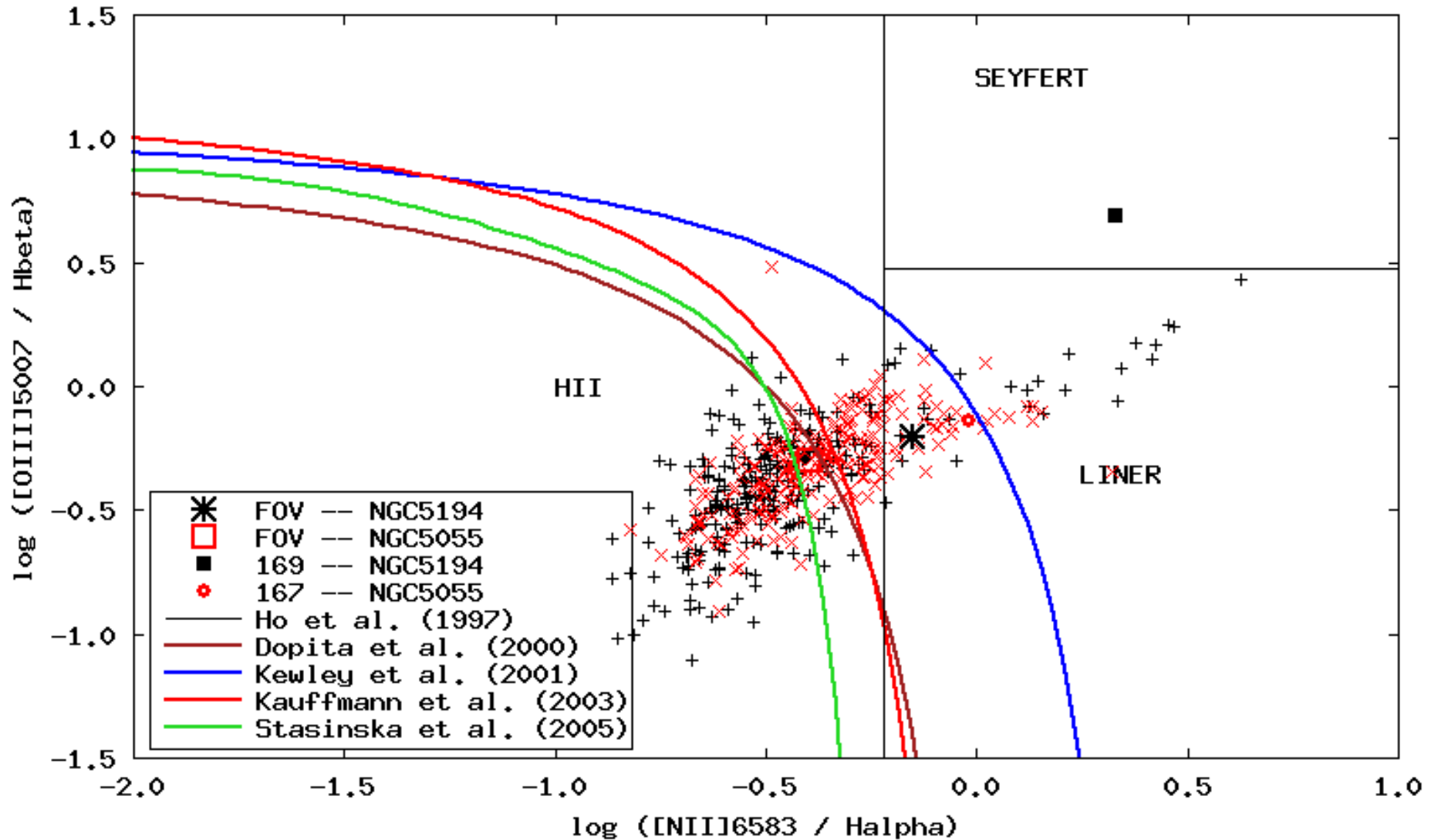
NGC5194 Diagnostic diagram [OIII]5007 / Hbeta vs. [SII]6731 / Halpha, S/N > 3

