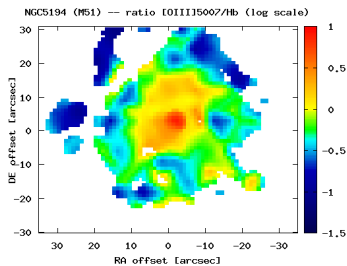
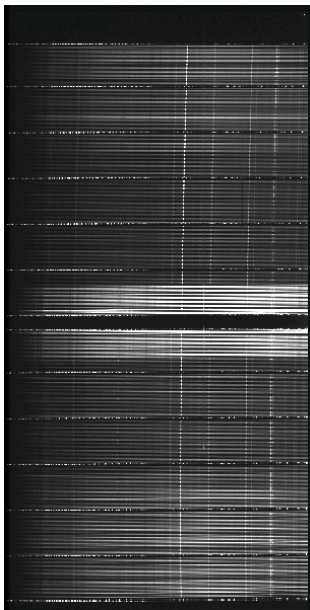


IFU spectroscopy of central parts of nearby Seyfert and normal galaxies

Tereza Bartáková, DTPA, Masaryk University, Brno
Supervisor: Bruno Jungwiert, ASL, Academy of Science, Prague

Ondřejov, 27th October 2009

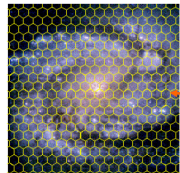
overview



- ▶ IFS – fundamentals, data cubes, motivation
- ▶ observed sample, aims
- ▶ data processing
- ▶ very first results – 2D maps
- ▶ status

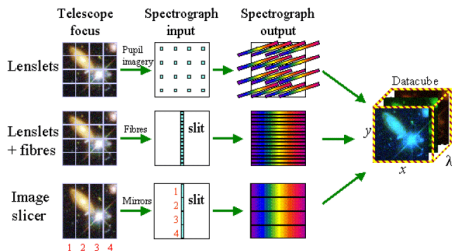
integral field spectroscopy (IFS)

- ▶ method to obtain many spectra from the target field simultaneously
- ▶ FOV divided into spatial elements (spaxels)
- ▶ today's integral field units (IFUs) have hundreds to a few thousand spaxels and FOV from a few arcseconds to an arcminute



when slit is not enough?

- ▶ to avoid light losses
- ▶ to avoid precise slit positioning
- ▶ to avoid slit effect
- ▶ asymmetric morphology and/or kinematics
- ▶ all spectra observed under the same conditions

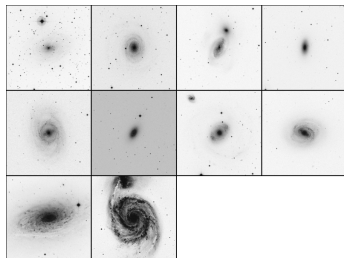


observed sample, aims

- ▶ PMAS-PPAK with FOV $\sim 1' \times 1'$ (currently the largest IFU)
- ▶ two setups: V300 (3650-6900 Å; FWHM ~ 500 km/s) and R1200 (6200-6850 Å; FWHM ~ 125 km/s)
- ▶ 5 pairs of AGN (Seyferts) and normal galaxies with the same Hubble type, similar inclination, distance and luminosity

to obtain

- ▶ 2D kinematics of stars and ionized gas
- ▶ stellar populations, reddening, electron density, temperature, ionization sources of the gas
- ▶ properties of the central kpc of active and non-active galaxies

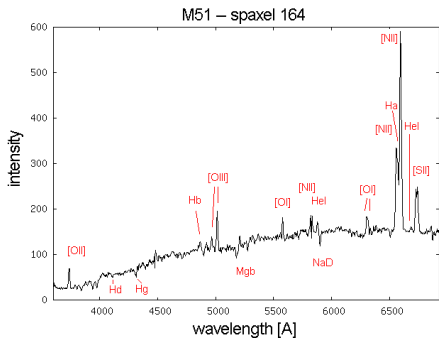
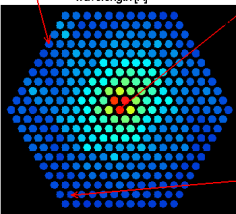
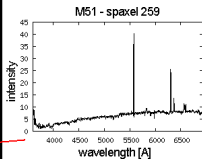
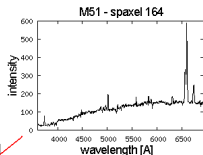
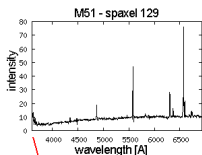


ic356, ngc2985, ngc3227, ngc3245, ngc3351, ngc4138, ngc4151, ngc 4579, ngc5055, ngc5194 (DSS images)

data – M51

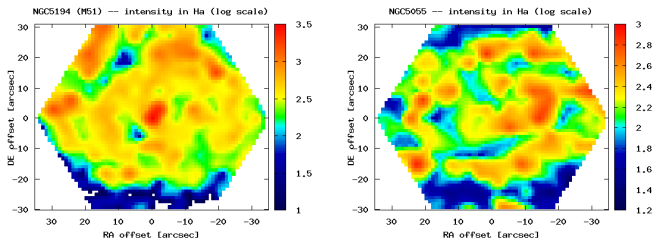


- ▶ 331 spectra of each galaxy
- ▶ emission line fluxes, velocities, FWHMs, ratios: $[NII]6583/H\alpha$, $[OIII]5007/H\beta$, $[OI]6300/H\alpha$, $H\alpha/H\beta$, $[SII]6731/[SII]6717$, ...
- ▶ fitting stellar population templates