NGC5055 Diagnostic diagram [OIII]5007 / Hbeta vs. [SII]6731 / Halpha, S/N > 3



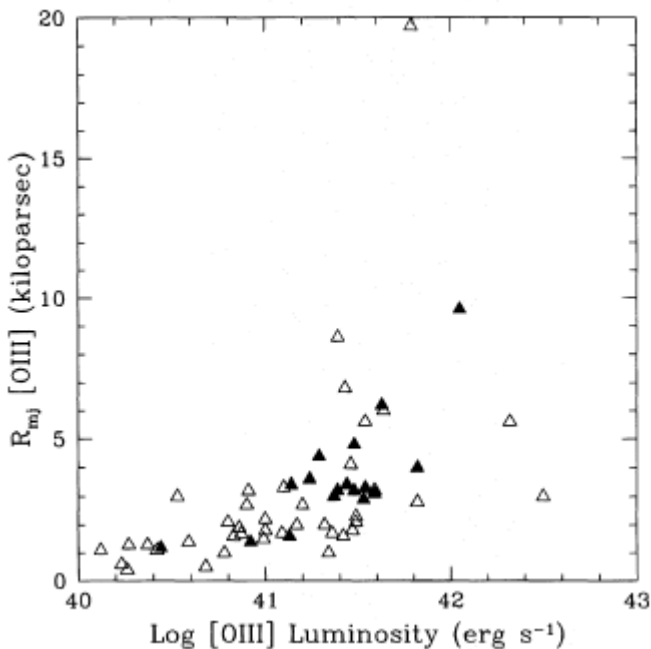
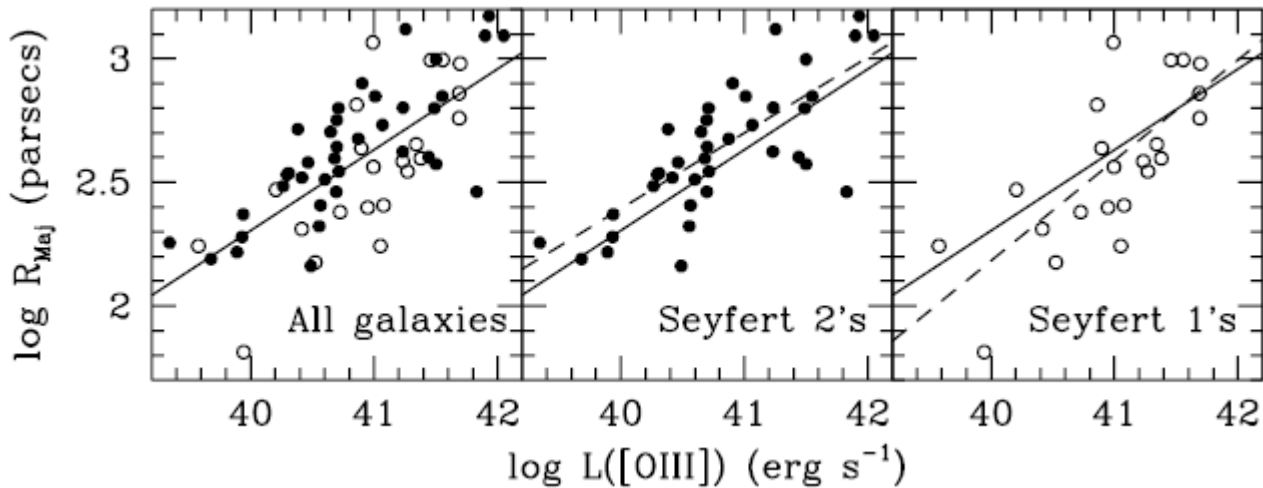
spatially resolved BPT diagrams

Comparison of NGC5194 & NGC5055 in diagnostic diagram [OIII]/H β vs. [NII]/H α , S/N>3



black points: NGC 5194 (M51) - Seyfert, **red points:** NGC 5055 - normal galaxy

size of NLR in CALIFA survey



- Schmitt et al. (2003)
- NLR size:
- 300 pc - 3.4 kpc
- for CALIFA with
- $0.005 < z < 0.03$:
- NLR $\sim 0.5''$ to $35''$ in diameter
- Mulchaey et al. (1996)
- NLR size:
- 300 pc - 14 kpc
- NLR $\sim 0.5''$ to $68''$ in diameter

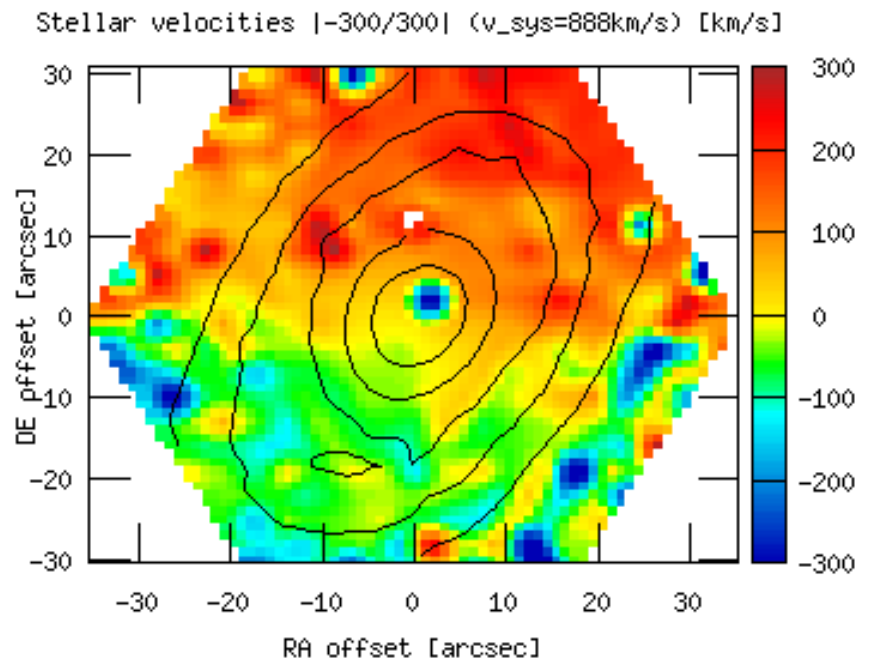
($H_0 = 71$ km/s/Mpc)