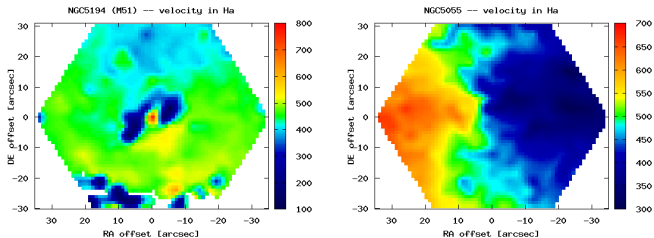


2D maps: M51 (AGN) vs. NGC5055 (normal g.)

flux in $H\alpha$

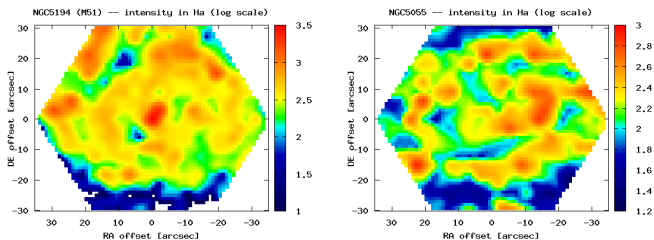


velocity in $H\alpha$

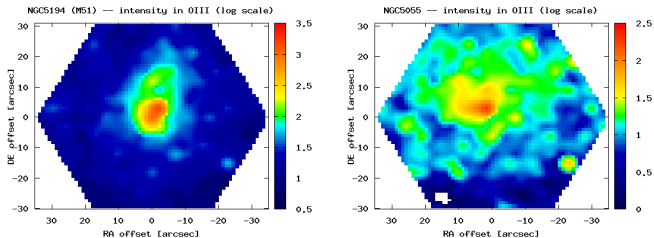


2D maps: M51 (AGN) vs. NGC5055 (normal g.)

flux in $H\alpha$



flux in [OIII]



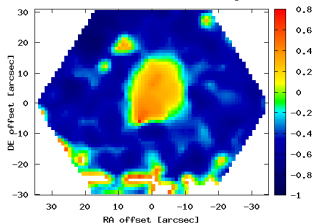
status

- ▶ data reduction – done
- ▶ emission line fitting – in progress
- ▶ 2D maps of fluxes, mean line-of-sight velocities, FWHMs, extinctions and finding ionization sources – in progress
- ▶ diagnostic diagrams – in progress
- ▶ stellar templates fitting to get age, metallicity, kinematics – in progress
- ▶ modelling kinematics of ionized gas – planned

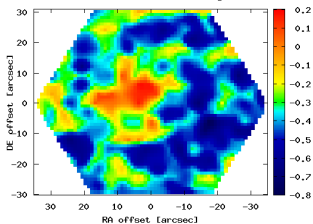
2D maps: M51 (AGN) vs. NGC5055 (normal g.)

ratio $[\text{NII}]\lambda 6583 / \text{H}\alpha$

NGC5194 (M51) -- ratio $[\text{NII}]\lambda 6583 / \text{H}\alpha$ (log scale)

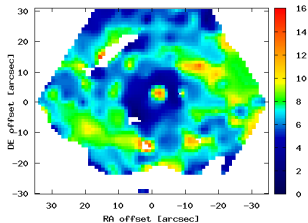


NGC5055 -- ratio $[\text{NII}]\lambda 6583 / \text{H}\alpha$ (log scale)

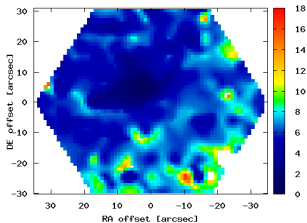


ratio $\text{H}\alpha / \text{H}\beta$

NGC5194 (M51) -- ratio $\text{H}\alpha / \text{H}\beta$

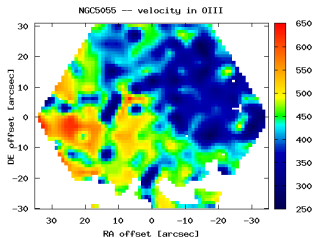
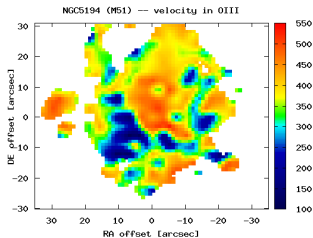


NGC5055 -- ratio $\text{H}\alpha / \text{H}\beta$



2D maps: M51 (AGN) vs. NGC5055 (normal g.)

velocity in [OIII]



ratio [OIII]5007 / H β