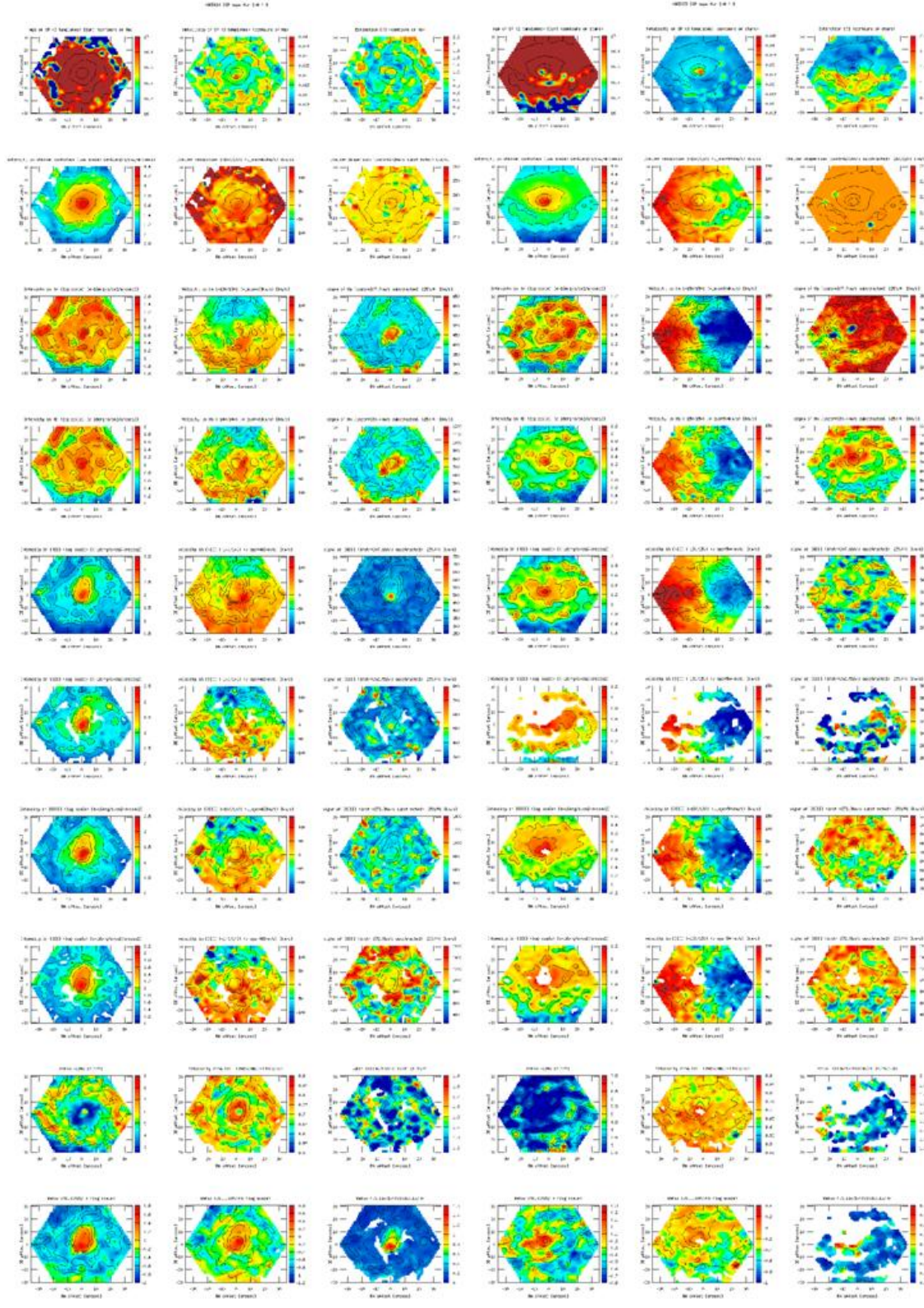
summary

- several aspects of PPAK data analysis were presented
 - study of complex structure of NGC 4138
 - role of stellar model subtraction
 - active vs non-active galaxy

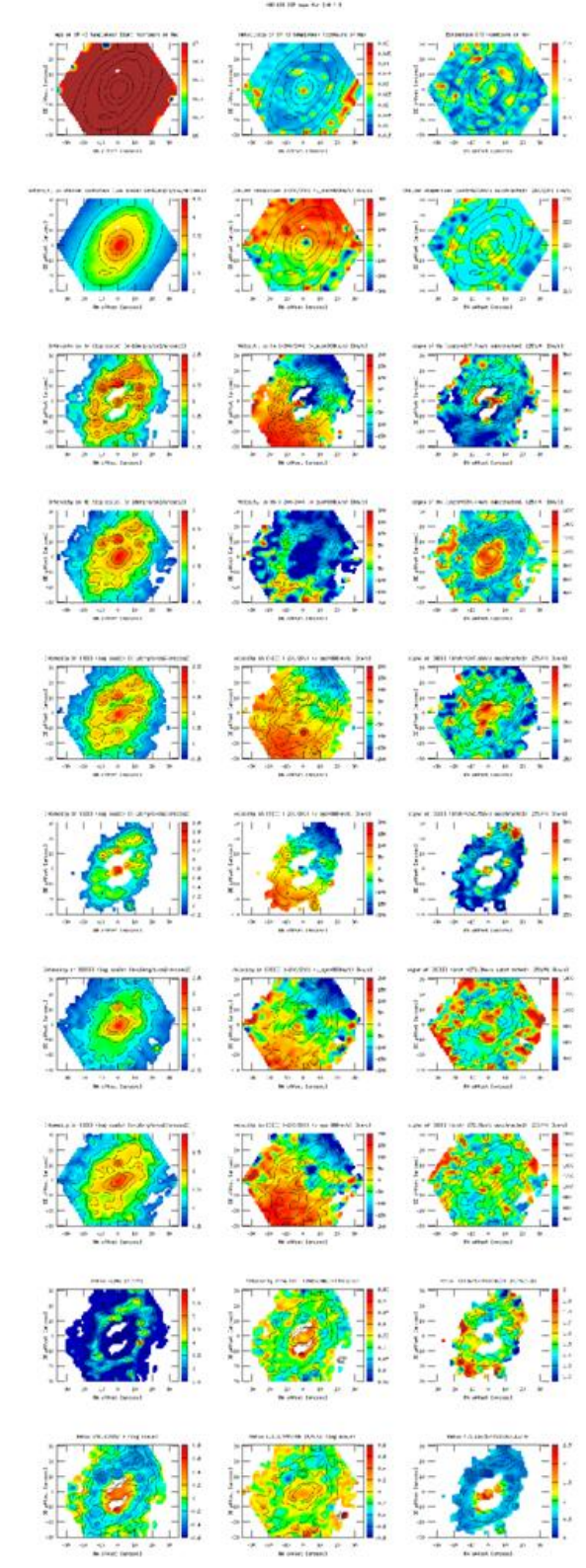
- multiple stellar kinematic components fitting is needed to disentangle complex dynamical structures

NGC 4138 - stellar mean LOS velocity





Thank
 You
 for
 Your
 Attention



references

- Afanasiev & Silchenko (2002), ApJ 124, 706
- Bennert et al. (2006), A&A 446, 919
- Bruzual & Charlot (2003), MNRAS 379, 1000
- Jore et al. (1996), AJ 112, 438
- Mulchaey et al. (1996), ApJ 467, 197
- Pogge & Eskridge (1987), AJ 92, 291
- Schmitt et al. (2003), ApJ 597, 768